The Hunt for the Most Luminous Galaxies with WISE



Chao-Wei Tsai

The Hunt for the Most Luminous Galaxies







Peter Eisenhardt (JPL)

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IPAC/Caltech Charles Beichman Roc Cutri Chris Gelino Frank Masci Lin Yan **GSFC** Dominic Benford David Leisawitz Deborah Padgett

WISI



UCLA Graduate students, 2017, August (~ 10 years + 1 week ago) Around 5:00 pm

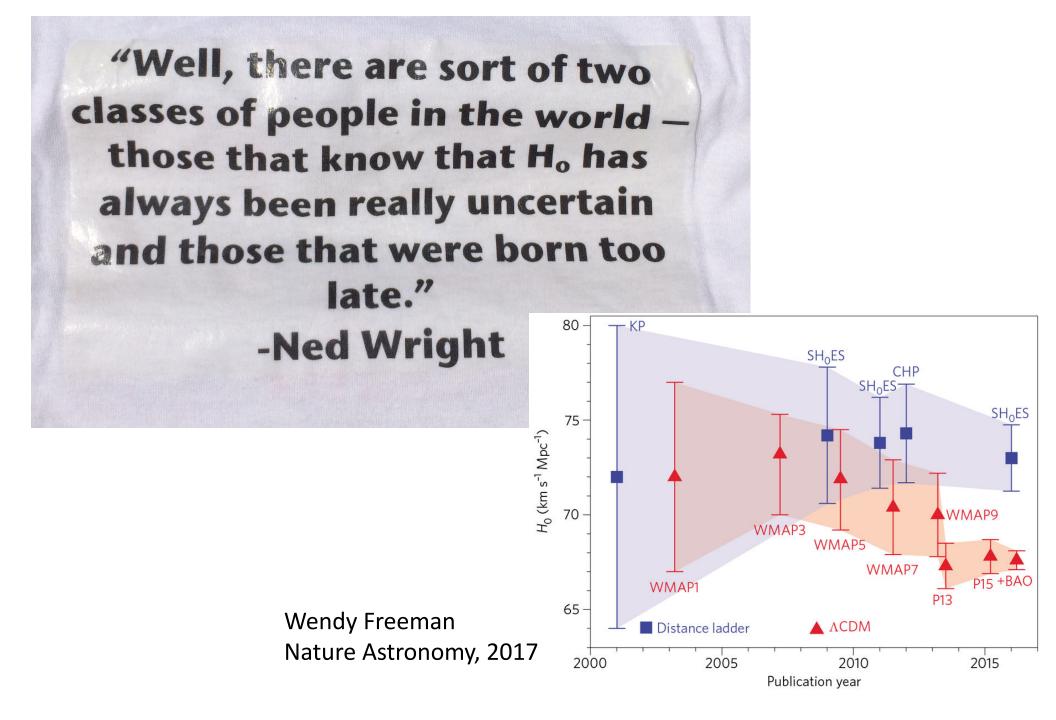


(Caltech)

(MMT)

Today





Wide-field Infrared Survey Explorer



Launched on December 14 2009, WISE mapped the sky in infrared light, searching for the nearest and coolest stars, the origins of stellar and planetary systems, and the most luminous galaxies in the Universe.

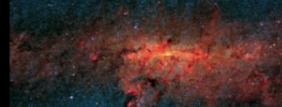
WISE delivered to the scientific community:

Over 1.5 million images covering the whole sky in 4 infrared wavelengths: 3.4, 4.6, 12, 22 micron

Catalogs of ~ 750 million objects seen in these 4 wavelengths









Galaxy

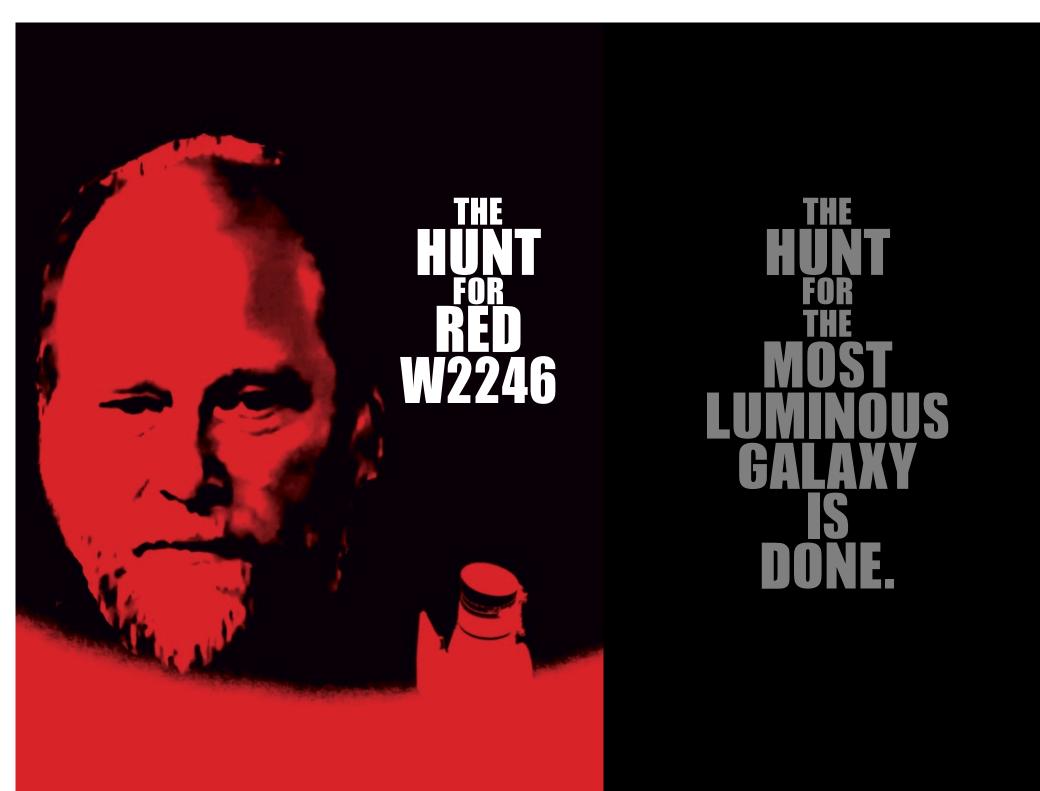
wise.astro.ucla.edu



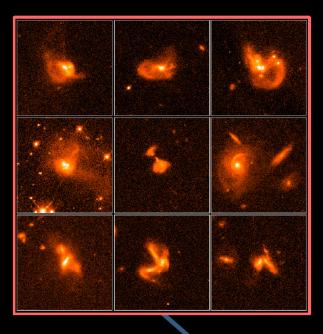
WISE All-Sky Data Release

WISE Science Data Center IPAC, Caltech March 14, 2012

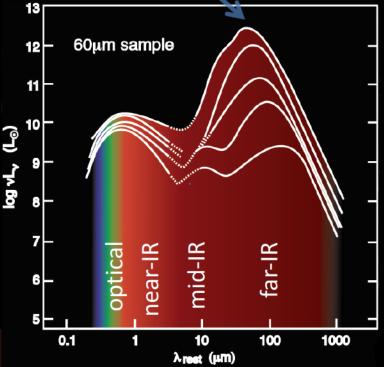








Spectral Energy Distribution (SED)



Infrared Astronomical Satellite Launched in 1983 Office Of

Discovery of Luminous Infrared Galaxies

- ~ 90 % of energy emitted at infrared wavelengths
- Total luminosities are ~ 10x 100x of Luminosity of Milky Way

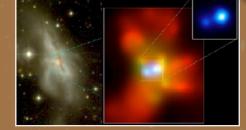
A bit of history – 1983

Merger Driven Galaxy Evolution

Coalescence/ULING



- BH grows rapidly
- Highly dust reddened
- dust/gas expelled

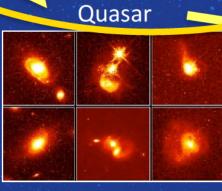


- gas inflows to center
- star formation triggered
- dusty starburst
- dominates luminosity

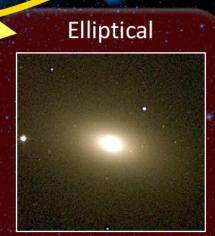
Galaxy Merger



- gas-rich merger



dust removedtraditional optical QSO

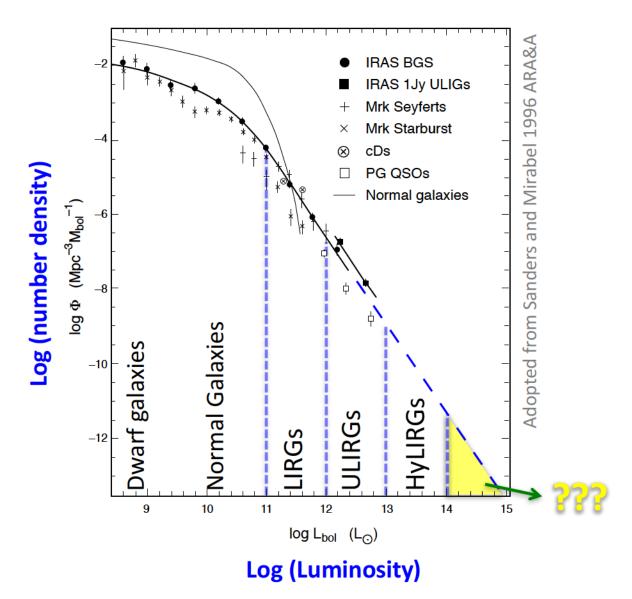


- red and dead

Sanders et al. (1988) Hopkins et al. (2008)



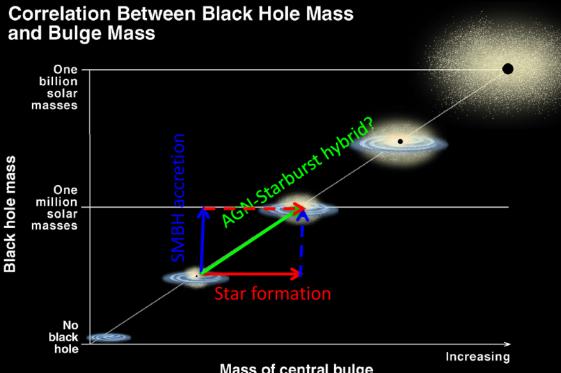
Tip of the Luminosity Function



Big Mystery of Galaxy Evolution

High luminosity – rapid mass growth phase

- Starburst dominated
- Rapid and violate star formation
- AGN dominated
- Powerful supermassive



Mass of central bulge

In the Most Luminous Galaxies,

mass of host galaxy or super massive black hole (SMBH) can

increase by a factor of 10 in ~ 100 - 300 Myr!



Tip of the Luminosity Function

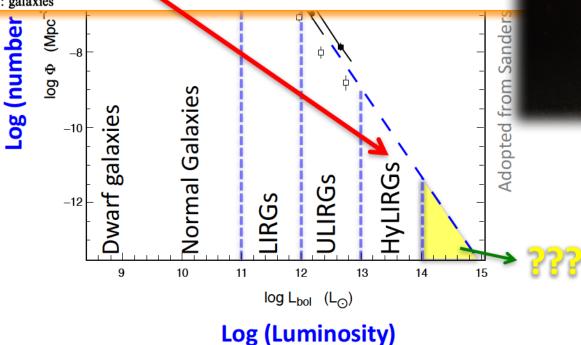
THE ASTROPHYSICAL JOURNAL, 424: L65–L68, 1994 April 1 © 1994. The American Astronomical Society. All rights reserved. Printed in U.S.A.

IRAS F15307+3252: A HYPERLUMINOUS INFRARED GALAXY AT $z = 0.93^{1}$ ROC M. CUTRI,² JOHN P. HUCHRA,³ FRANK J. LOW,² ROBERT L. BROWN,⁴ AND PAUL A. VANDEN BOUT⁴ Received 29 October 1993; accepted 1994 January 6

ABSTRACT

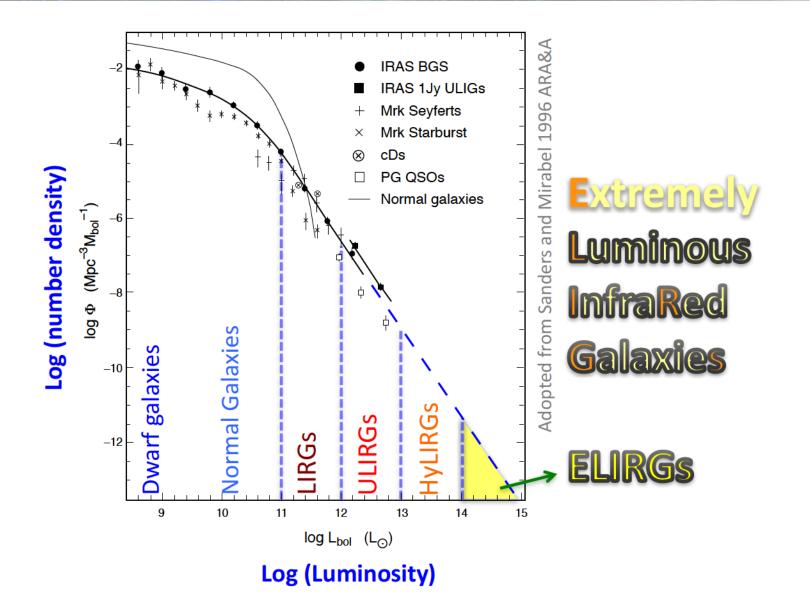
A new hyperluminous infrared galaxy in the *IRAS* Faint Source Catalog is identified at a redshift of 0.93. This object has a bolometric luminosity of $\sim 10^{13} L_{\odot}$, a very large ratio of infrared-to-optical luminosity, warm dust emission, a rate of infrared-to-radio flux densities consistent with other infrared galaxies, and an optical spectrum similar to a Seyfert 2 galaxy. IRAS F15307+3552 shares these characteristics and its radio-to-optical spectral energy distribution with two other infrared galaxies, F10214+4724 and P09104+4109. Discovery of a third object with these properties defines an extreme subclass of ultraluminous galaxies powered primarily by star formation. The systematic method used to find this object begins the process of determining the space density of these most luminous examples of the infrared galaxy phenomenon.

Subject headings: galaxies: individual (IRAS F15387+3252) — galaxies: photometry — galaxies: starburst — infrared: galaxies

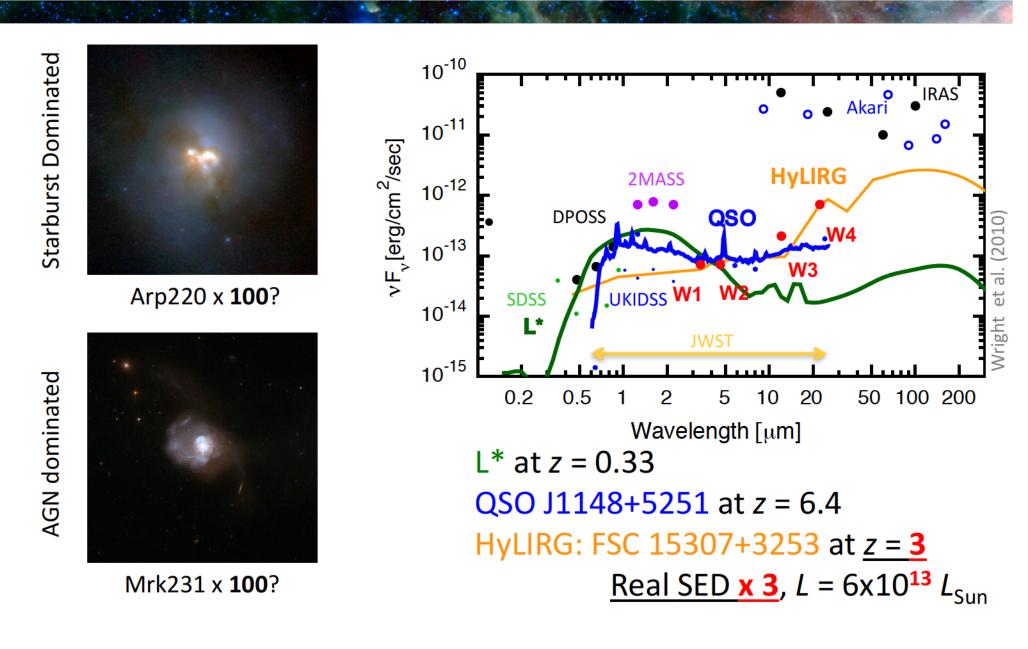




Tip of the Luminosity Function – Extreme Luminous Infrared Galaxies (ELIRGs)

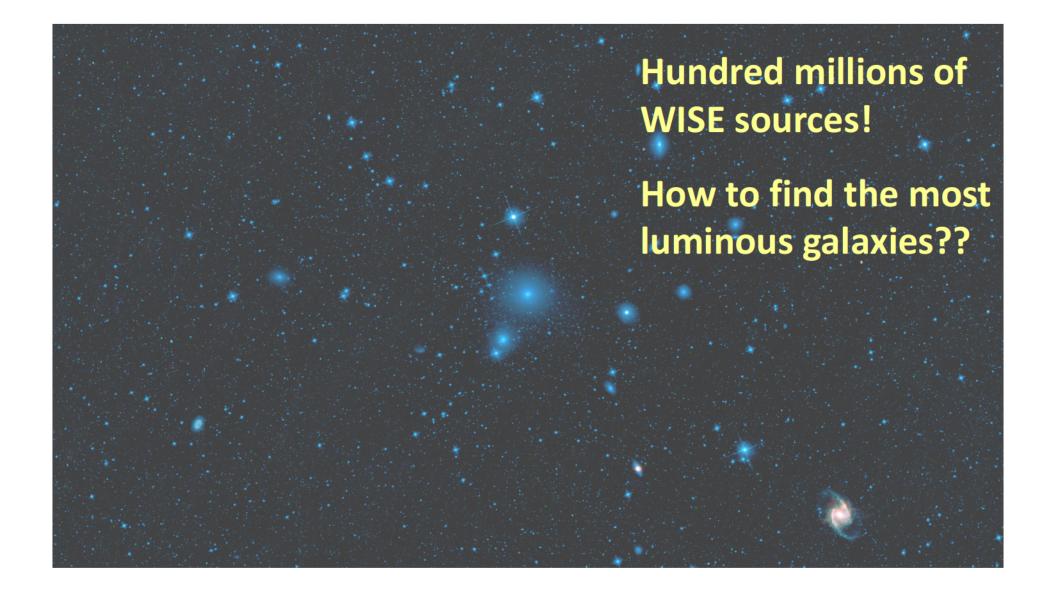


Hunt for **ELIRGs** – with **WISE**

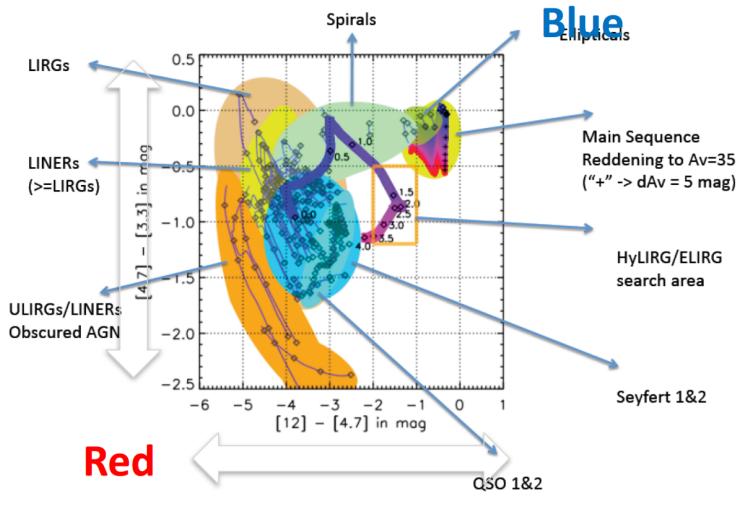




Ah... How to Identify ELIRGs?



WISE Colors of U/Hy/E LIRGs

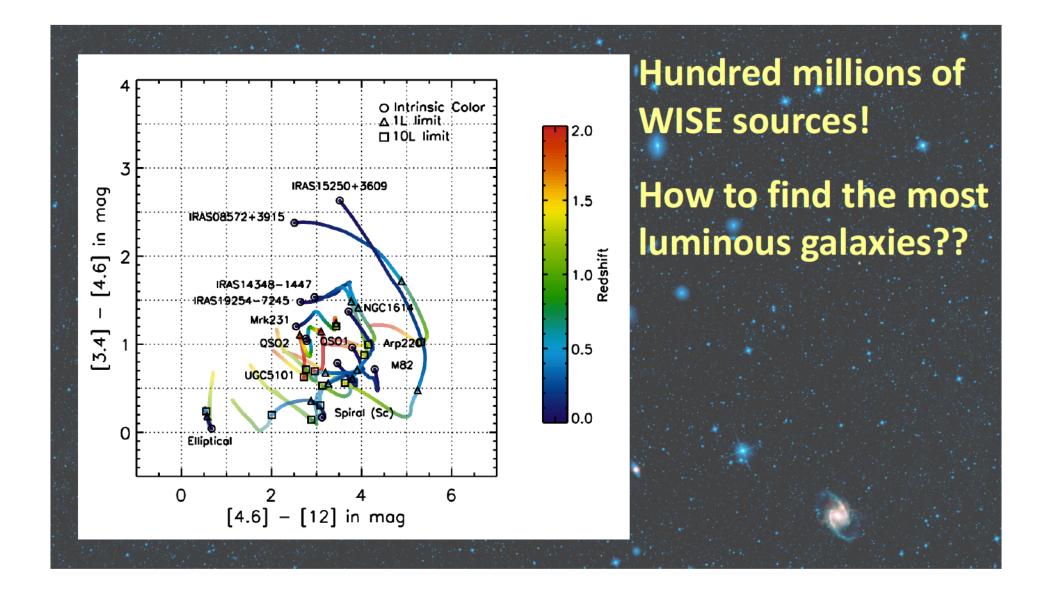


Based on GRASIL SED models and SWIRE SED templates

Chao-Wei Tsai - 2010 WISE Science Team Meeting

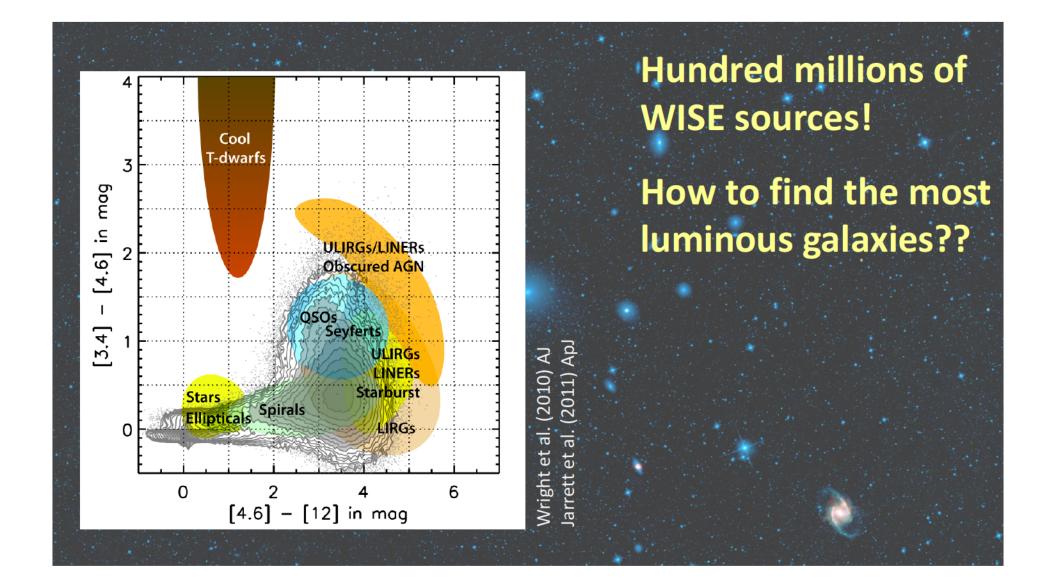


Ah... How to Identify ELIRGs?





Inhabitants of WISE Color Space



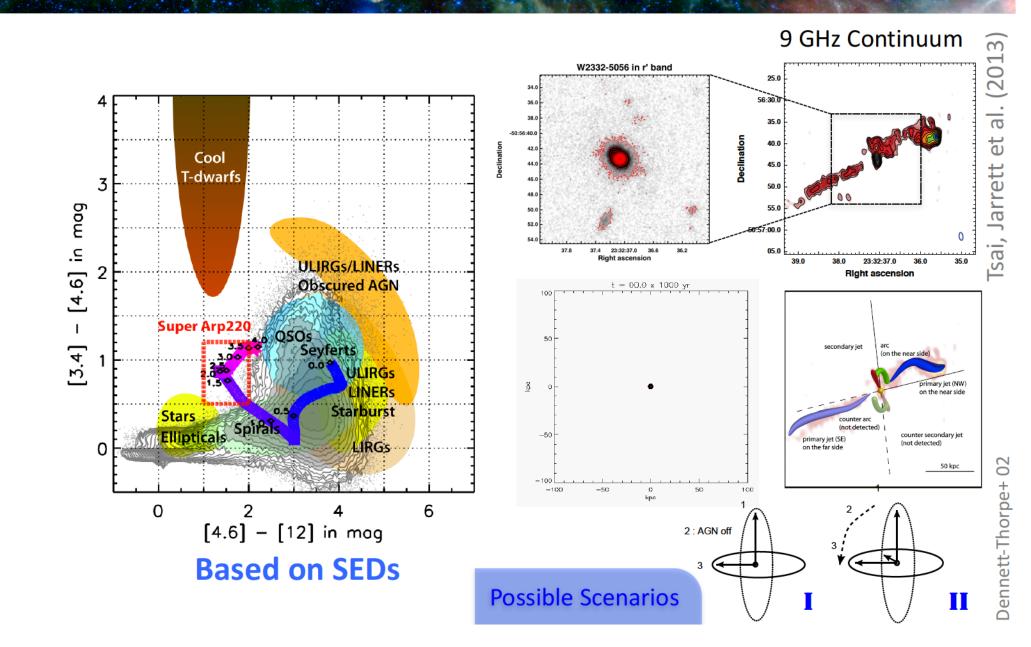
doi:10.1088/0004-6256/140/6/1868

THE ASTRONOMICAL JOURNAL, 140:1868–1881, 2010 December © 2010. The American Astronomical Society. All rights reserved. Printed in the U.S.A.

THE WIDE-FIELD INFRARED SURVEY EXPLORER (WISE): MISSION DESCRIPTION AND INITIAL ON-ORBIT PERFORMANCE

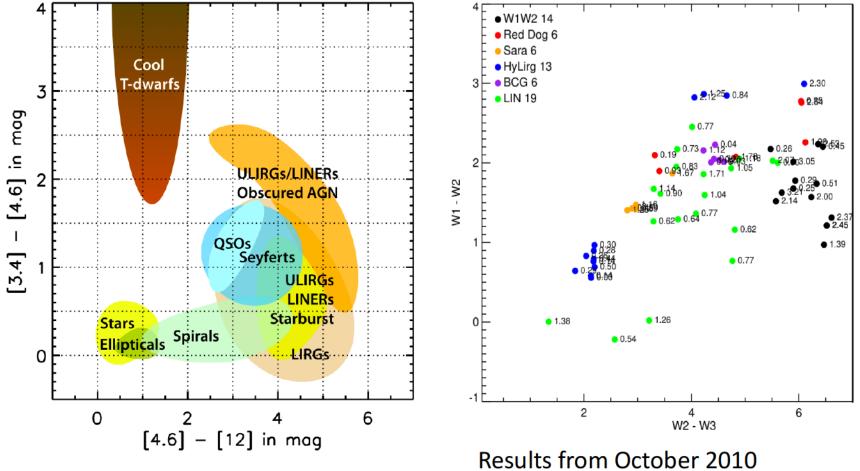
EDWARD L. WRIGHT¹, PETER R. M. EISENHARDT², AMY K. MAINZER², MICHAEL E. RESSLER², ROC M. CUTRI³, THOMAS JARRETT³, J. DAVY KIRKPATRICK³, DEBORAH PADGETT³, ROBERT S. MCMILLAN⁴, MICHAEL SKRUTSKIE⁵, S. A. STANFORD^{6,7}, MARTIN COHEN⁸, RUSSELL G. WALKER⁸, JOHN C. MATHER⁹, DAVID LEISAWITZ⁹, THOMAS N. GAUTIER III², IAN MCLEAN¹, DOMINIC BENFORD⁹, CAROL J. LONSDALE¹⁰, ANDREW BLAIN¹¹, BRYAN MENDEZ¹², WILLIAM R. IRACE², VALERIE DUVAL², FENGCHUAN LIU², DON ROYER², INGOLF HEINRICHSEN², JOAN HOWARD¹³, MARK SHANNON¹³, MARTHA KENDALL¹³, AMY L. WALSH¹³, MARK LARSEN¹⁴, JOEL G. CARDON¹⁴, SCOTT SCHICK¹⁵, MARK SCHWALM¹⁶, MOHAMED ABID², BETH FABINSKY², LARRY NAES^{17,18}, AND CHAO-WEI TSAI³ ¹ UCLA Astronomy, P.O. Box 951547, Los Angeles, CA 90095-1547, USA; wright@astro.ucla.edu ² Jet Propulsion Laboratory, 4800 Oak Grove Drive, Pasadena, CA 91109, USA ³ Infrared Processing and Analysis Center, California Institute of Technology, Pasadena, CA 91125, USA ⁴ University of Arizona, 1629 East University Boulevard, Tucson, AZ 85721, USA ⁵ Department of Astronomy, University of Virginia, Charlottesville, VA 22903, USA ⁶ Physics Department, University of California, Davis, CA 95616, USA ⁷ Institute of Geophysics and Planetary Physics, LLNL, Livermore, CA 94551, USA ⁸ Monterey Institute for Research in Astronomy, 200 8th Street, Marina, CA 93933, USA ⁹ NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA ¹⁰ National Radio Astronomy Observatory, Charlottesville, VA 22903, USA ¹¹ California Institute of Technology, Pasadena, CA 91125, USA ¹² Space Sciences Laboratory, University of California, Berkeley, CA 94720, USA ¹³ Ball Aerospace & Technologies Corporation, 1600 Commerce Street, Boulder, CO 80301, USA ¹⁴ Space Dynamics Laboratory, 1695 North Research Park Way, North Logan, UT 84341, USA ¹⁵ Practical Technology Solutions, Inc., P.O. Box 6336, North Logan, UT 8434, USA ¹⁶ L-3 Communications SSG-Tinsley, Wilmington, MA 01887, USA ¹⁷ Lockheed Martin Advanced Technology Center, Palo Alto, CA 94304, USA Received 2010 July 25; accepted 2010 September 22; published 2010 November 9

Searching for Most Luminous Galaxies



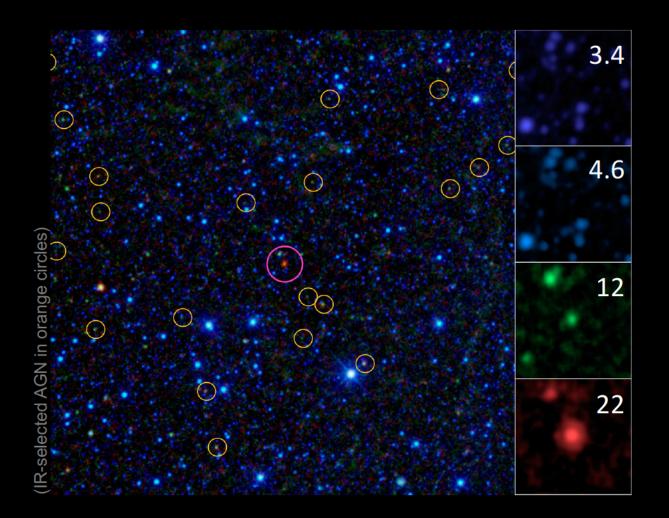


Inhabitants of WISE Color Space



Keck LRIS Optical Spectroscopic Follow-up

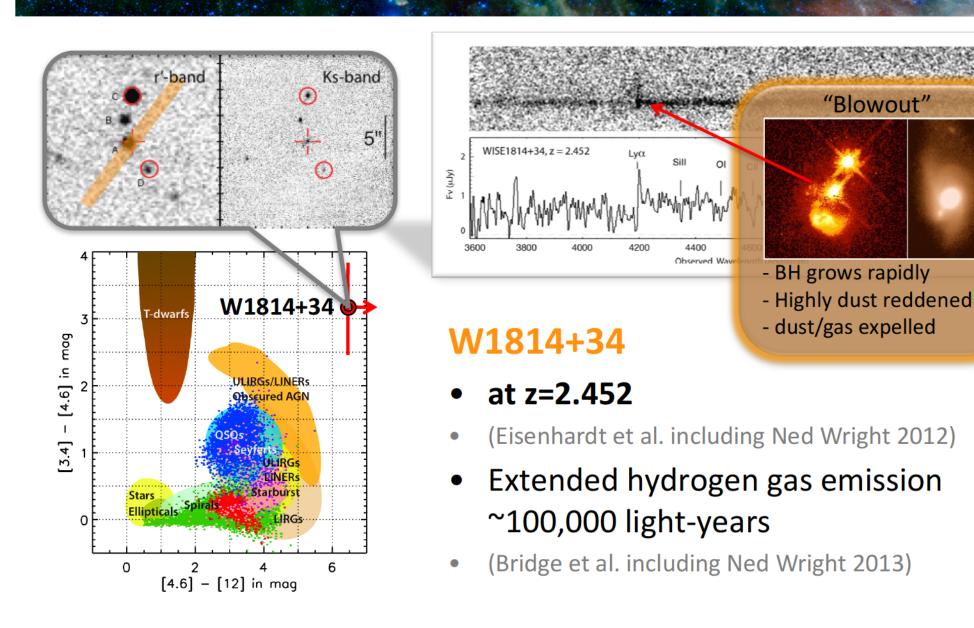
W1W2 Dropouts



WISE objects well detected at 12 μ m and 22 μ m faint or invisible at 3.4 μ m and 4.6 μ m



First HyLIRG Discovered by WISE

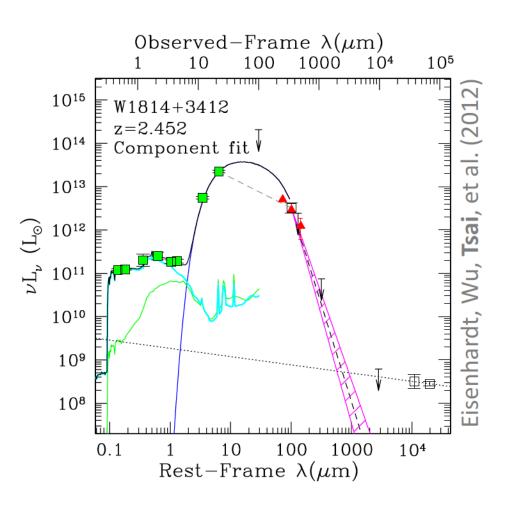


SED Decomposition of W1814+34

- $L_{\rm tot} = 4 8 \times 10^{13} L_{\odot}$
 - Very difficult to be powered by starburst
 - Implies supermassive black hole
 (~ 10^9 M_☉)
- Obscured AGN
 - i.e. visible light obscured by dust
- Starburst

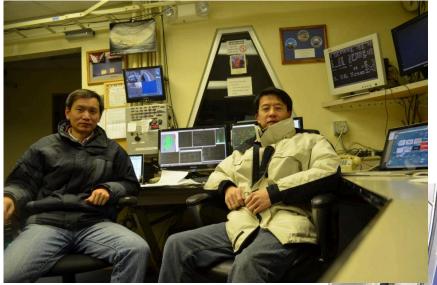
WISE

- SFR $\sim 500~M_{\odot}/yr$ (10% of $L_{bol})$
- Spiral Galaxy
 - Faint compared to AGN
 - Implies relatively low ratio of star mass to black hole mass compared to today



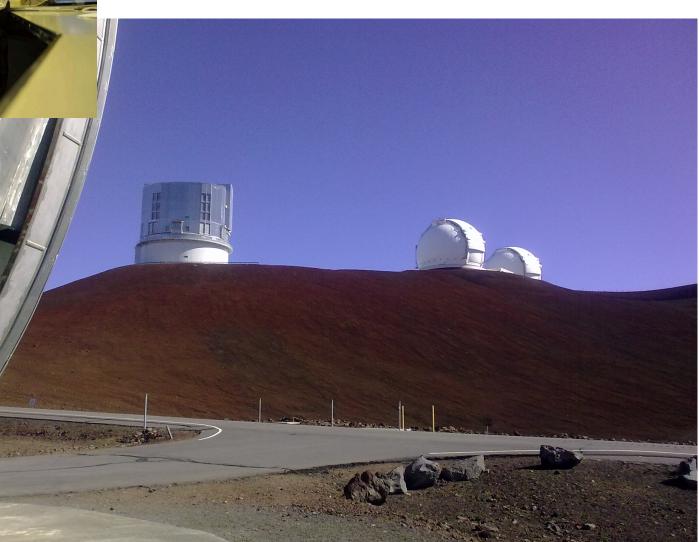
Follow-up Observation Programs

X-ray	UV –Optical	Near-IR	Mid-IR	Far-IR	Sub-mm – mm	Radio
NuSTAR	HST	HST	WISE & NEOWISE	Herschel	CSO	VLA
Chandra	Keck	Keck	Spitzer		JCMT	ATCA
XMM	VLT	VLT			SMA	GBT
Suzaku	Gemini N/S	Gemini N/S				
	Palomar	Palomar			CARMA	
	SOAR	SOAR			IRAM30m	
	WYIN	WYIN			PdBI	
	Lick	Lick			ALMA	
	Magellan	Bok				



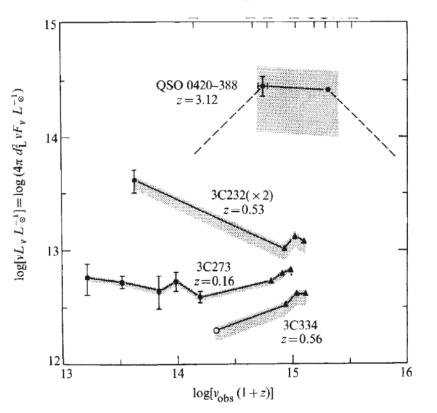
2010-2012 CSO submillimeter Follow-up

Often getting the optical spectroscopic data from Keck at the same time

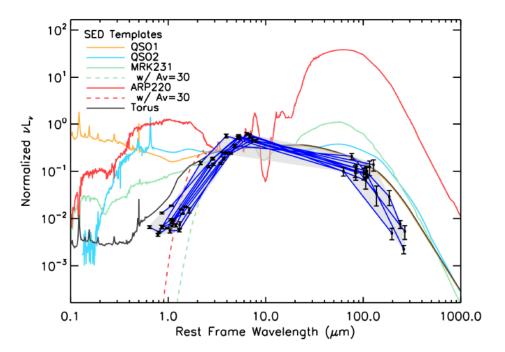


Far-IR Follow-up on W1W2 Dropouts

Infrared observations of the most luminous quasar



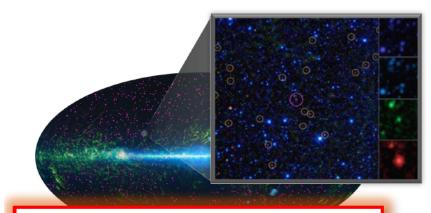
Submm follow-up of WISE-selected Hyperluminous Galaxies



Wu et al. including Ned Wright (2012)

Wright and Kleinmann (1978, Nature)



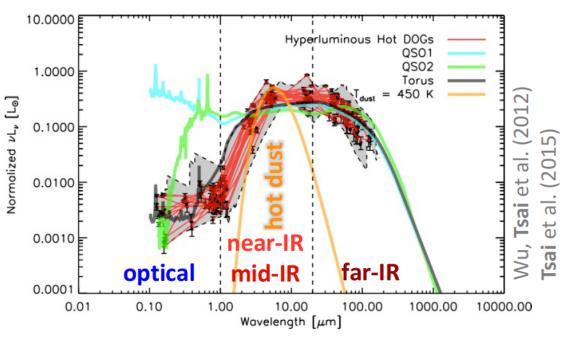


Hot Dust Obscured Galaxies

– Hot DOGs



(Dog Haus Pasadena)

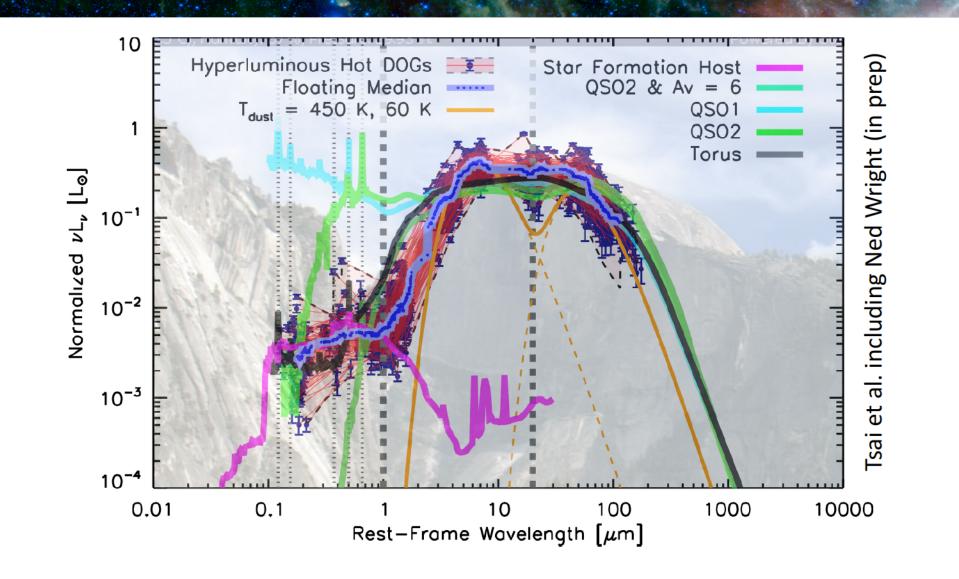


- Hyperluminous: $L > 10^{13} L_{Sun}$
- mid-IR excess:

hot dust emission ($T_{\rm D}$ ~ 450 K)

 Hot Dust Obscured Galaxies (Hot DOGs)





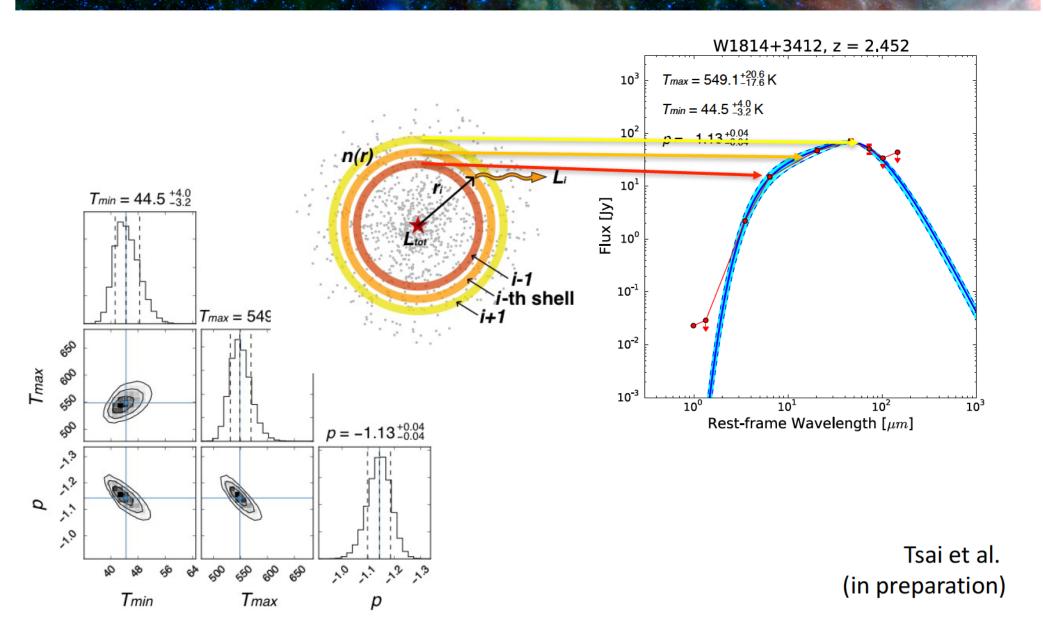
Yosemite "Half-Dome" shaped SED





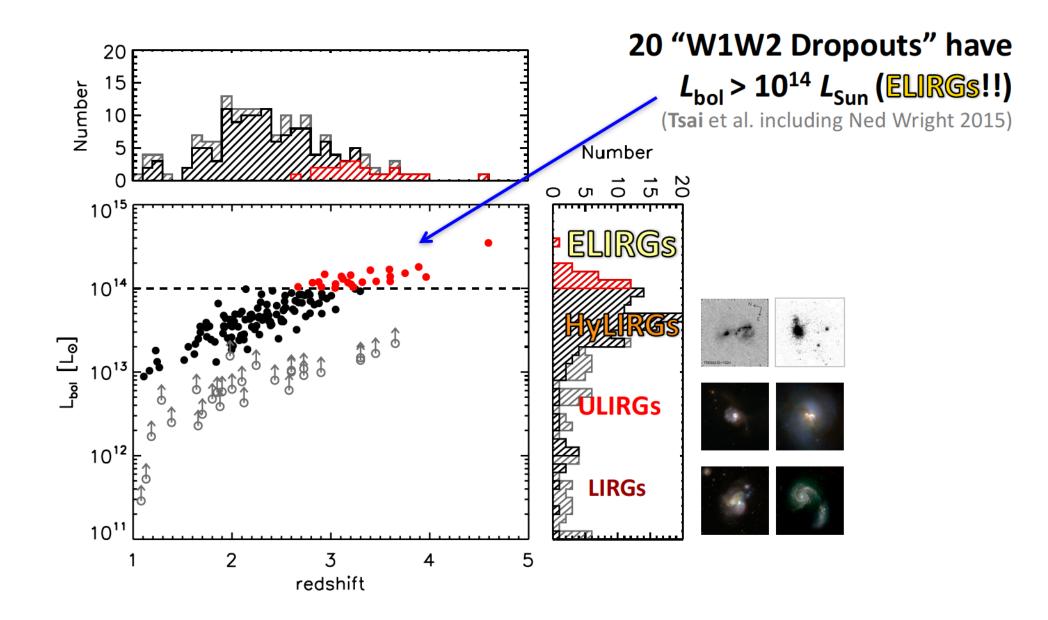
Yosemite "Half-Dome" shaped SED





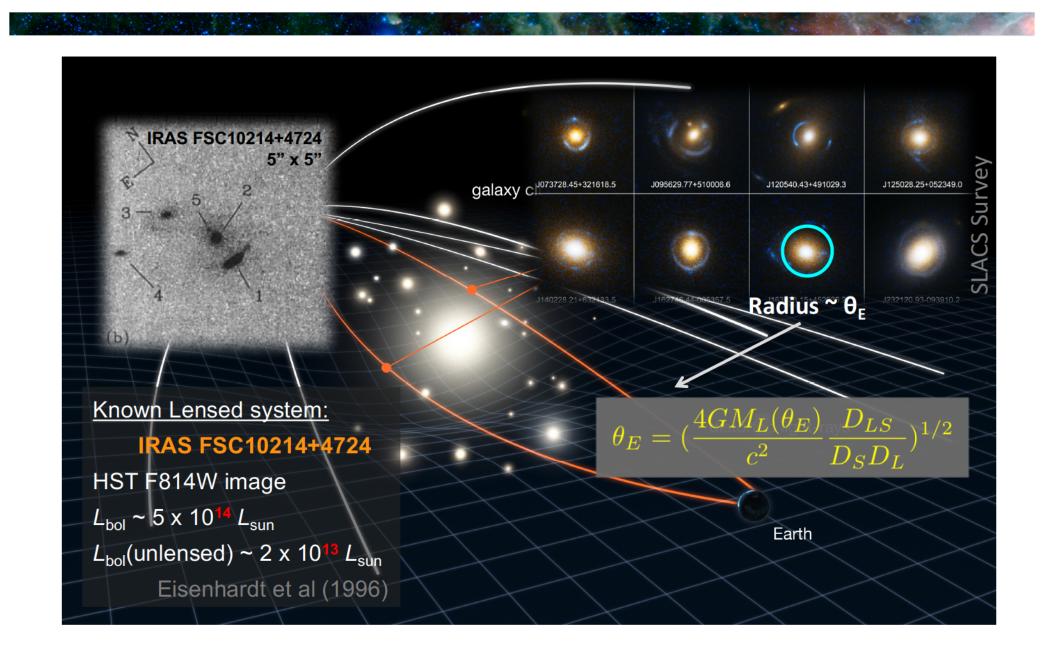


Mission Accomplished!



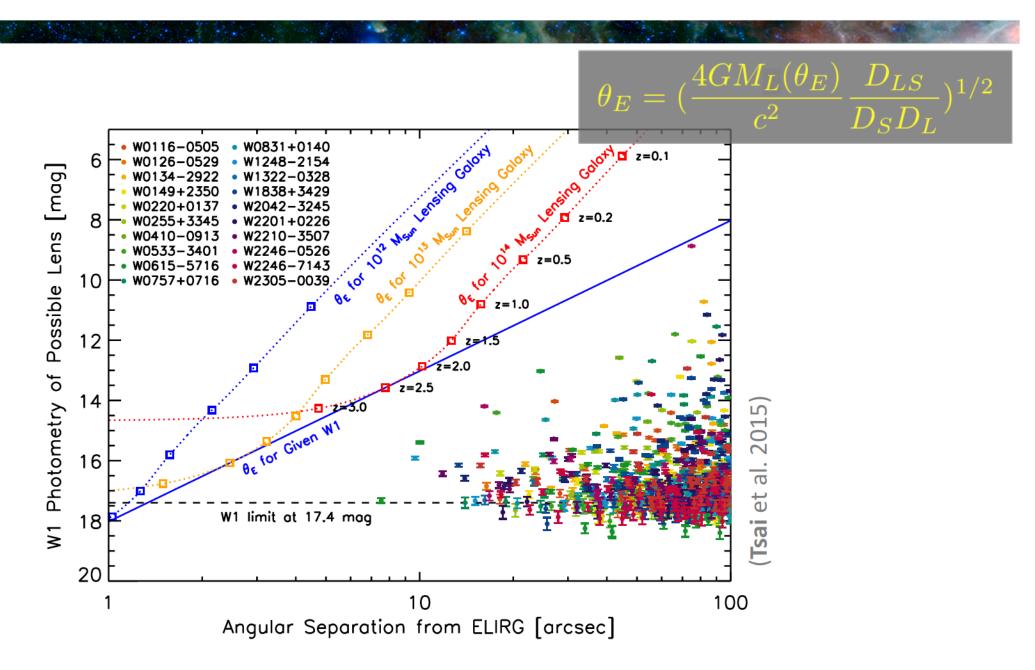


Gravitationally lensed?



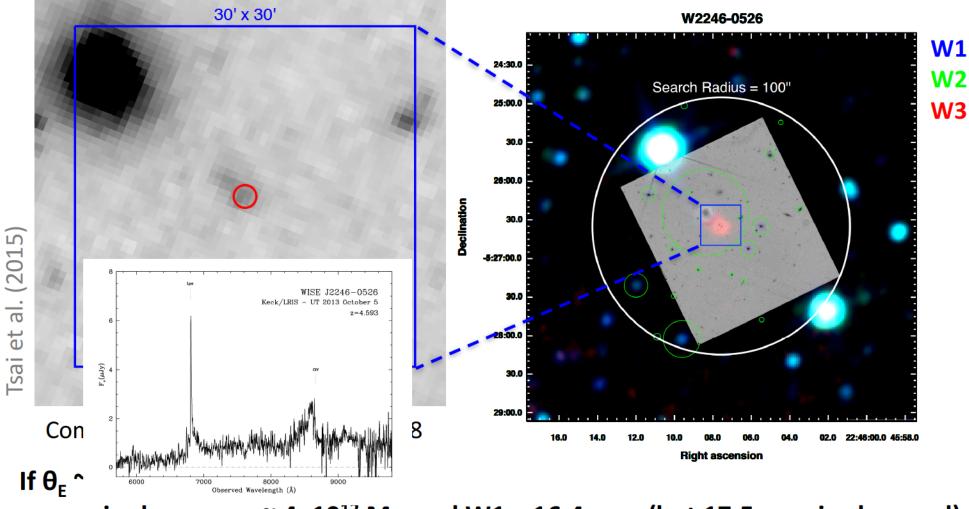


Possible Foreground Lensing Galaxies at Separation > WISE W1 Resolution





HFSACF11 (530 %) image



-> require lens mass ~ 4×10^{12} M_{\odot} and W1 < 16.4 mag (but 17.5 mag is observed)



JET PROPULSION LABORATORY DISCOVERY

The most luminous galaxy found

LA CAÑADA FLINTRIDGE » A distant galaxy sizzling with light from more than 300 trillion stars has been named the most luminous galaxy, NASA's Jet Propulsion Laboratory announced Thursday.

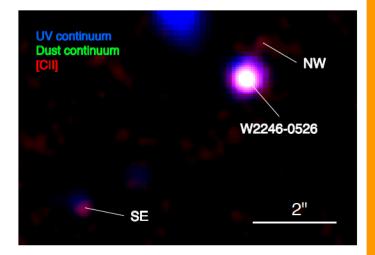
NASA's Wide-field Infrared Survey Explorer spacecraft, managed and operated by JPL, discovered the bright galaxy, which belongs to a new class of objects called extremely luminous infrared galaxies. The light may originate from a major growth spurt of the galaxy's black hole, said Chao-Wei Tsai of JPL and lead author of a report published Friday.

by Zen Vuong, Pasadena Star-News Based on NASA Press-Release 15-095 and **Tsai et al. 2015, ApJ**



NASA/JPL-CALTECH

This artist's concept depicts the current record holder for the most luminous galaxy in the universe. The galaxy, WISE J224607.57-052635.0, is erupting with light equal to more than 300 trillion suns. It was discovered using data from NASA's WISE mission.

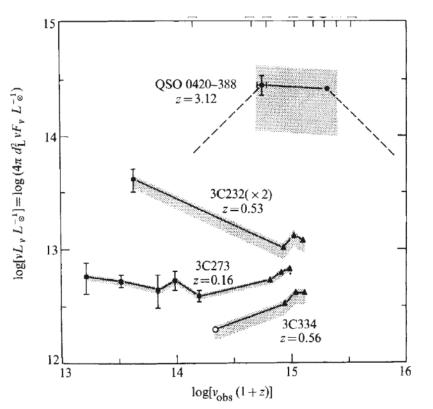


ALMA and HST images of W2246-0526 Diaz-Santos et al. 2016, ApJL



About 40 Years Ago

Infrared observations of the most luminous quasar

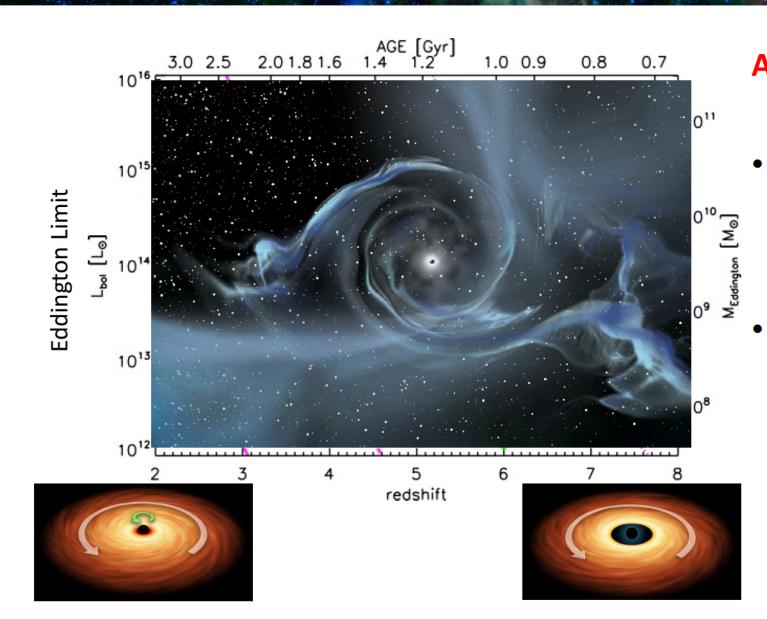


Wright and Kleinmann (1978, Nature)

- With an assumed spectral index and moderm cosmology:
 - Lbol ~ 3.0 x 10¹⁴ L_{Sun}
 - LIR ~ 1.7–4 x 10¹⁴ L_{sun}
 - Could qualify as an ELIRG.
- After adding WISE data, the conservative estimate:
 - L_{bol} ~ 1.2 x 10¹⁵ L_{sun}
 - L_{IR} (8-1000µm) ~ 2.0 x 10¹⁴ L_{Sun}
 - No Far-IR data available
 - It IS an ELIRG!!



SMBH Accretion History

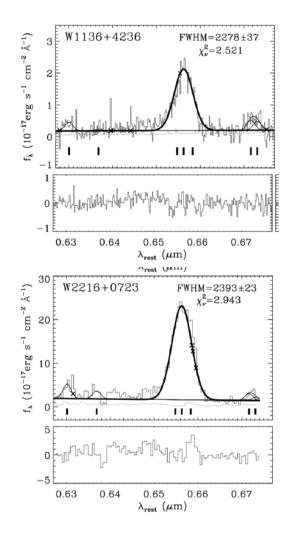


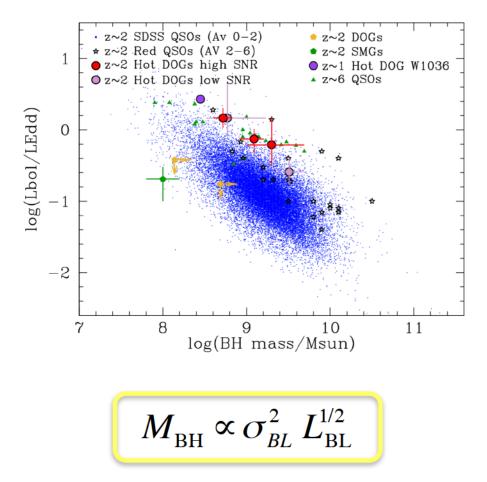
Averaged η < 0.15

- Black holes in
 ELIRGs are constantly at
 low spin!
- Chaotic accretion history

SMBH Mass and Eddington Ratios

Keck MOSFIRE and Gemini FLAMINGO2 spectra of H α

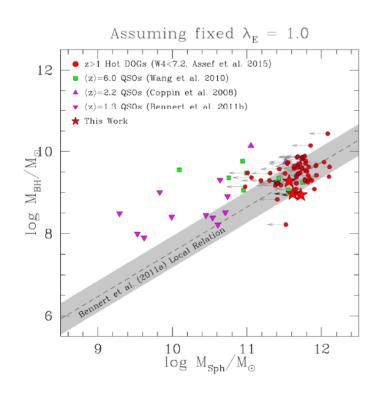


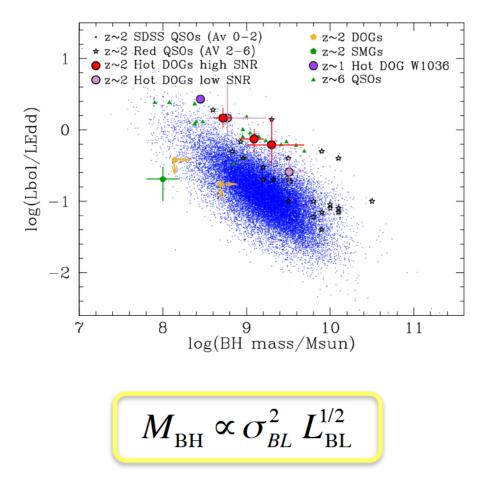


Wu et al. including Ned Wright (submitted)

SMBH Mass and Eddington Ratios

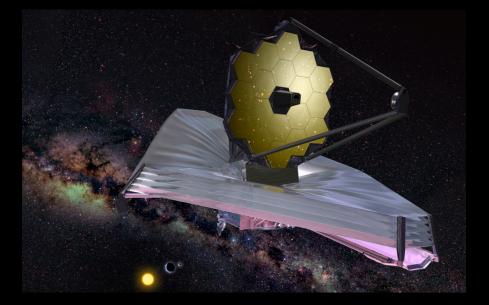
Keck MOSFIRE and Gemini FLAMINGO2 spectra of H α





Wu et al. including Ned Wright (submitted)

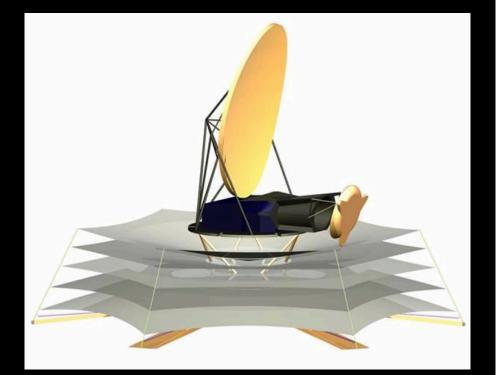
In the Future





Far-IR Surveyor

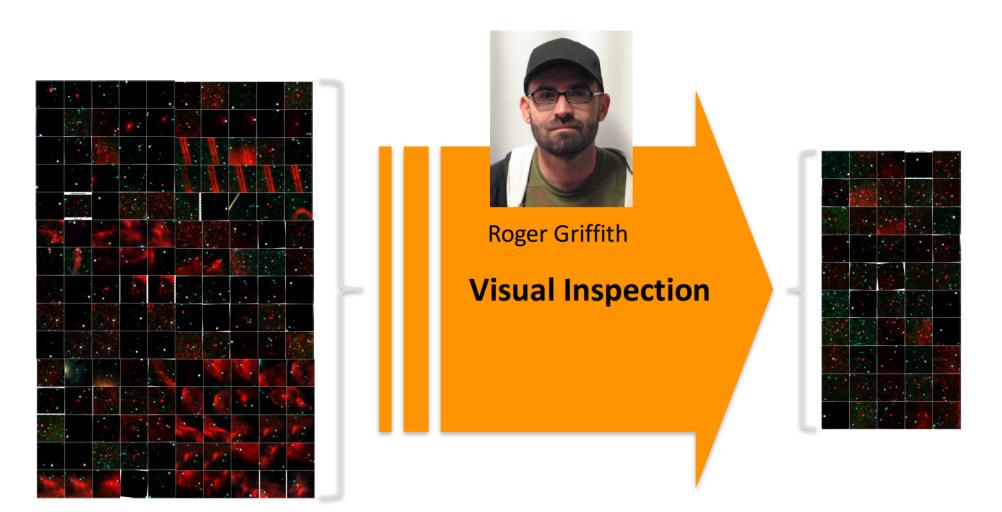
CALISTO concept for 2020 Decadal





What's Next?

Broadening the Selection Criteria



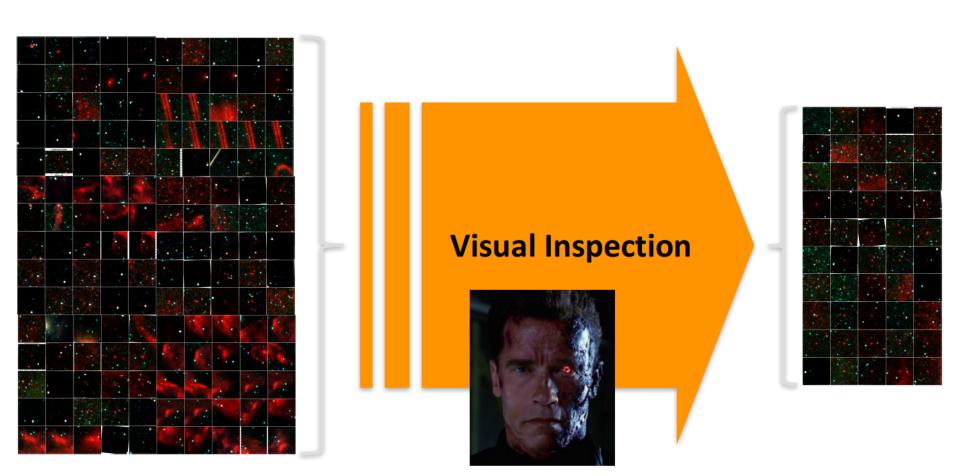
Archival Selection Results

Good ELIRG candidates



What's Next?

Broadening the Selection Criteria



Archival Selection Results

T-800

Good ELIRG candidates

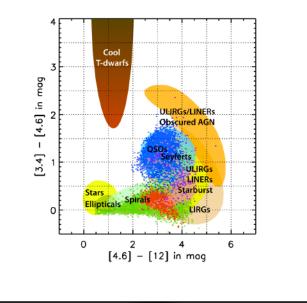
Because – What can go wrong?!

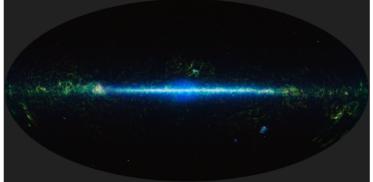


The Terminator, Orion Pictures (1984)



Inhabitants of WISE Color Space

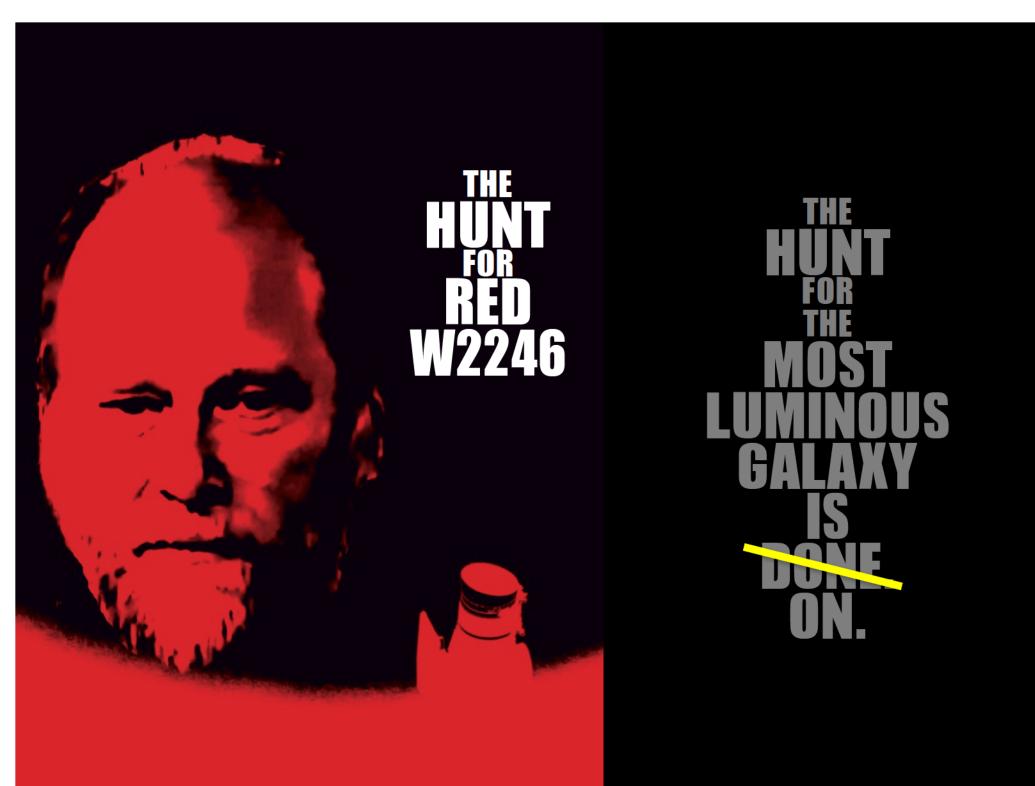






T-800





WISE One-Year-in-Orbit Party Peter Eisenhardt's House in Altadena, CA 2010/12/15





Salt Lick BBQ Driftwood, Texas (right outside Austin) 2012/01/12, AAS meeting



Ned's 60th Birthday Party

