

# The SENSEI<sup>†</sup> project

a zero noise detector for DM searches

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Fermi National Laboratory

February 18, 2018

<sup>†</sup> **Sub-Electron-Noise SkipperCCD Experimental Instrument**

## Origin of the Collaboration



**SnowPAC 2013**

**SnowDARK 2013**

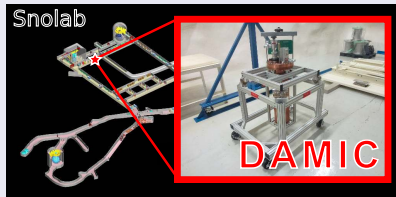
**SnowCluster 2013**

**March 22 - 25, 2013**

- Scientific CCDs as low-energy-threshold/low-noise particle detectors
- SENSEI project: status and prospects
- Application to light DM searches

# Current experiments using Scientific CCDs

## DAMIC



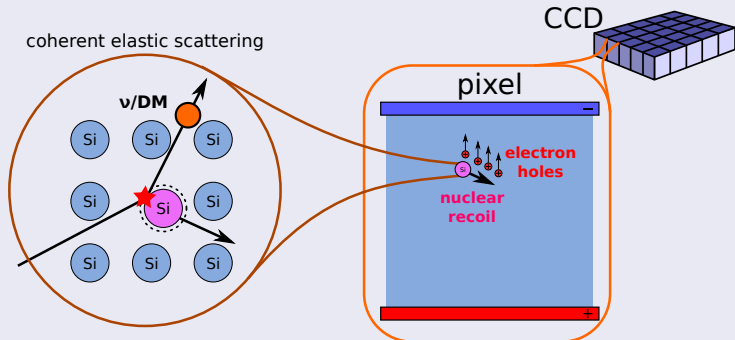
- Low mass Dark Matter search (WIMP/NR-optimized)
- Installed at Snolab on Dec-2012
- Currently taking data

## CONNIE



- Coherent  $\nu$ -nucleous interaction
- Installed next to Angra nuclear power plant on Dec-2014
- technique could be used for  $SB\nu$ -Ex
- Currently taking data

# DAMIC & CONNIE use CCDs as targets to detect coherent DM/ $\nu$ -nucleus interactions by measuring the ionization produced by the nuclear recoils

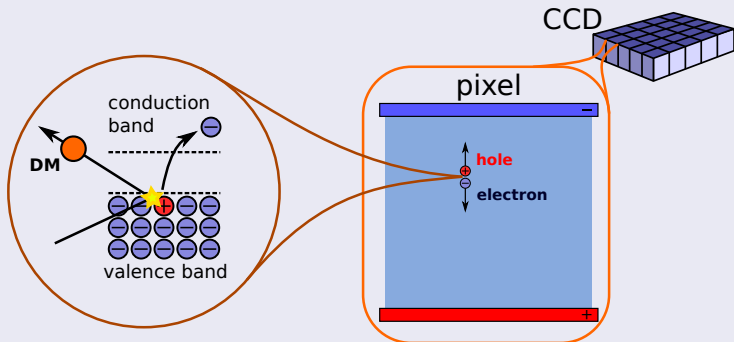


Sensitivity is limited by readout noise and NR-ionization yield

**SENSEI: lower the energy threshold to look for light DM candidates**

**Detect DM-e interactions by measuring the ionization produced by the electron recoils. See arXiv:1509.01598**

**Idea: use electrons in the CCDs as target**



This requires very low noise!

## SENSEI LDRD Collaboration (2015)

Develop a CCD-based detector with an energy threshold close to the silicon band gap (1.1 eV) using SkipperCCDs produced at LBL MSL

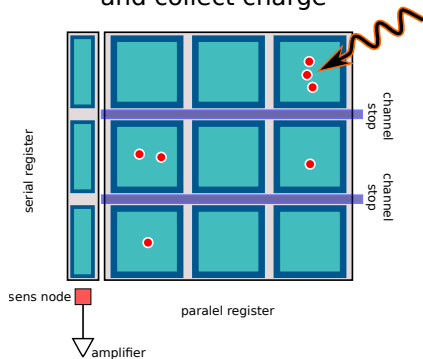
- **Fermilab:** Tiffenberg, Guardincerri, Sofo Haro
- **Stony Brook:** Rouven Essig
- **LBNL:** Steve Holland, Christopher Bebek
- **Tel Aviv University:** Tomer Volansky
- **CERN:** Tien-Tien Yu
- **Stanford University\*:** Jeremy Mardon

## Main goals

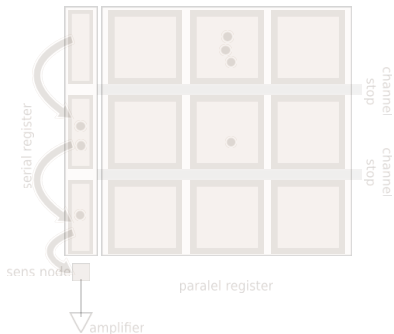
- Build the first working detector using Skipper-CCDs.
- Validate the technology for DM and  $\nu$  experiments.
- Probe DM masses at the MeV scale through electron recoil.
- Probe axion and hidden-photon DM with masses down to 1 eV.

## 3x3 pixels CCD

Expose the CCD to particles and collect charge

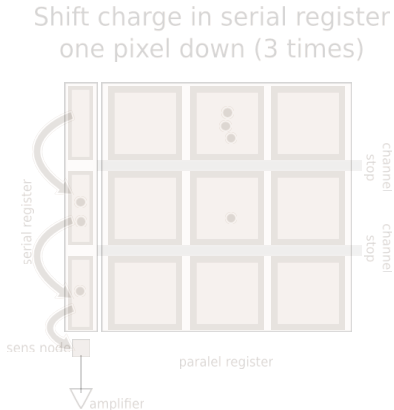
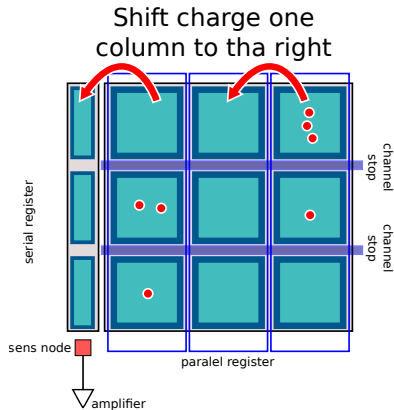


Shift charge in serial register one pixel down (3 times)

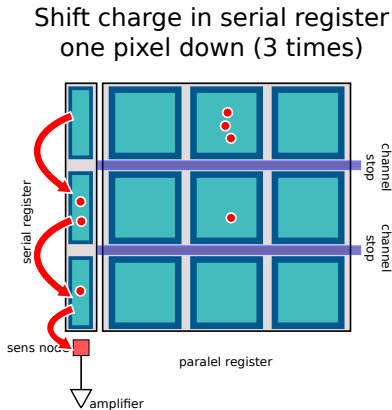
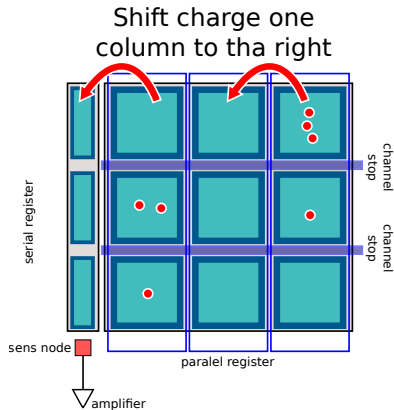




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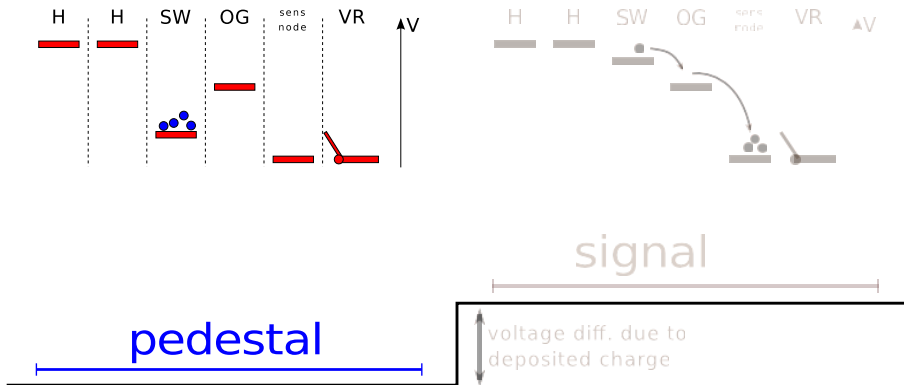


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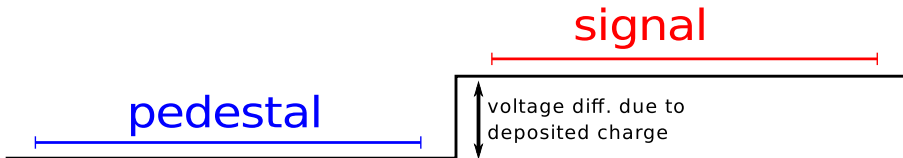
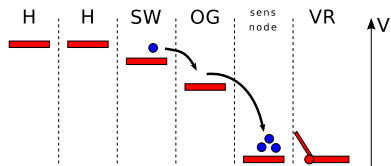
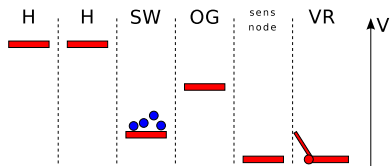


capacitance of the system is set by the SN:  $C=0.05\text{pF} \rightarrow 3\mu\text{V}/e$

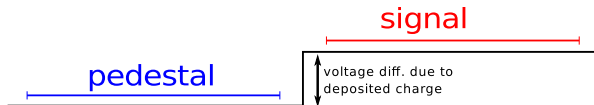
# CCD: readout



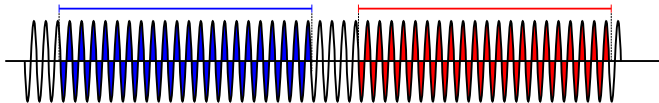
# CCD: readout



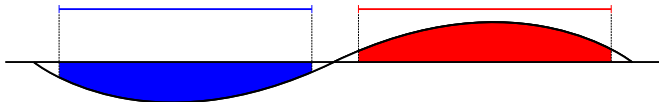
pixel charge  
measurement



high frequency  
noise

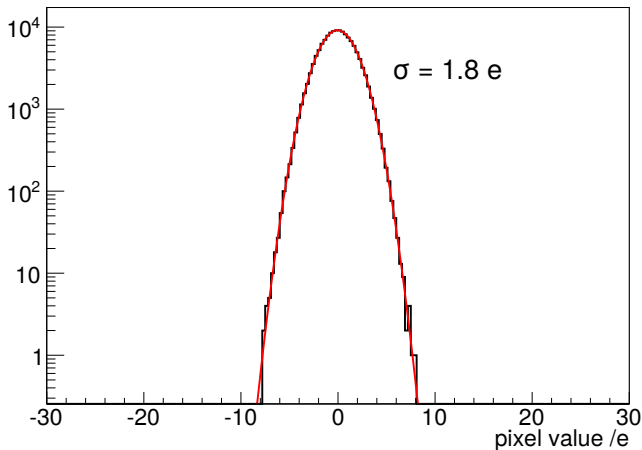


low frequency  
noise



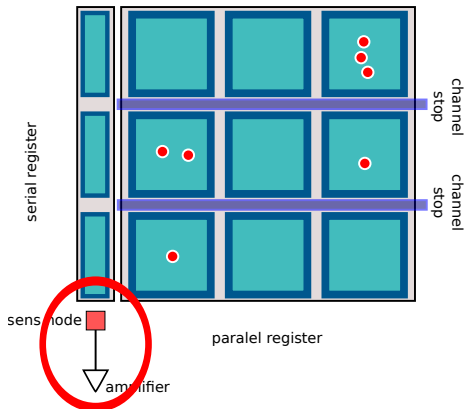
excellent for removing high frequency noise but sensitive to low frequencies

## Readout noise: empty pixels distribution, regular scientific CCD



**$2 \text{ e}^-$  readout noise roughly corresponds to 50 eV energy threshold**

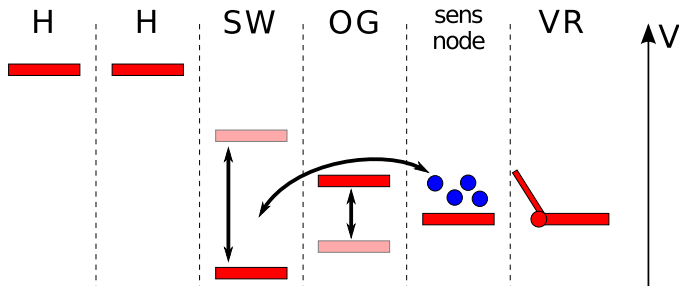
## Lowering the noise: Skipper CCD



**Only the readout stage is modified**

## Lowering the noise: Skipper CCD

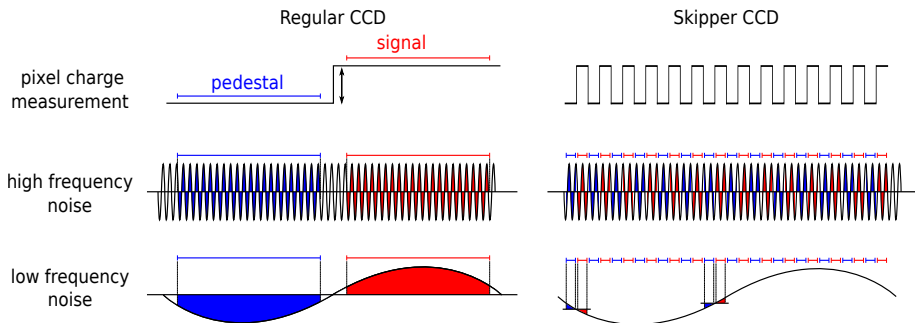
- **Main difference:** the Skipper CCD allows multiple sampling of the same pixel without corrupting the charge packet.
- The final pixel value is the average of the samples  
**Pixel value** =  $\frac{1}{N} \sum_i^N (\text{pixel sample})_i$
- Idea proposed in 1990 by Janesick et al. (doi:10.1117/12.19452)





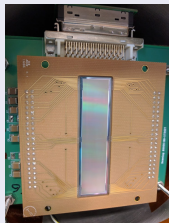
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# SENSEI: First working instrument using SkipperCCD tech

## Sensors



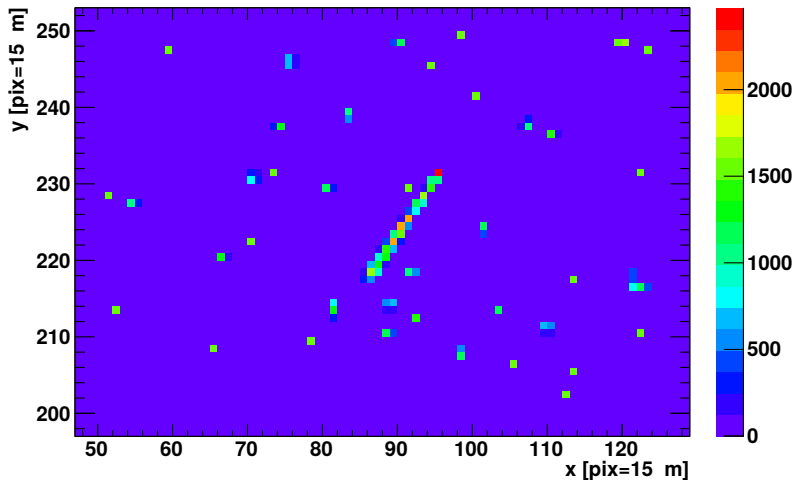
- Skipper-CCD prototype designed at LBL MSL
- 200 & 250  $\mu\text{m}$  thick, 15  $\mu\text{m}$  pixel size
- Two form factors 4k $\times$ 1k (0.5gr) & 1.2k $\times$ 0.7k pixels
- Parasitic run, optic coating and Si resistivity  $\sim 10\text{k}\Omega$
- 4 amplifiers per CCD, three different RO stage designs

## Instrument

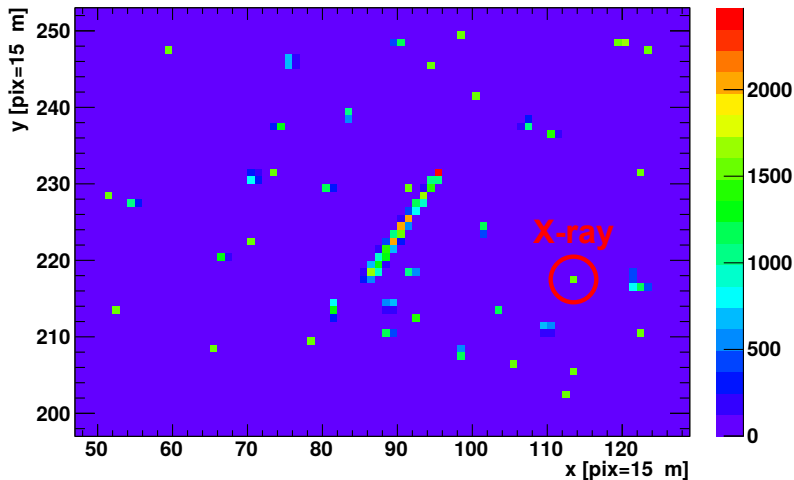


- System integration done at Fermilab
- Custom cold electronics
- Modified DES electronics for read out
- Firmware and image processing software
- Optimization of operation parameters

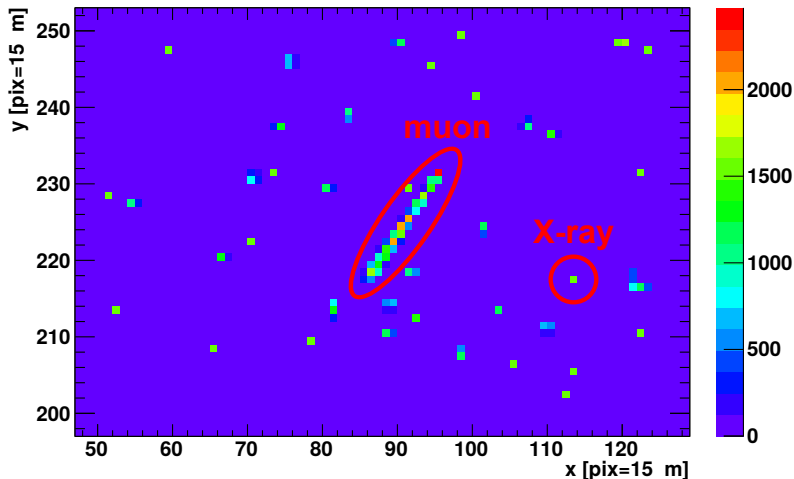
# Image taken with SENSEI: 4000 samples per pixel (processed)



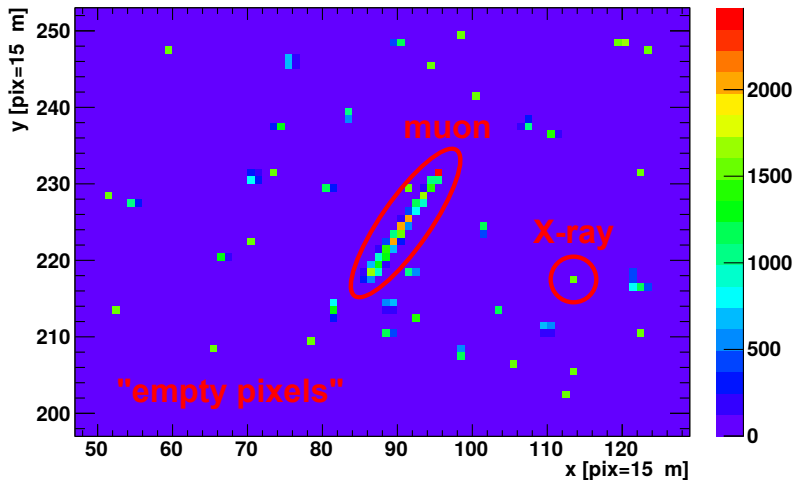
# Image taken with SENSEI: 4000 samples per pixel (processed)



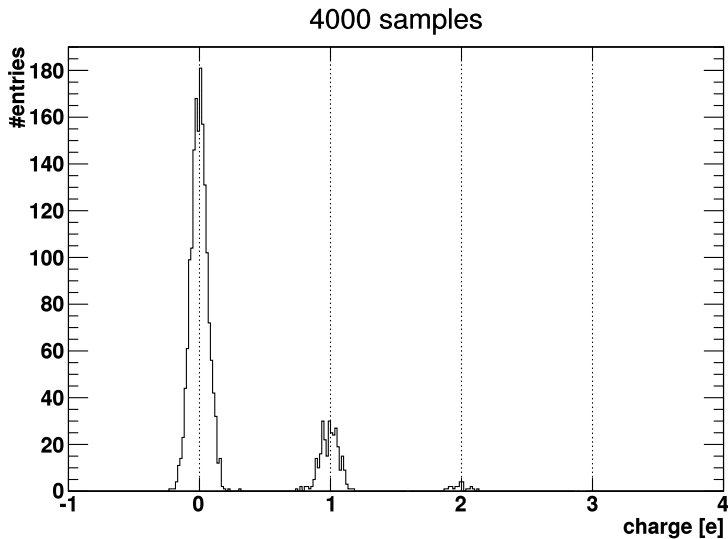
# Image taken with SENSEI: 4000 samples per pixel (processed)



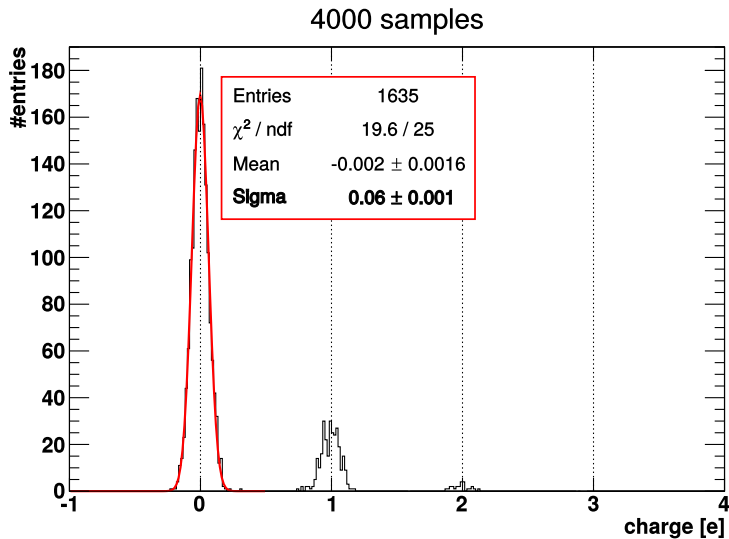
# Image taken with SENSEI: 4000 samples per pixel (processed)



## Charge in pixel distribution. Counting electrons: 0, 1, 2..



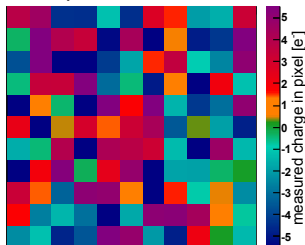
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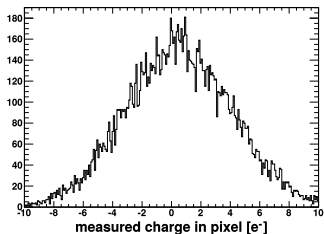


# Counting electrons: 0, 1, 2..

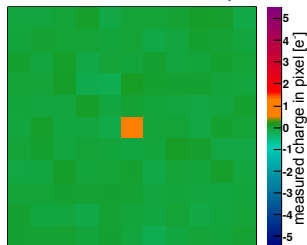
Standard CCD mode: charge in each pixel is measured once



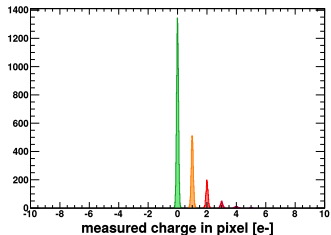
Readout-noise: 3.5 e RMS



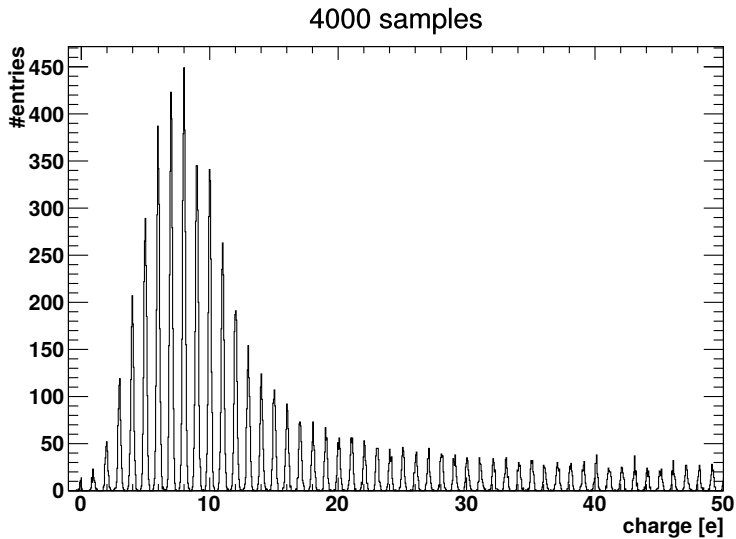
New Skipper CCD: charge in each pixel is measured multiple times

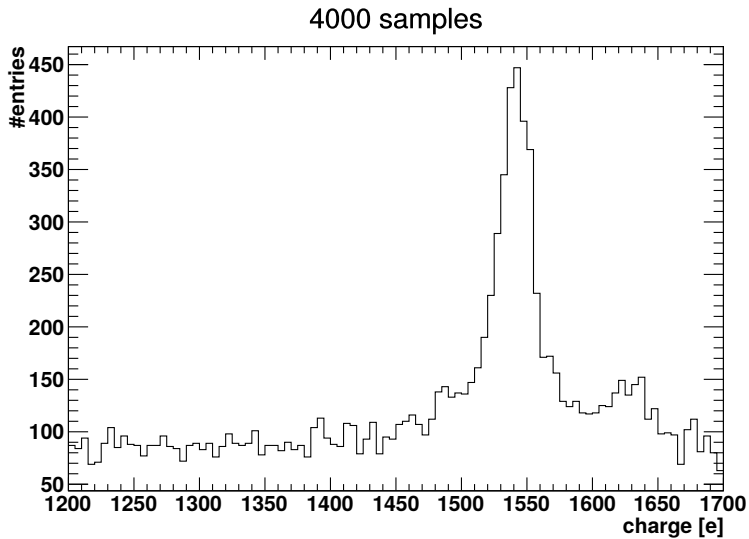


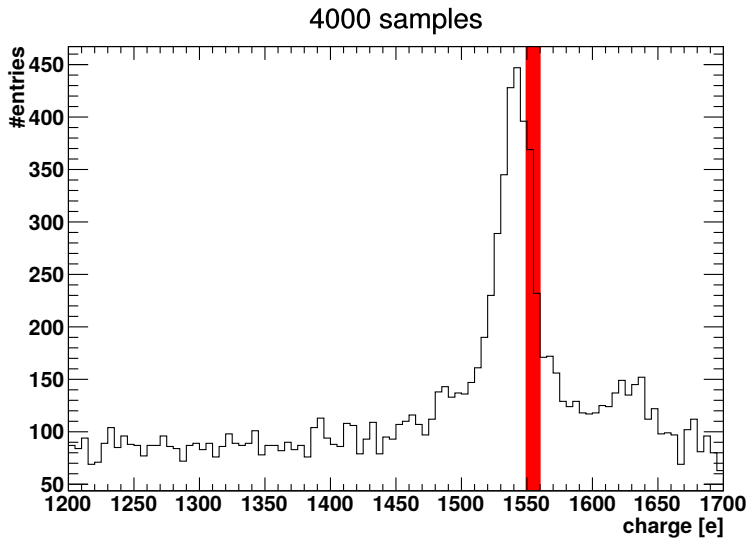
Readout-noise: 0.06 e RMS



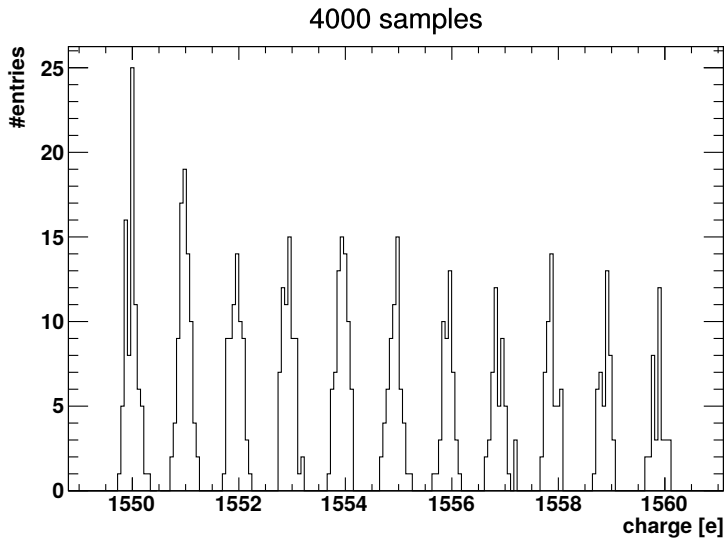
## Counting electrons: ..48, 49, 50..



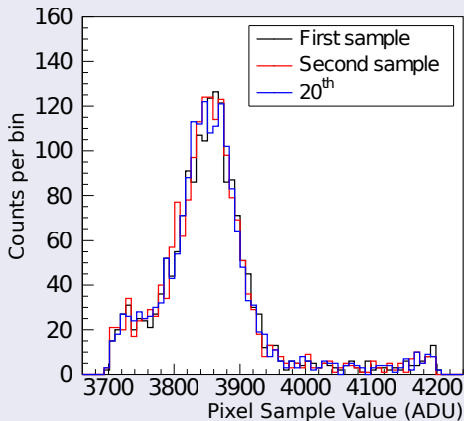




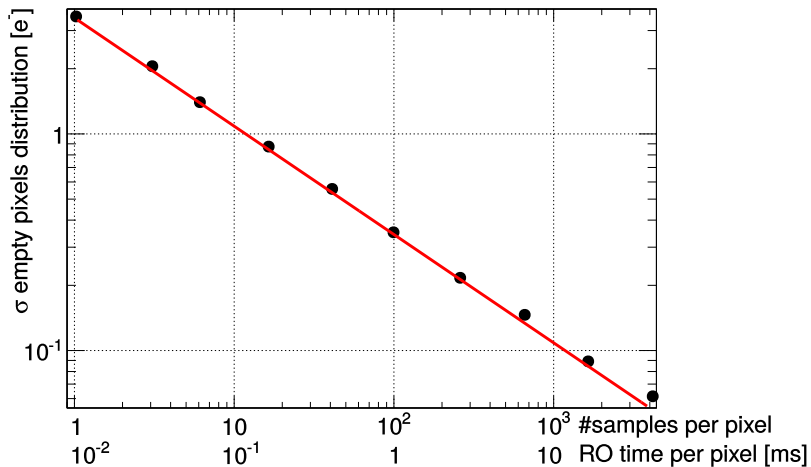
keep counting: ..1550, 1551, 1552..



### Single pixel distribution: X-rays from $^{55}\text{Fe}$



**The gain is the same for all the samples**



## SENSEI: DM search operation mode

- Counting electrons  $\Rightarrow$  **noise has zero impact**
- It can take about 1h to read the sensors
- Dark Current is the limiting factor**

It's better to readout continuously to minimize the impact of the DC

Dark Current [ $\text{e}^- \text{pix}^{-1} \text{day}^{-1}$ ]	$\geq 1\text{e}^-$ [pix]	$\geq 2\text{e}^-$ [pix]	$\geq 3\text{e}^-$ [pix]
$10^{-3}$	$1 \times 10^8$	$3 \times 10^3$	$7 \times 10^{-2}$
$10^{-5}$	$1 \times 10^6$	$3 \times 10^{-1}$	$7 \times 10^{-8}$
$10^{-7}$	$1 \times 10^4$	$3 \times 10^{-5}$	$7 \times 10^{-14}$

Measured upper limit for the DC in CCDs is:

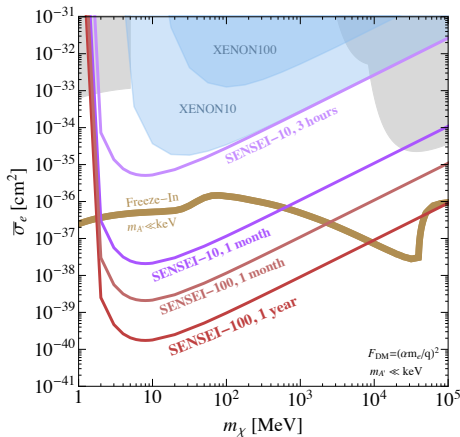
$$1 \times 10^{-3} \text{ e pix}^{-1} \text{day}^{-1} \quad \text{arXiv:1611.03066}$$

Could be orders of magnitude lower. **Theoretical prediction is  $O(10^{-7})$**

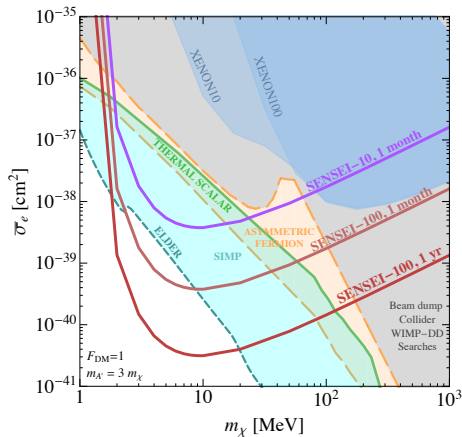


# SENSEI: reach of a 100g, zeroish-background experiment

## Light Dark Photon



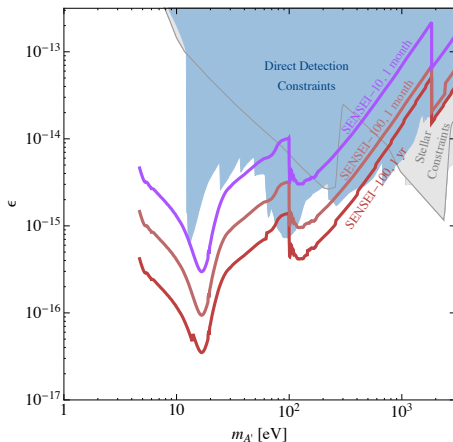
## Heavy Dark Photon



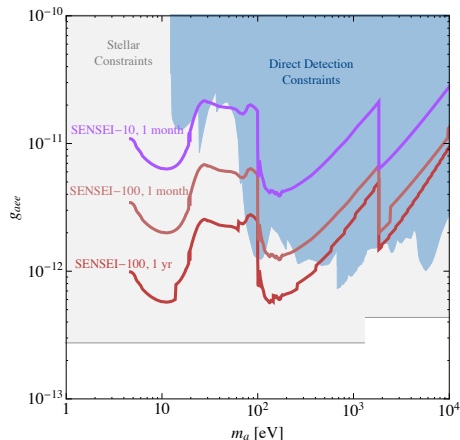
Rouven Essig, Tomer Volansky & Tien-Tien Yu.

# SENSEI: reach of a 100g, zeroish-background experiment

## Dark photon ( $A'$ )



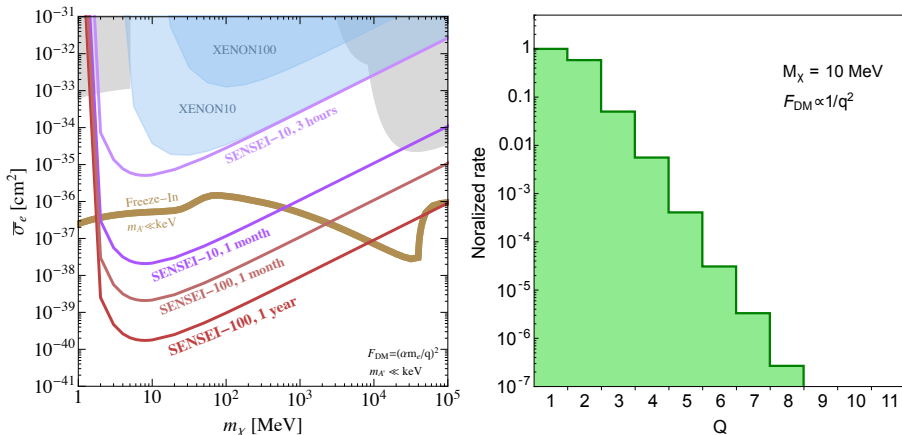
## Axion-like-particle (ALP)



Rouven Essig, Tomer Volansky & Tien-Tien Yu.

# SENSEI: electron recoil background requirements

The sensitivity is dominated by the lowest energy/charge bin



Rouven Essig, Tomer Volansky & Tien-Tien Yu.

### Back of the envelope calculation

A 100g detector that takes data for one year  $\rightarrow$  **Expo = 36.5kg  $\cdot$  day**

Assuming same background as in DAMIC:

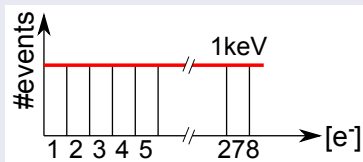
- **5 DRU** ( $\text{events} \cdot \text{kg}^{-1} \cdot \text{day}^{-1} \cdot \text{keV}^{-1}$ ) in the 0-1keV range  
 $\rightarrow$   **$N_{\text{bkg}} = 36.5 \text{ kg} \cdot \text{day} \times 5 \text{ DRU} = 182.5$  events**
- Dominated by external gammas  $\rightarrow$  **flat Compton spectrum**

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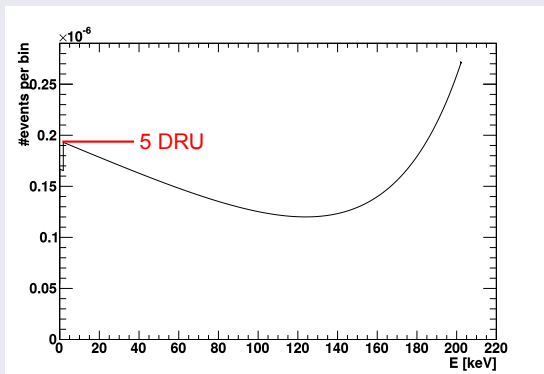
182.5 events over the 278 charge bins in the 0-1keV range

**Expect 0.65 bkd events in the lowest (2  $e^-$ ) charge-bin**

# Skipper CCD - electron recoil background requirements

A more detailed analysis: Klein-Nishina + binding energy correction

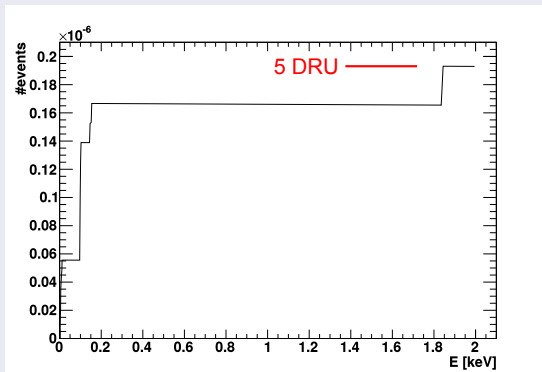
- at lower energies atomic binding energies are relevant
- partial energy depositions populate low E region (thin det)



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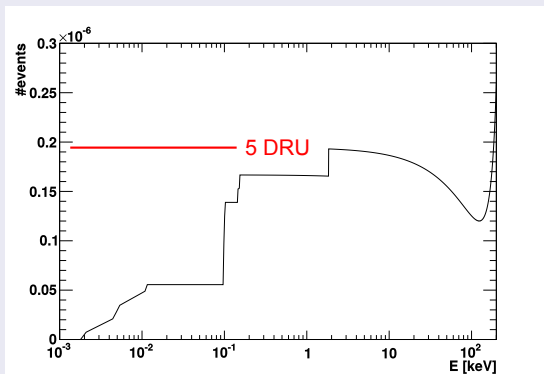
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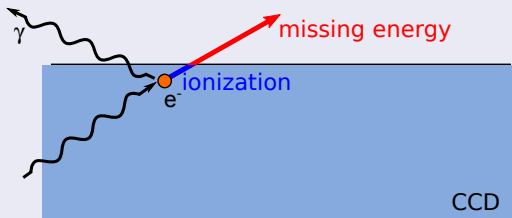




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A more detailed analysis: MC simulation, G4 3D Monash model

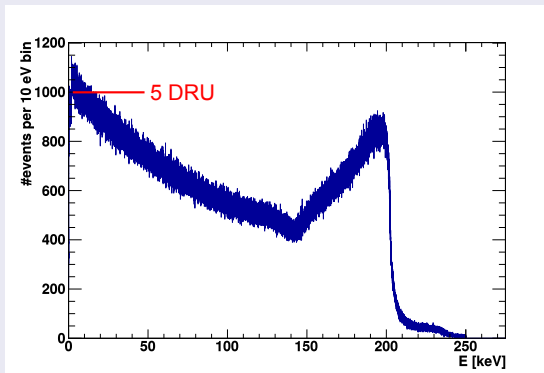
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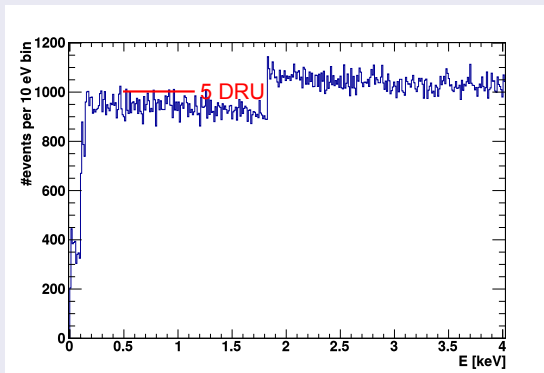
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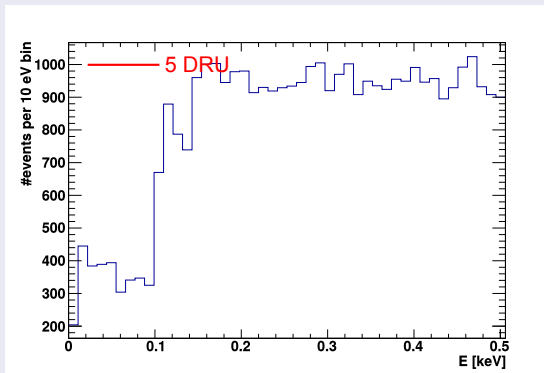
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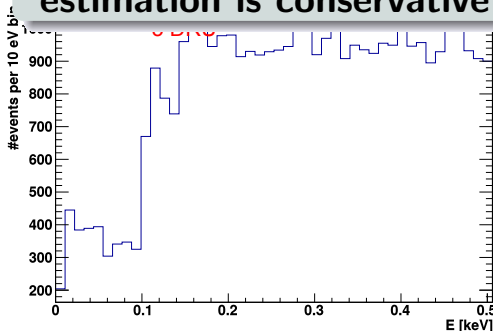


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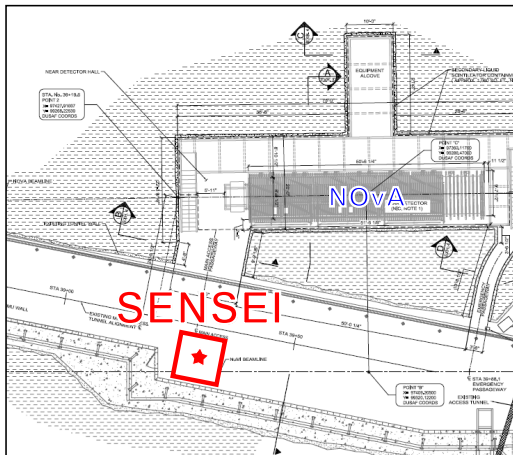
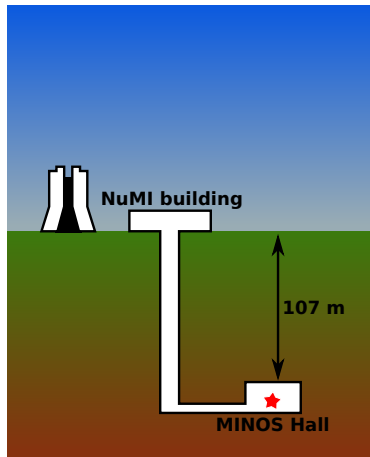
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**Back of the envelope  
estimation is conservative**

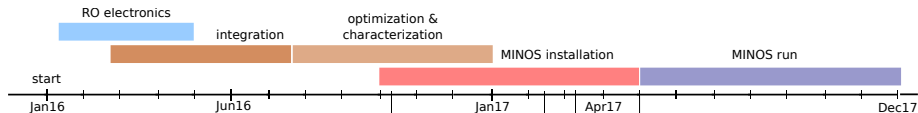


# Whats going on now: Installation @MINOS

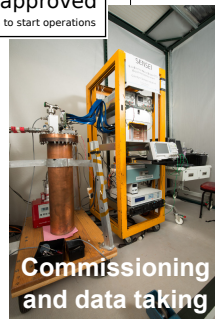
Technology demonstration: installation at shallow underground site



# Whats going on now: Installation @MINOS



TSW approved  
permission to start operations



**Taking data to understand if current (parasitically-fabricated) detectors are good enough to produce a science result**

## Timeline

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**2016**

LDRD funded,  
fabrication of SkipperCCD  
prototype

**2017**

testing of prototype,  
received funding from HSF  
for S-10 and S-100

**2018**

assembly and testing of S-10,  
take data

**2019**

take more data with S-10, begin analysis  
assembly and testing of S-100

**2020**

continue S-10 analysis,  
take data with S-100

**2021**

S-100 analysis



## Summary

- Demonstrated technology: working detector.
- Demonstrated bkg: no R&D needed.
  - ▶ required bkg level already reached by running experiments.
- Minimal R&D required for the packaging of the sensors.
- 10g & 100g desing/construction started.
  - ▶ Grant from Heising-Simons Foundation
  - ▶ Full technical support from Fermilab
- Complementary to LDMX.
- 10g Scientific Skipper-CCDs will start taking data at MINOS by the end of 2018.
- MINOS site is good up to a 10g experiment. Deeper location (Snolab/SURF) is required for 100g.