The LZ Experiment Tom Shutt SLAC

SURF South Dakota



(10 tons LXe total)

Center for Underground Physics (Korea)

LIP Coimbra (Portugal) MEPhI (Russia) Edinburgh University (UK)

University of Liverpool (UK)

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Imperial College London (UK)

University College London (UK) University of Oxford (UK) STFC Rutherford Appleton, and Daresbury, Laboratories (UK) University of Sheffield (UK)

LUX - ZEPLIN

31 Institutions, ~200 people

University of Alabama University at Albany SUNY Berkeley Lab (LBNL), UC Berkeley **Brookhaven National Laboratory Brown University** University of California, Davis Fermi National Accelerator Laboratory Lawrence Livermore National Laboratory University of Maryland Northwestern University University of Rochester University of California, Berkeley University of California, Santa Barbara University of South Dakota South Dakota School of Mines & Technology South Dakota Science and Technology Authority SLAC National Accelerator Laboratory Texas A&M Washington University

University of Wisconsin Yale University



Principle of Operation

- 3D imaging rejects external backgrounds
- Electron-recoil backgrounds distinguished by ratio of charge / light ratio
- High purity LXe target
- Single photon and electron sensitivity





The LZ Detector



Low background Ti vessels

HV UMBILICAL AND CONNECTION TO CATHODE







Performance drivers

- Backgrounds
- Purity for charge drift
- Light Collection
- Drift field, low electron + photon emission
 - -Discrimination
 - -Threshold
 - -Grids: surface fields vs light collection
- Extensive test program: small chambers, HV in LAr, and System Test - T. Biesiadzinksi's talk

	Requirement / Baseline	Goal
Cathode HV	50 kV	100 kV
Light collection	7.5%	12%
e ⁻ lifetime (µs)	850	2800
N-fold trigger coincidence	3	2
²²² Rn	20 mBq	1 mBq

Prototype TPC Section









Discrimination



- LUX new standard for discrimination calibration
- Discrimination strongest
 at lowest energy
- LZ requirement: ≥99.5%



Signal production in liquid Xe



Electron Recoils Low field, low energy

Signal production in liquid Xe



Figure: Gibson/Shutt

Electron Recoils High field, high energy

Signal production in liquid Xe



Nuclear Recoils

Absolute calibration: the "Doke" plot



Comprehensive framework captured in NEST MC package



Calibrations

- Expand upon successful LUX program
- Spatial response, temporal variation

 -^{83m}Kr, ^{131m}Xe
- Outer LXe and Gd-scintillator

-²²⁰Rn, movable gamma ray sources

- Electron and
 - Nuclear recoils
 - -Tritium
 - Variety of high and low energy neutron sources





Outer Detector System

- LXe skin scintillation: 4-8 cm (walls), ~20 cm (dome)
- Gd-loaded liquid scintillator (LAB): 60 cm, 21.5 tons.
- ~97% efficient for neutrons
- Hermetic measurement of all penetrating backgrounds





Backgrounds



- Significant screening effort: K. Oliver-Mallory talk
- Assessment of backgrounds: M.E. Monzani's talk
- Internal backgrounds dominate: Kr, Rn. Goal: *Neutrinos dominate!* (and are interesting signal).

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- Rn (and Kr) dominant internal radioactive background
- Emanates from most materials
- 20 mBq requirement, 1 mBq goal
- Four separate measurements systems, ~0.1 mB sensitivity
 - -Experience from SNO, KamLAND, EXO, NEMO, Borexino
- "Scrubber" processes purge
 Xe from warm breakout regions





Xe Purification and Cryogenics

- Kr removal via chromatography
- Gas phase purification through getter 10 tons / 2.5 days
- Trap-enhanced mass spec: ~ppt
- High efficiency two-phase heat exchange
- LN thermosyphon-based cryogenics multiple cooling locations. Stirling LN refrigerator.
- Most aspects tested in System Test
- 10 tons Xe in hand or under contract for 2018 delivery





Kr removal



DAQ, Electronics, Control, Offline



Slow control in use at System Test

Schedule

• CD1: March 2015

-Conceptual Design Report arXiv:1509.02910

- CD2: April 2016
- LUX removed Feb 2017
- Underground installation beings May 2018
- Operations begin May 2019

Projected Sensitivity

Summary

- LXe is pre-eminent target for high mass WIMPs
- LZ leverages LUX innovations in calibrations, cryogenics, purification
- Robust test program optimizing detector performance
- Very high fiducial volume fraction due to outer detector and low background cryostat
- Goal: neutrino-limited sensitivity of ~1x10⁻⁴⁸ cm²
- Goal: significant neutrino physics, including strong solar ⁸B v-nuclear coherent scattering signal

Spin Dependent Sensitivity

Nuclear Recoil Spectrum

Double Beta decay Sensitivity

Solar Axions

