

# How Did I End Up in the Galactic Center?

Frederick Baganoff

MIT Kavli Institute for Astrophysics  
and Space Research

Nedfest 2017 --- UCLA

# Ned @ 40

- My first research advisor
  - Quasar reddening
  - IRC +10 216
- I learned a lot about programming
- Ned's & Mark's 40<sup>th</sup> birthday party
  - Ned: "brevity is the soul of wit" – Hamlet
  - Francoise hiding Ned's rock 'n roll albums
- Playing tennis with Ned
  - Ned: "You don't *have* to hit the ball when you toss it."

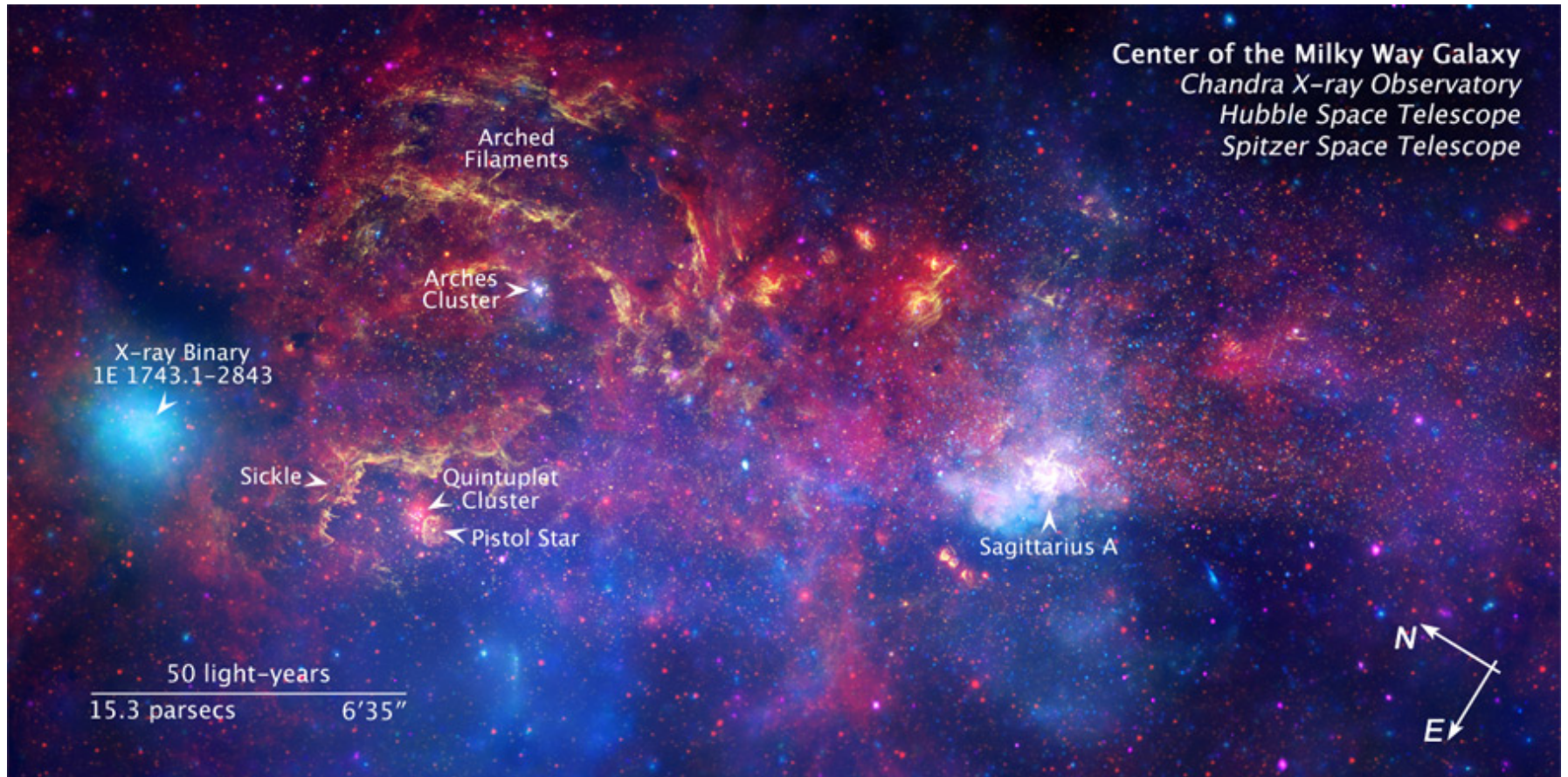
# Good Advice!



# Ned @ 40

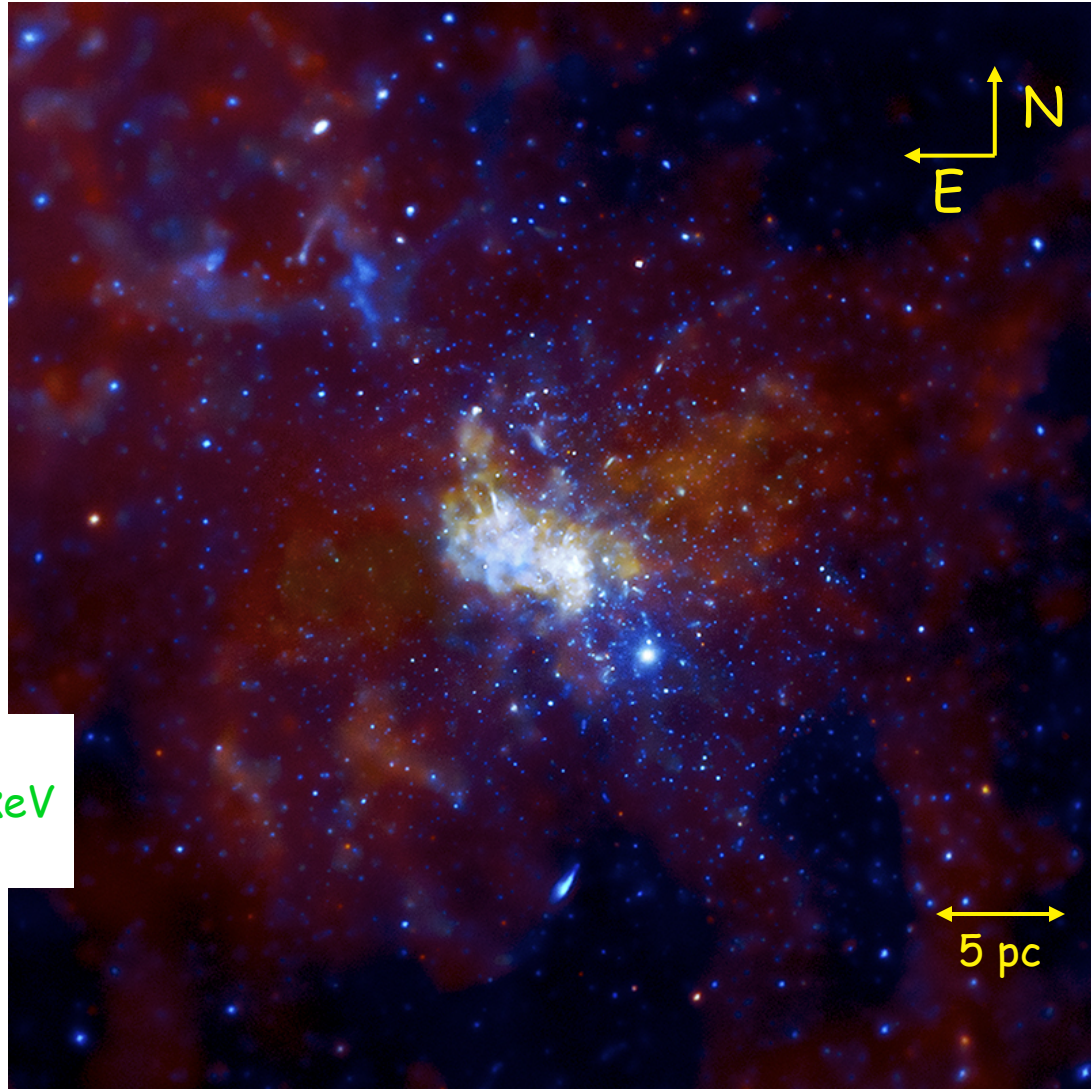
- Thesis project?
  - COBE launch 2+ years away
  - Ned: "All I have are programming projects. You don't want to do that."
- Off to work with Matt on AGN variability
- Which lead to Sgr A\* and the Galactic Center

# Ned @ 70



Chandra, Hubble & Spitzer Views of the Galactic Center

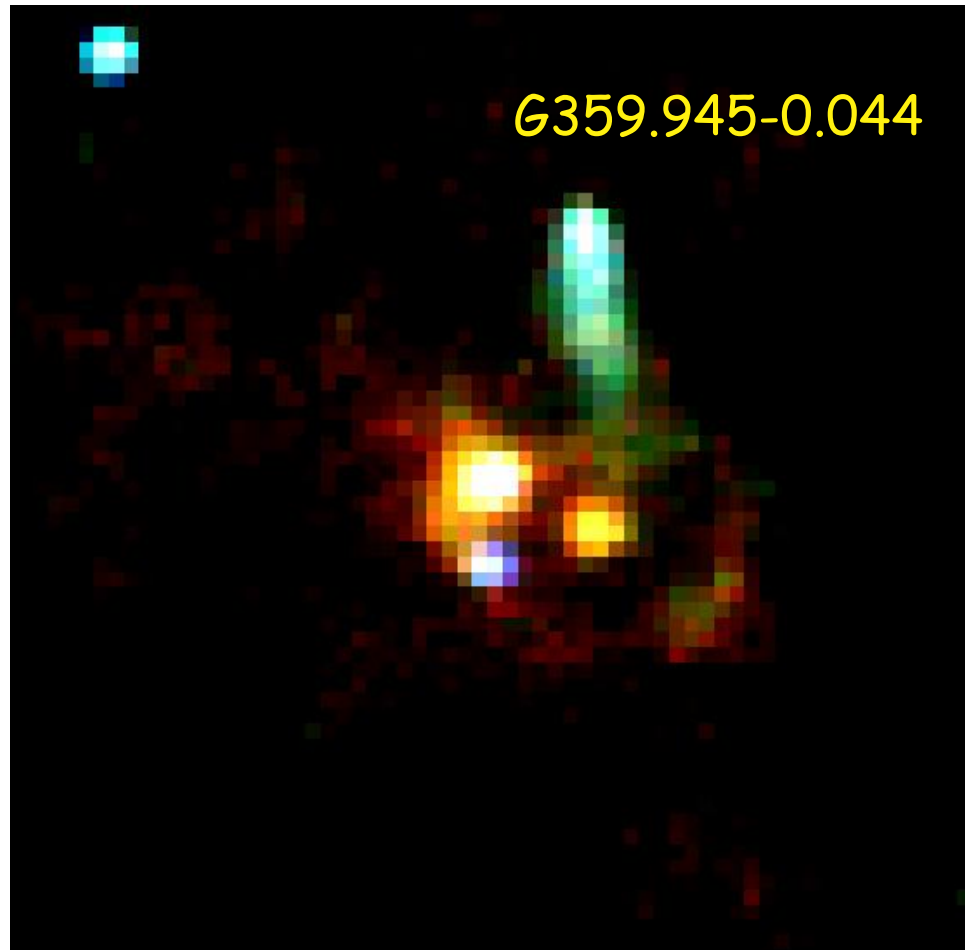
# Chandra Image of the Galactic Center



Red: 2-3.7 keV  
Green: 3.7-4.5 keV  
Blue: 4.5-8 keV

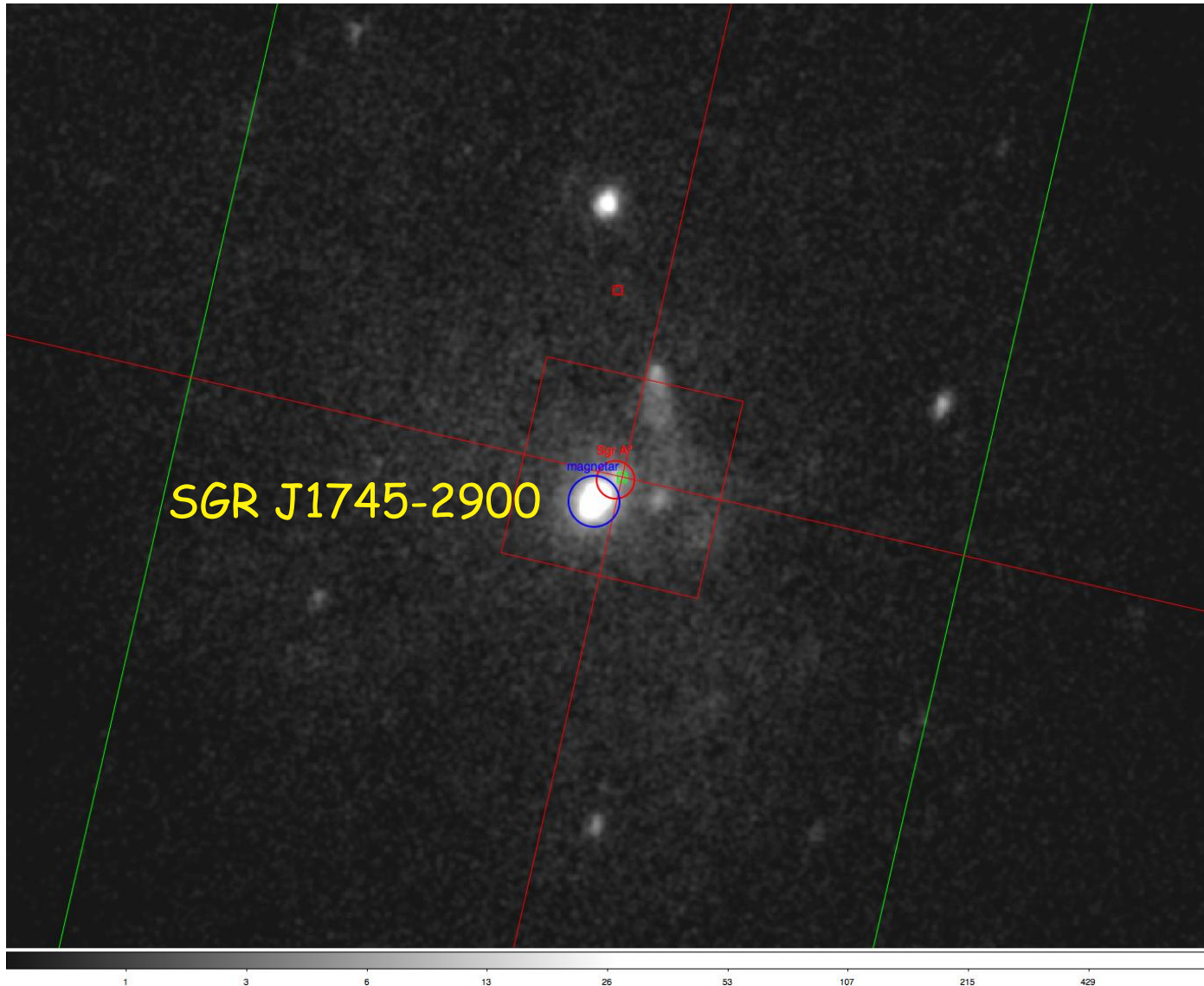
5 pc

# Central Parsec



2-3.5, 3.5-5, 5-8 keV

# GC Magnetar Outburst





# GC Magnetar SGR J1745-2900

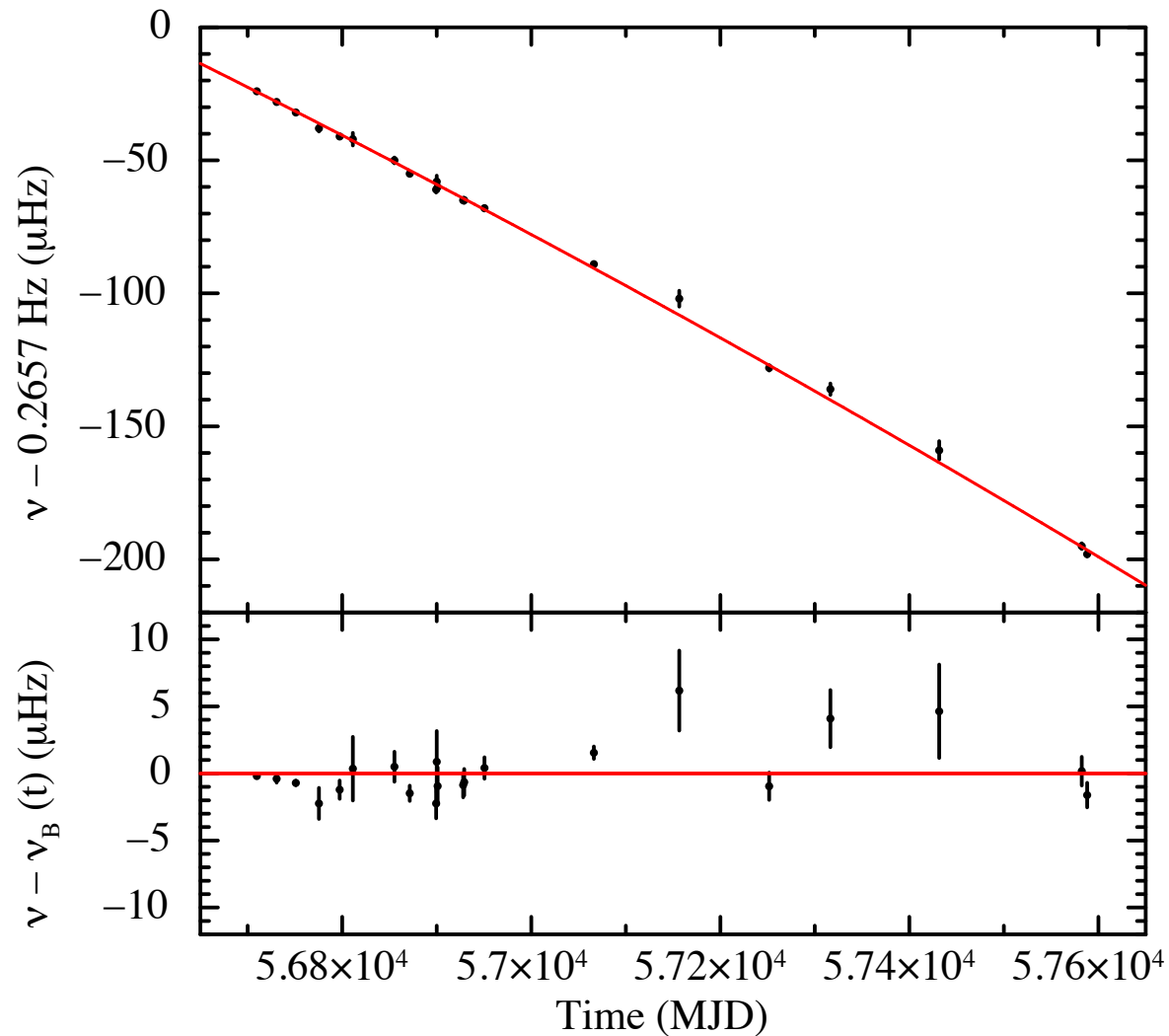
- Swift/BAT discovered outburst 2013 Apr 24 (Kennea et al. 2013)
- NuSTAR follow-up revealed 3.76 s period (Mori et al. 2013)
- Chandra imaging places it 2.4" SE of Sgr A\* (Baganoff et al. 2013)
- One of a handful of radio magnetars
- Eatough et al. (2013) infer super-equipartition  $B \sim 8$  mG near Bondi radius

# GC Magnetar SGR J1745-2900

- VLBA proper motion yields transverse velocity  $236 \pm 11$  km/s wrt Sgr A\* (Bower et al. 2015)
- Radial velocity unknown but chance alignment unlikely
- Velocity & position consistent with bound orbit originating within CW disk of massive stars orbiting Sgr A\*
- Orbital period 700+ yr

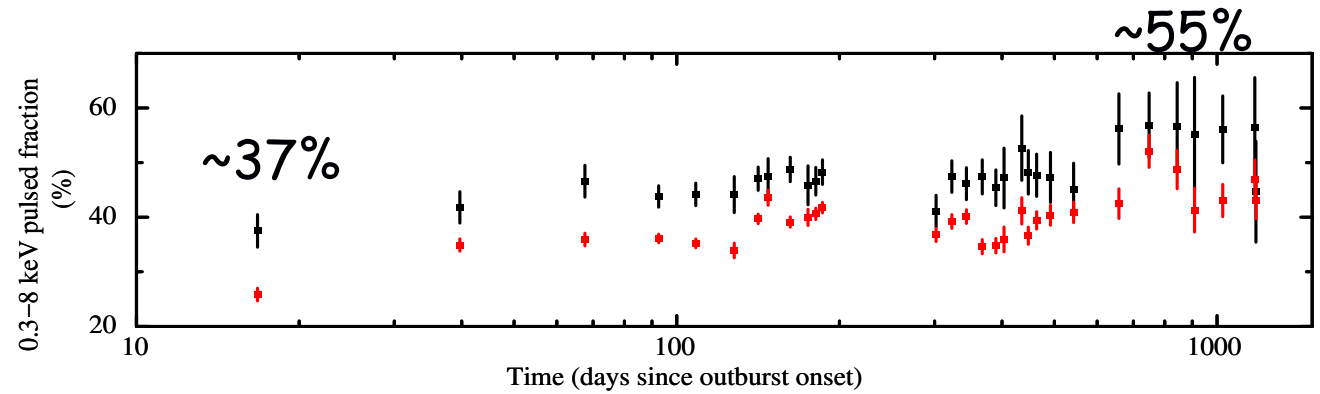
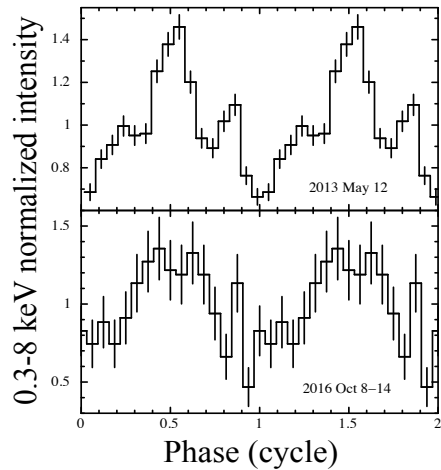
**Chandra monitoring of initial 3.5 yr of  
outburst decay analyzed by Coti Zelati  
et al. (2015, 2017)**

# Temporal Evolution of Spin Frequency

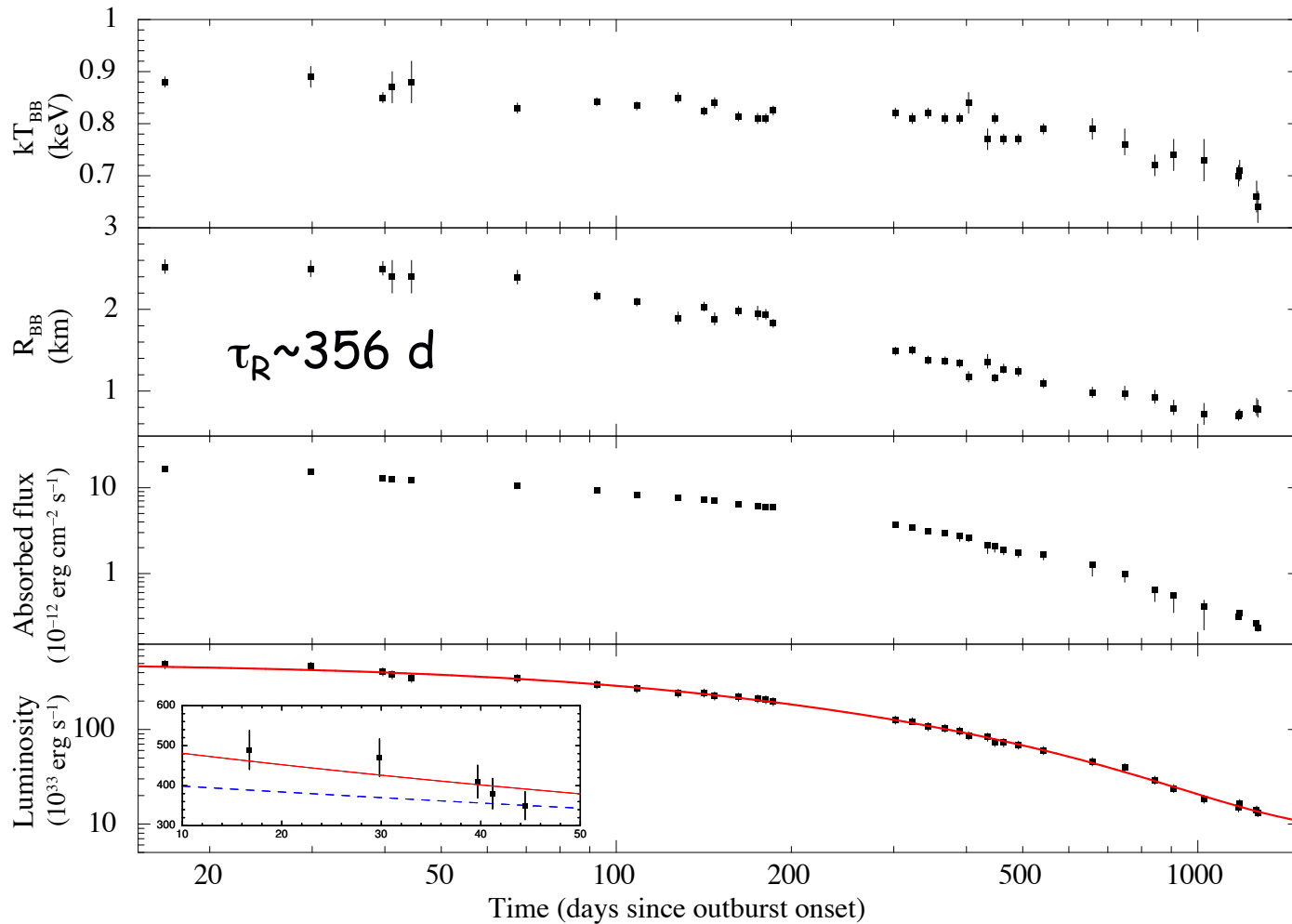


Approximately linear spin down

# Pulsed Fraction Slightly Rising

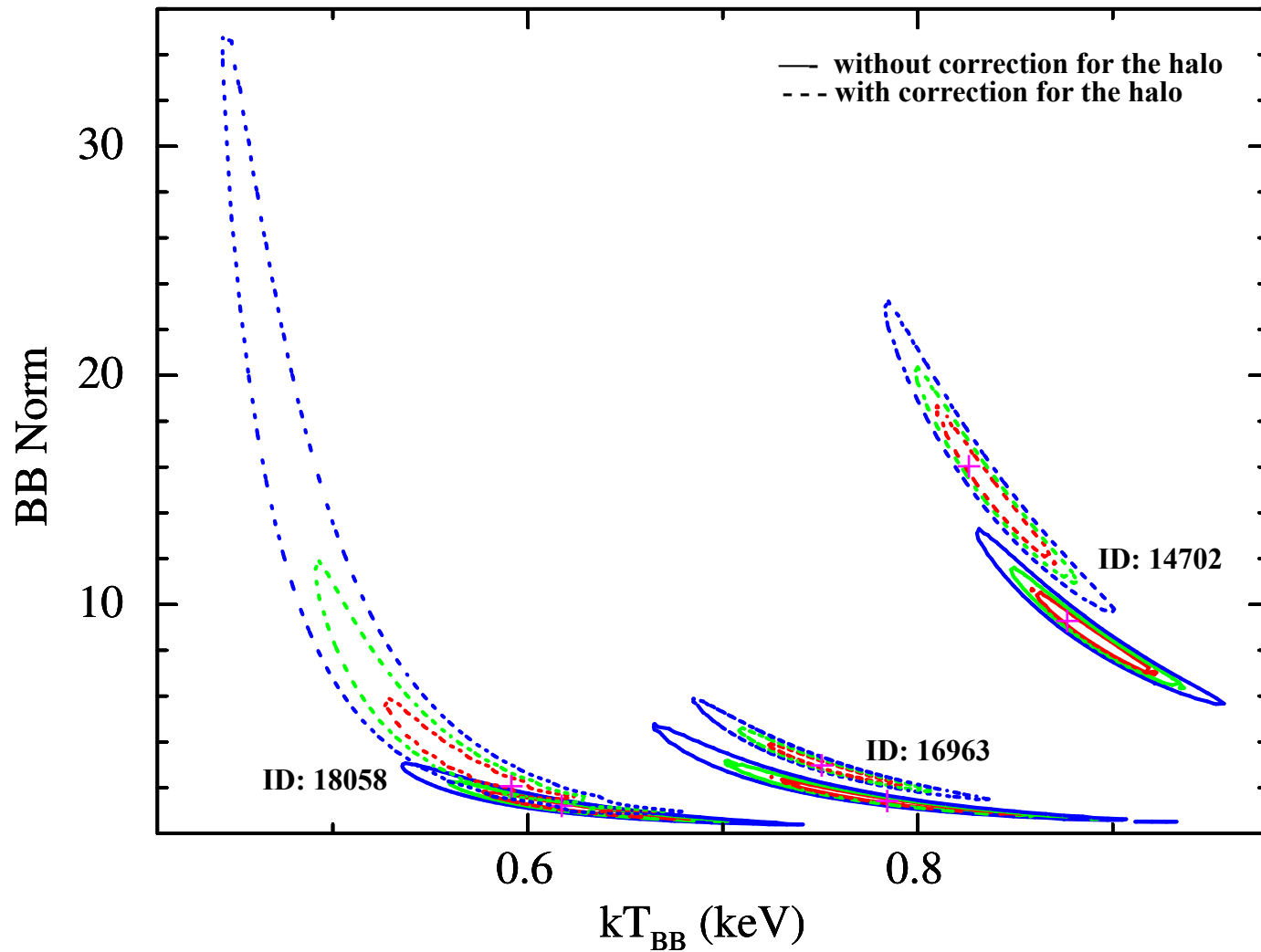


# Time Evolution of Blackbody T & R

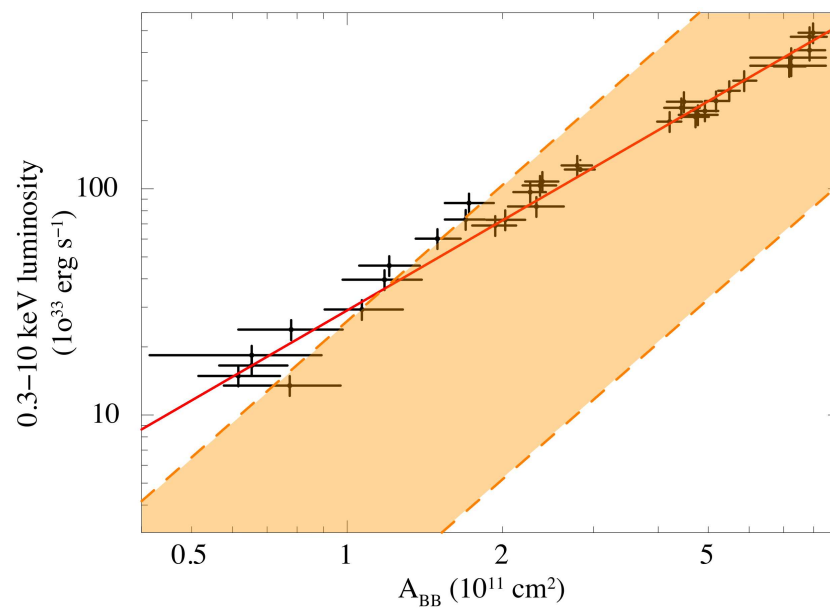
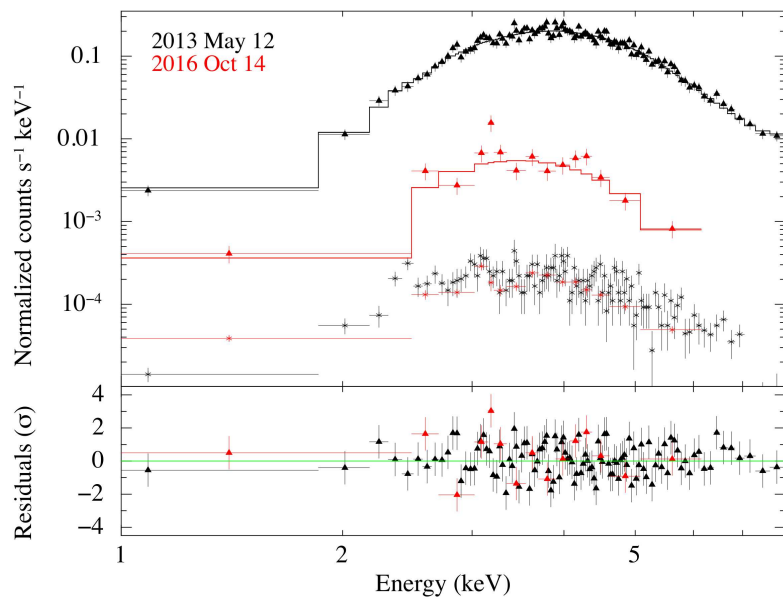


L has double exponential decay:  $\tau_1 \sim 96$  d,  $\tau_2 \sim 326$  d

# Correcting for Dust Scattering Halo



# Spectral Fits & Beloborodov Model





# Summary

- Crustal cooling model predicts hotspot should remain fairly constant in size while it cools as internal heat conducted up to outer crust; inconsistent with our observations
- Alternative model of untwisting bundle of current-carrying closed field lines in magnetosphere works better (Beloborodov 2009)
- Other models being tested

# Summary

- Magnetar was still visible to Chandra in 2017 July
- Day-long exposures measured spin
- Two more long Chandra observations approved for 2018 July => follow spin evolution for 5 yr

**Lots more to do in the GC**

**Thanks, Ned!**

**Happy Birthday!**

**Ned @ 100?**