

# Particle Energization and the Tearing-Driven Turbulent Cascade

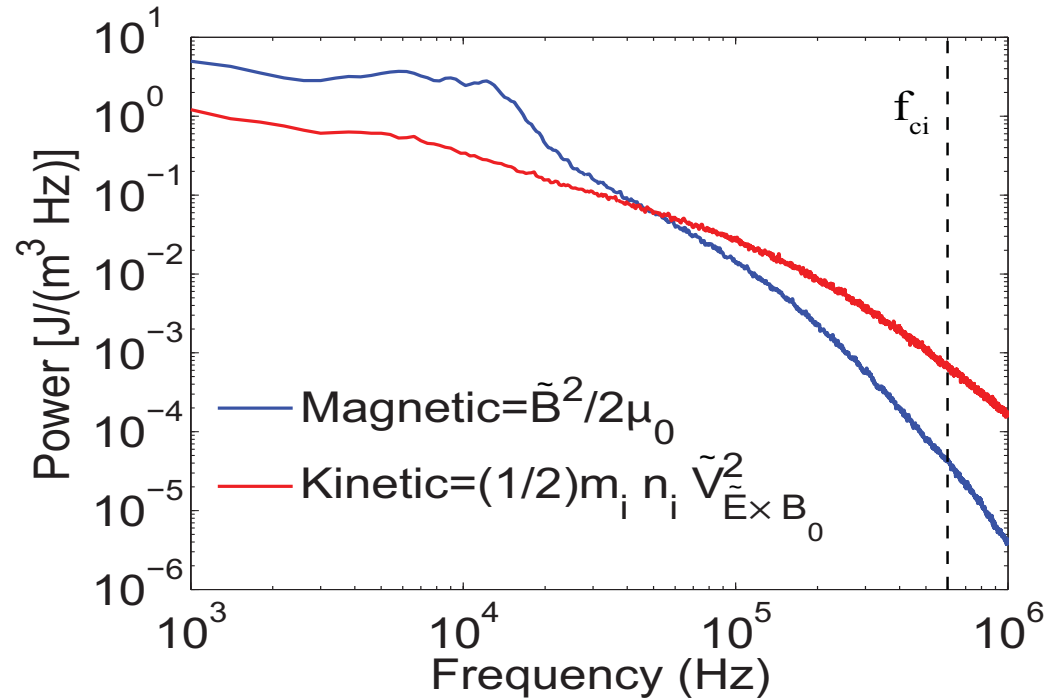
John Sarff

A. Almagri, J. Anderson, A. DuBois, D. Craig\*, D. Den Hartog,  
C. Forest, K. McCollam, M. Nornberg, P. Terry, D. Thuecks  
and the MST Team  
(\*Wheaton College)



**WISCONSIN**  
UNIVERSITY OF WISCONSIN-MADISON

# The tearing-driven turbulent cascade and dissipation

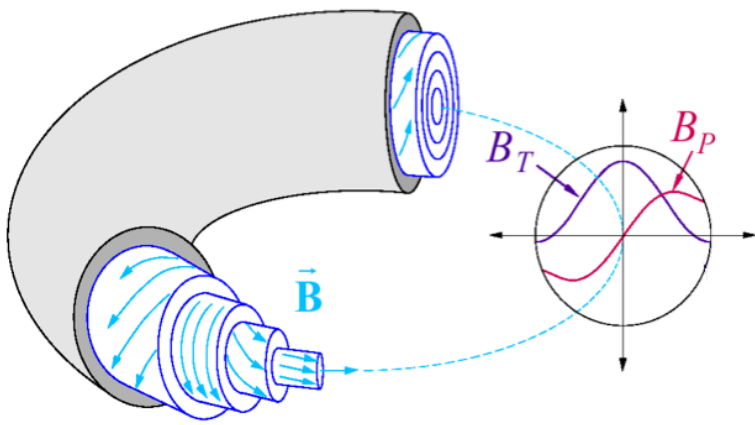
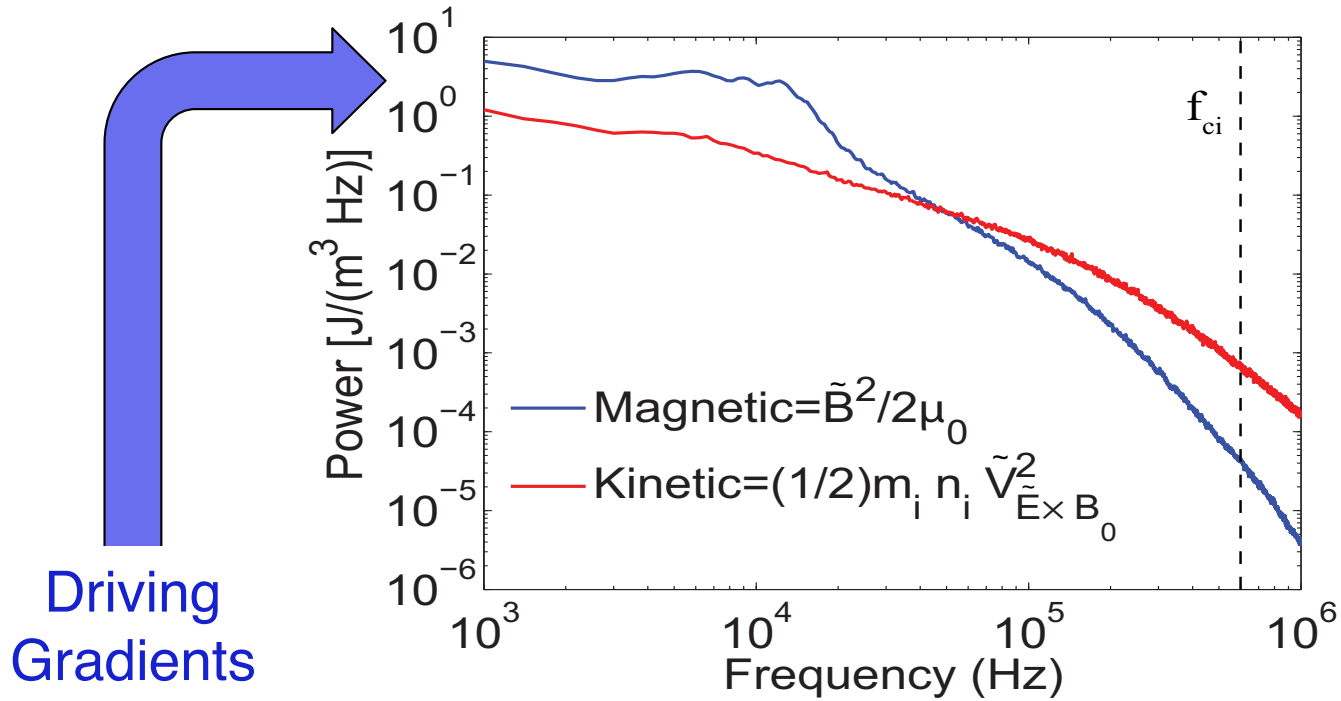


- **Connected challenges:**

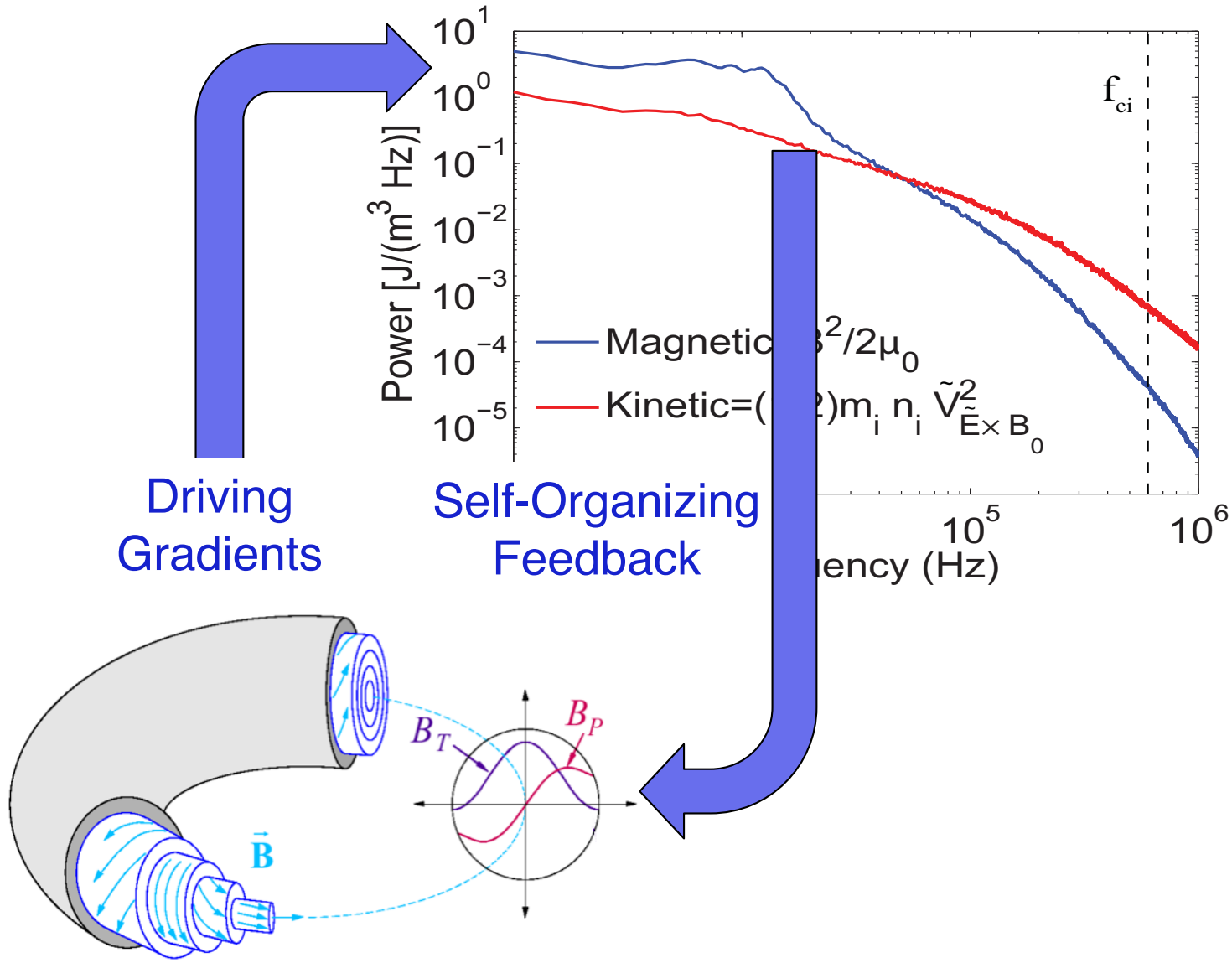
- Multi-scale modeling framework that spans global instability to turbulent dissipation
- Diagnostics and methods that isolate and distinguish particle energization processes



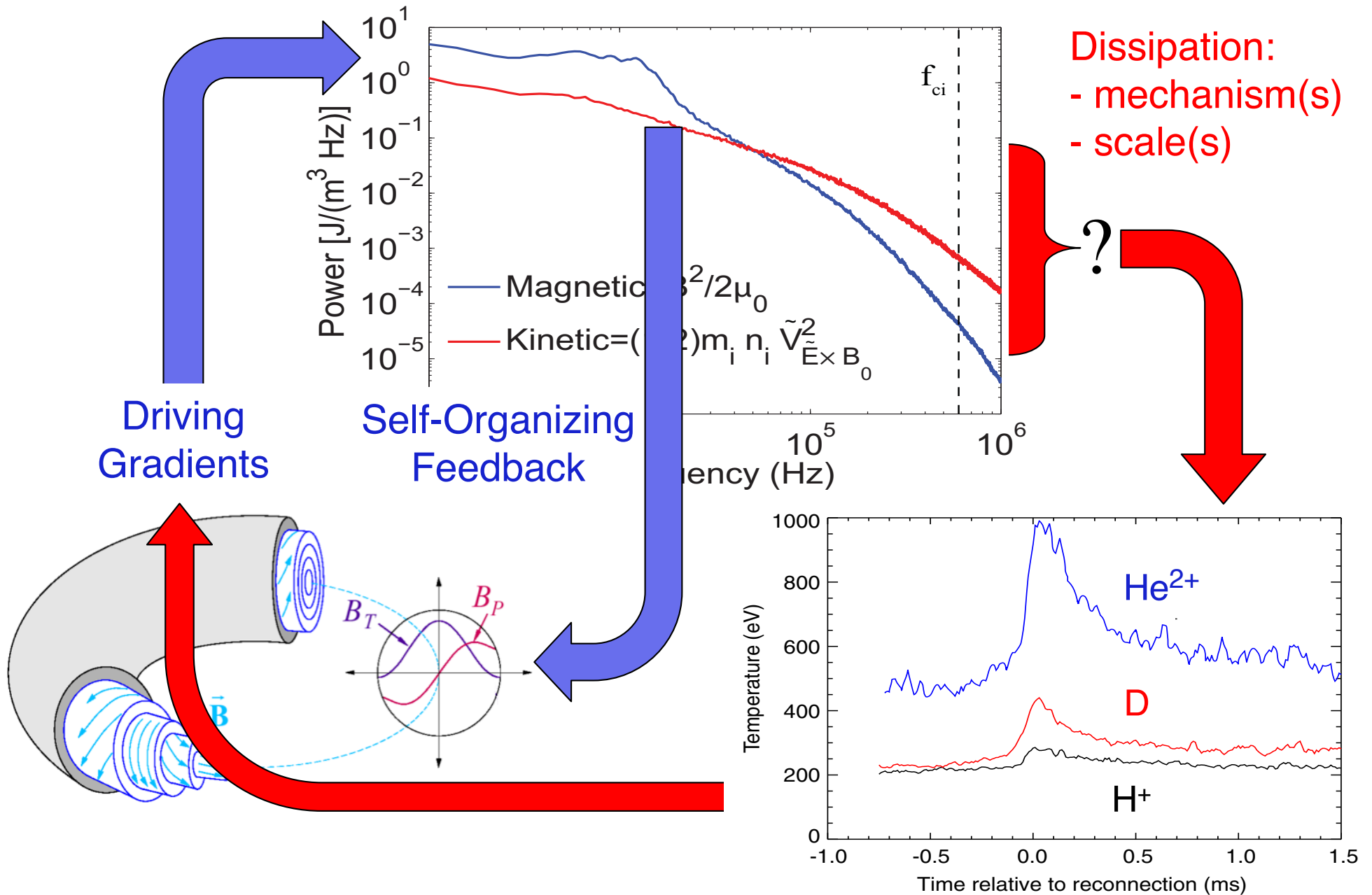
# The tearing-driven turbulent cascade and dissipation



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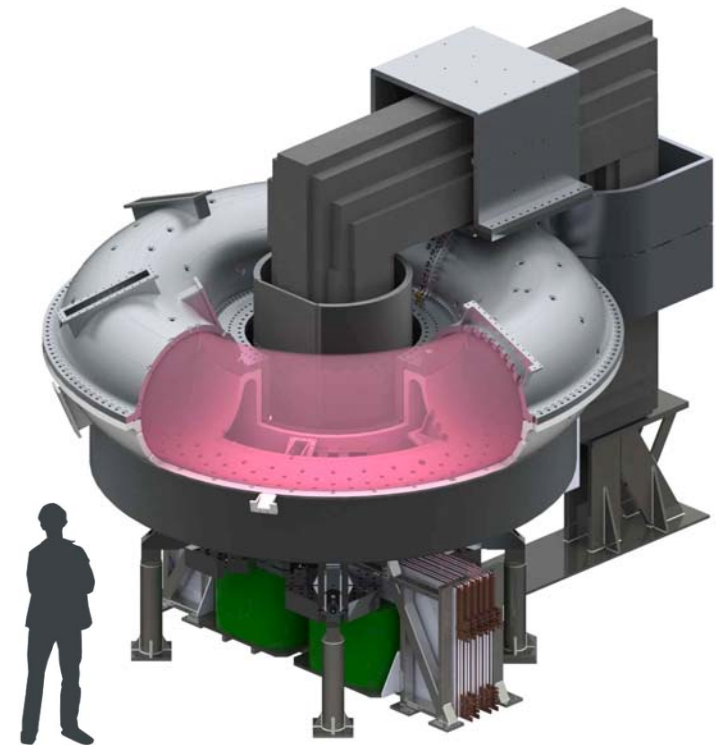
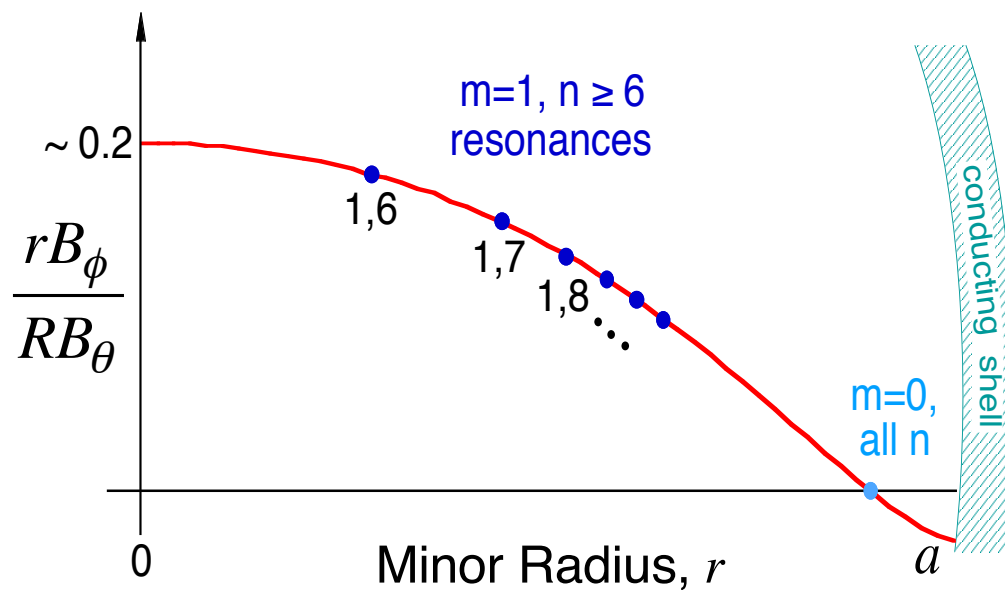
# The tearing-driven turbulent cascade and dissipation



# MST reversed field pinch plasmas



- Magnetic induction is used to drive a large current in the plasma
  - Plasma current,  $I_p = 50\text{-}600$  kA ;  $B < 0.5$  T
  - Inductive ohmic heating: 5-10 MW (input to electrons)
  - $T_i \sim T_e < 2$  keV, despite weak  $i$ - $e$  collisional coupling ( $n \sim 10^{19}$  m $^{-3}$ )
  - Minor radius,  $a = 50$  cm ; ion gyroradius,  $\rho_i \approx 1$  cm ;  $c/\omega_{pi} \approx 10$  cm
  - $\beta < 25\%$  ; Lundquist number  $S = 5 \times 10^{4-6}$

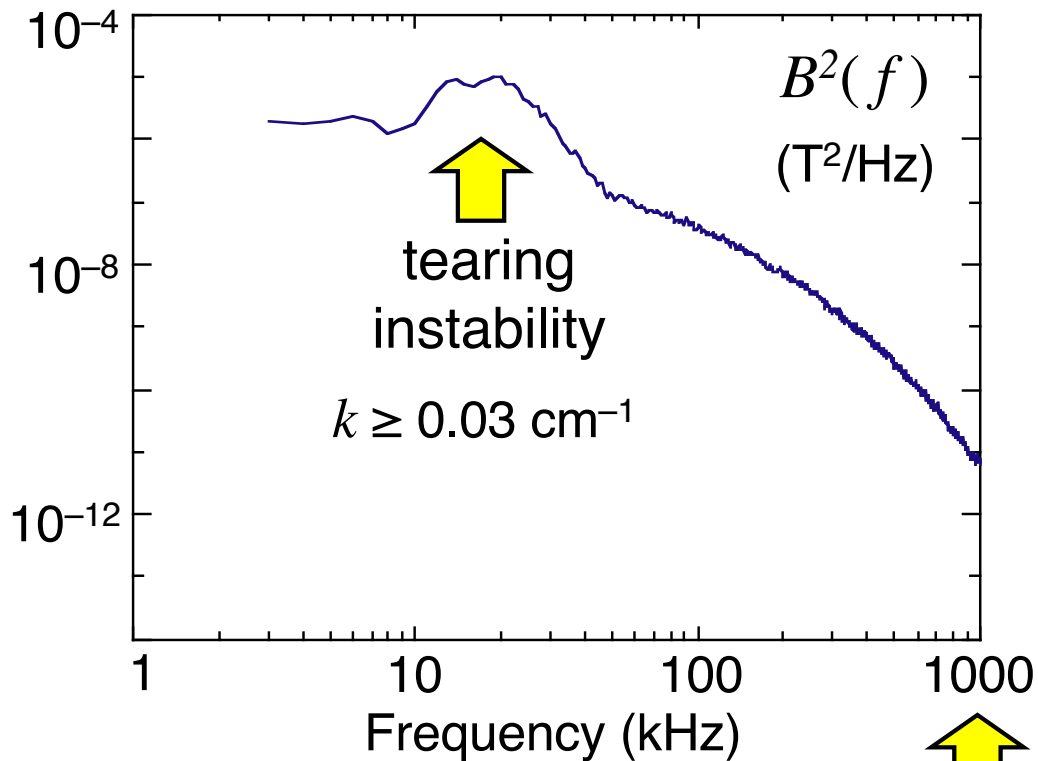


# Tearing instability drives a broadband turbulent cascade

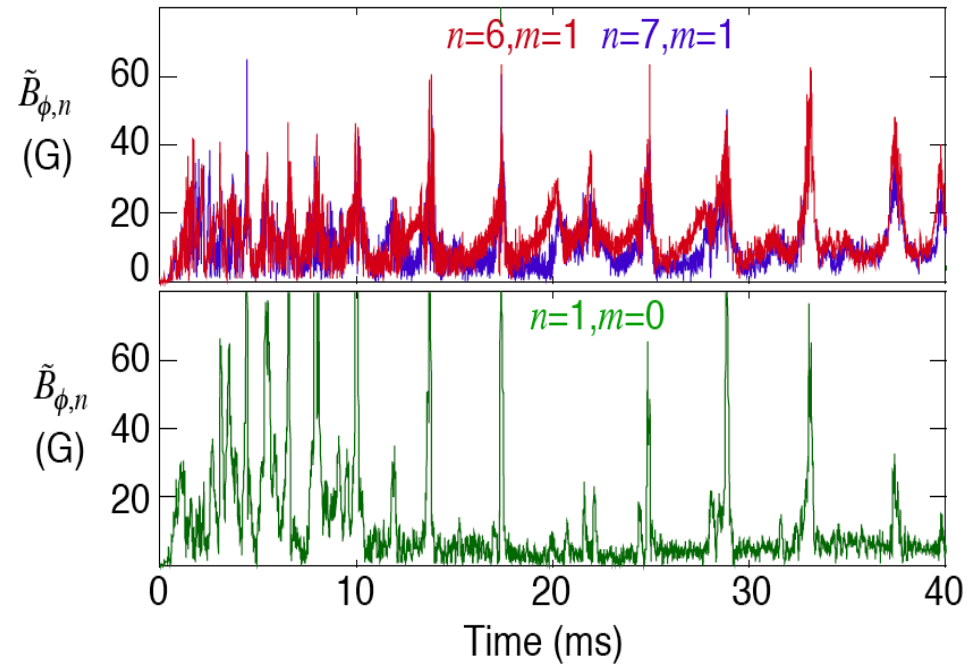


Toroidicity allows distinct  $k_{\parallel} = 0$  resonant modes at many radii in the plasma:

$$0 = \mathbf{k} \cdot \mathbf{B} = \frac{m}{r} B_{\theta} + \frac{n}{R} B_{\phi} \quad \begin{array}{l} m = \text{poloidal mode number} \\ n = \text{toroidal mode number} \end{array}$$



$\omega_{ci} / 2\pi$   
 $\rho_i \sim 1 \text{ cm}$

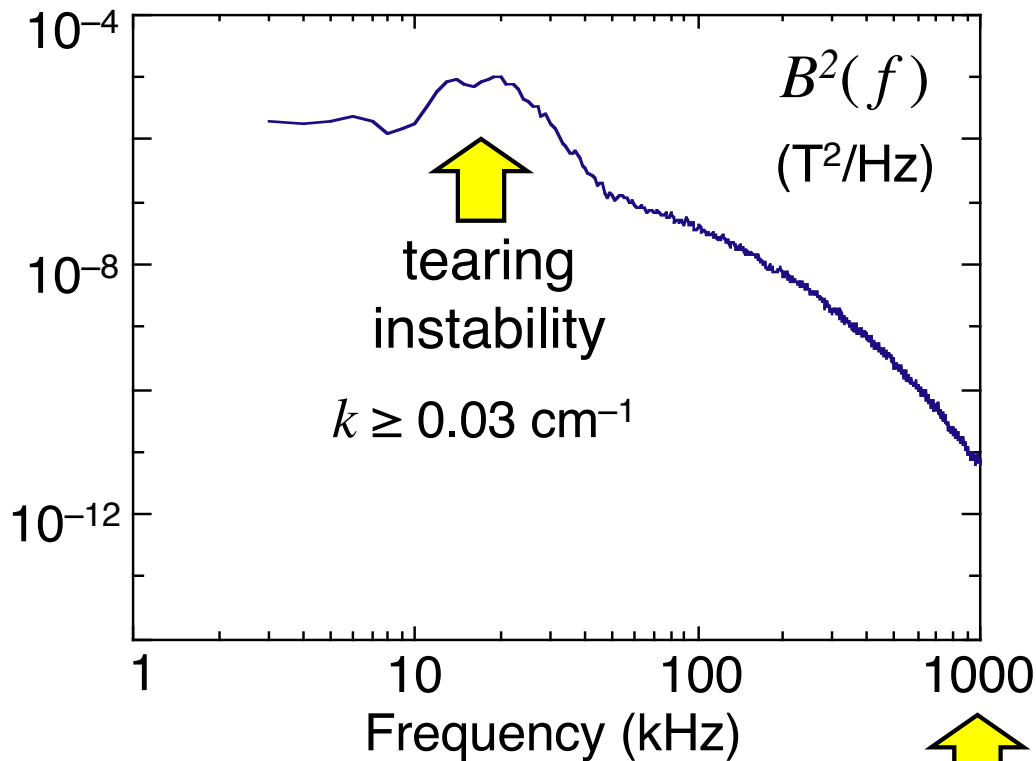


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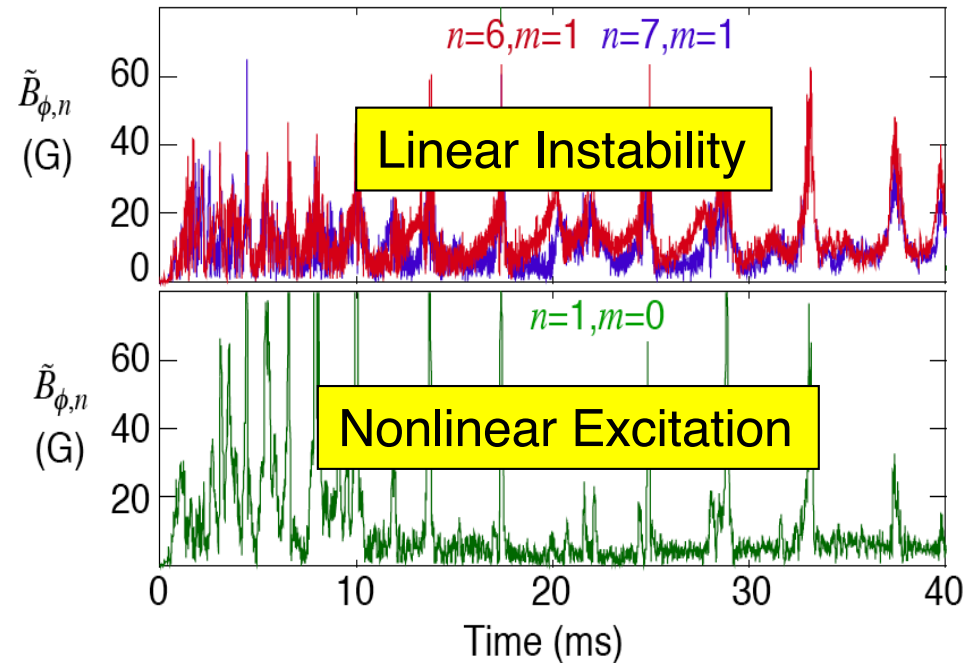


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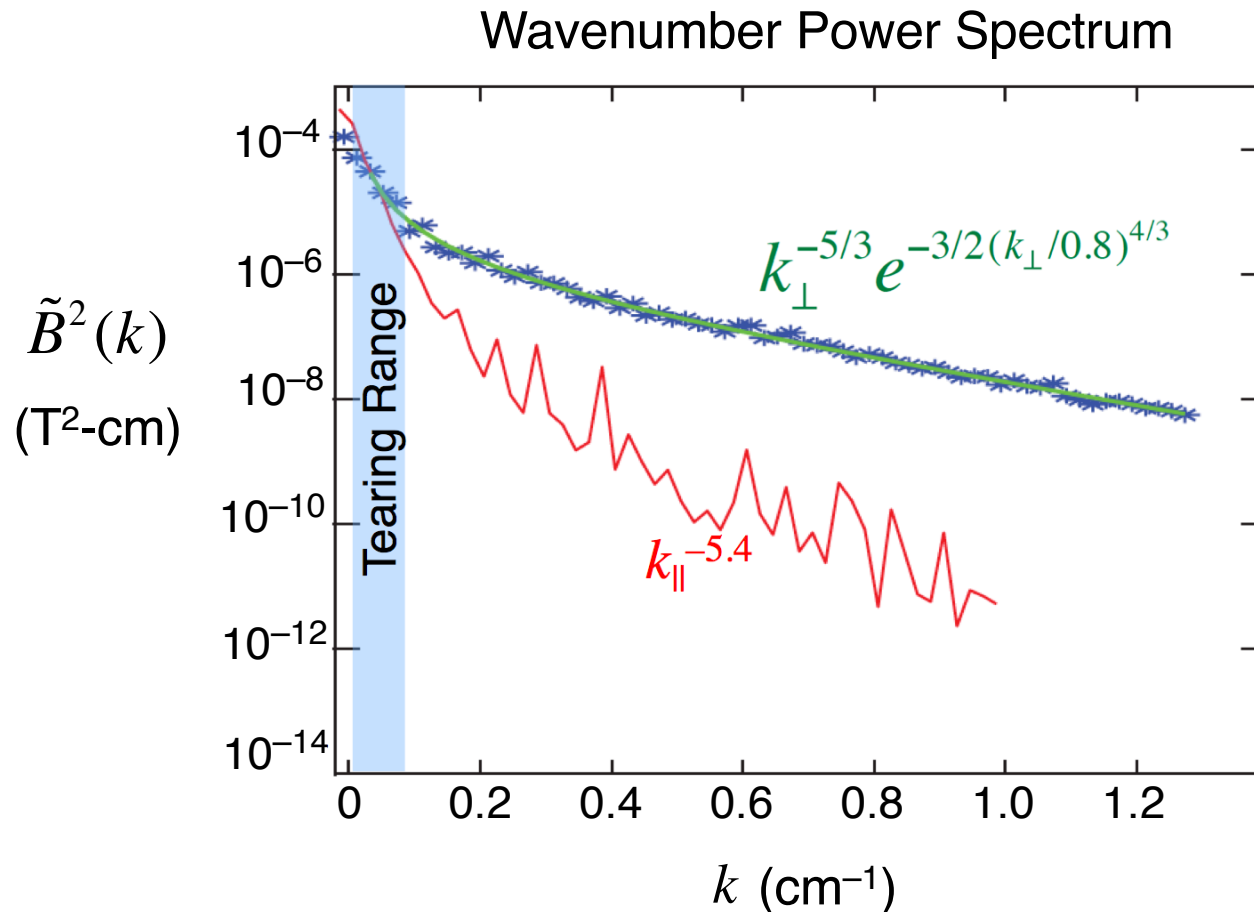




# The cascade is anisotropic and hints at a non-classical dissipation mechanism



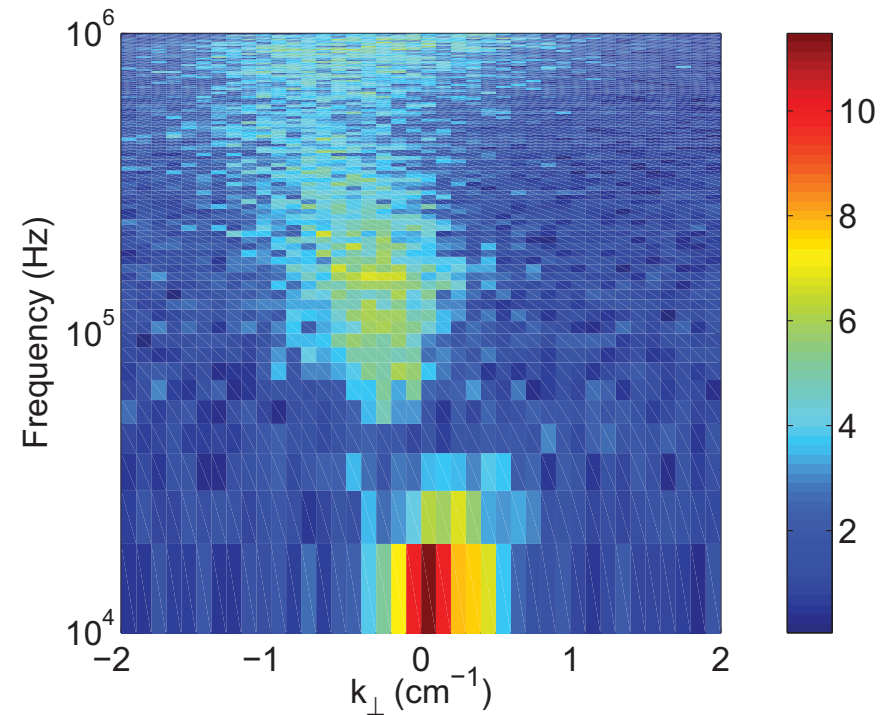
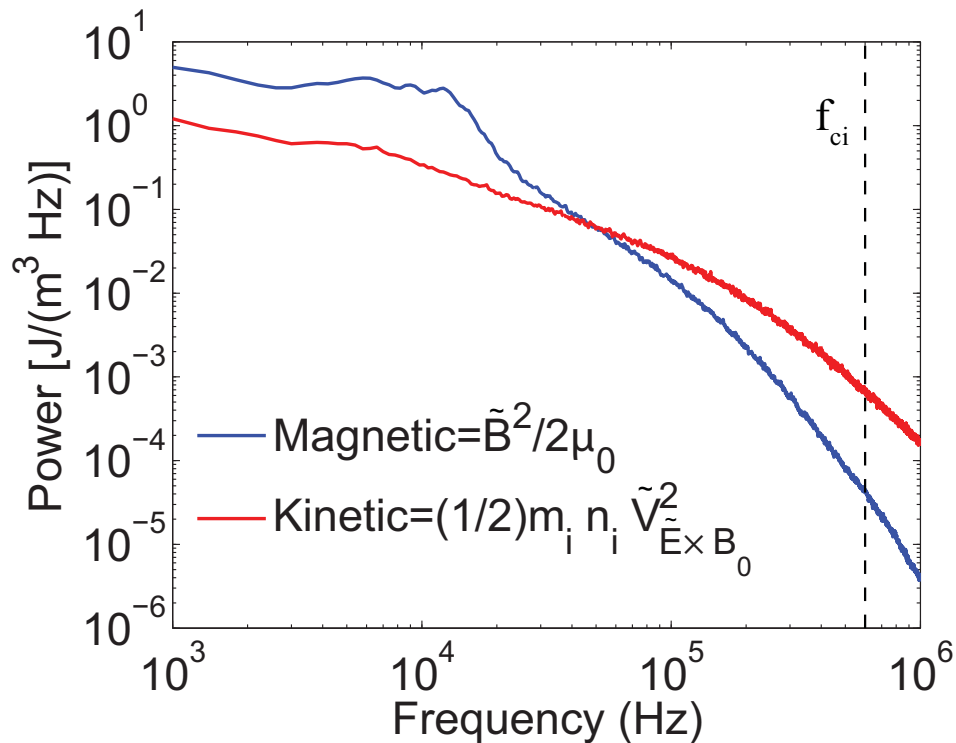
- The  $k_{\perp}$  spectrum fits a dissipative cascade model (Y. Ren, PRL 2011; P. Terry, PoP 2009)
- Onset of exponential decay (dissipation) occurs at a smaller  $k_{\perp}$  than expected for classical dissipation



# New evidence for the onset of drift waves at intermediate scales



- Turbulence becomes kinetic energy dominant at  $k_{\perp} \rho_s \approx 0.2$
- Signatures consistent with drift waves (Thuecks et al, PoP 2017)

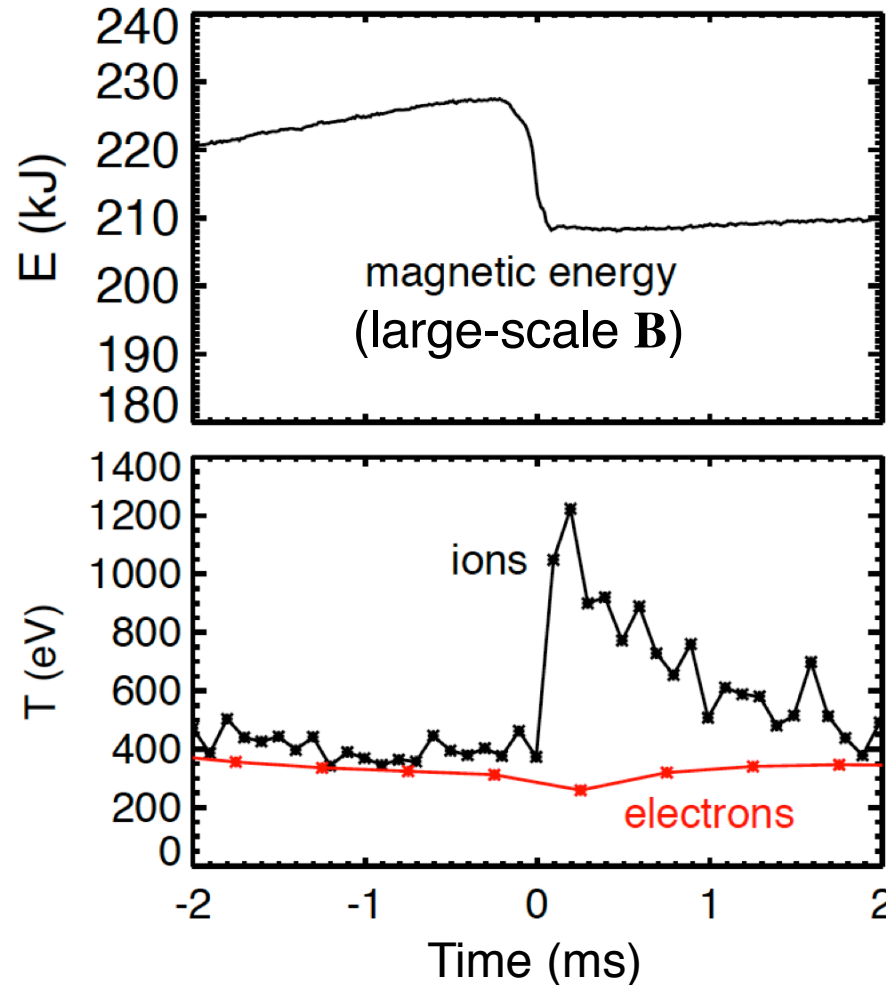


# Non-collisional Ion Energization

# Powerful ion energization is associated with impulsive magnetic reconnection events



- Instantaneous heating rate up to 10 MeV/s (50 MW)



$$T_i > T_e$$

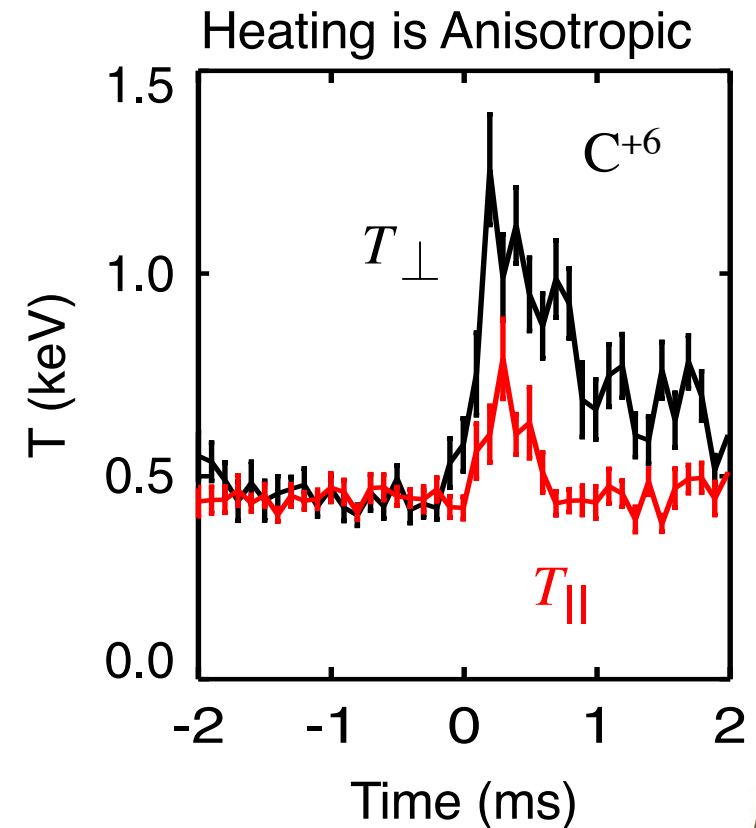
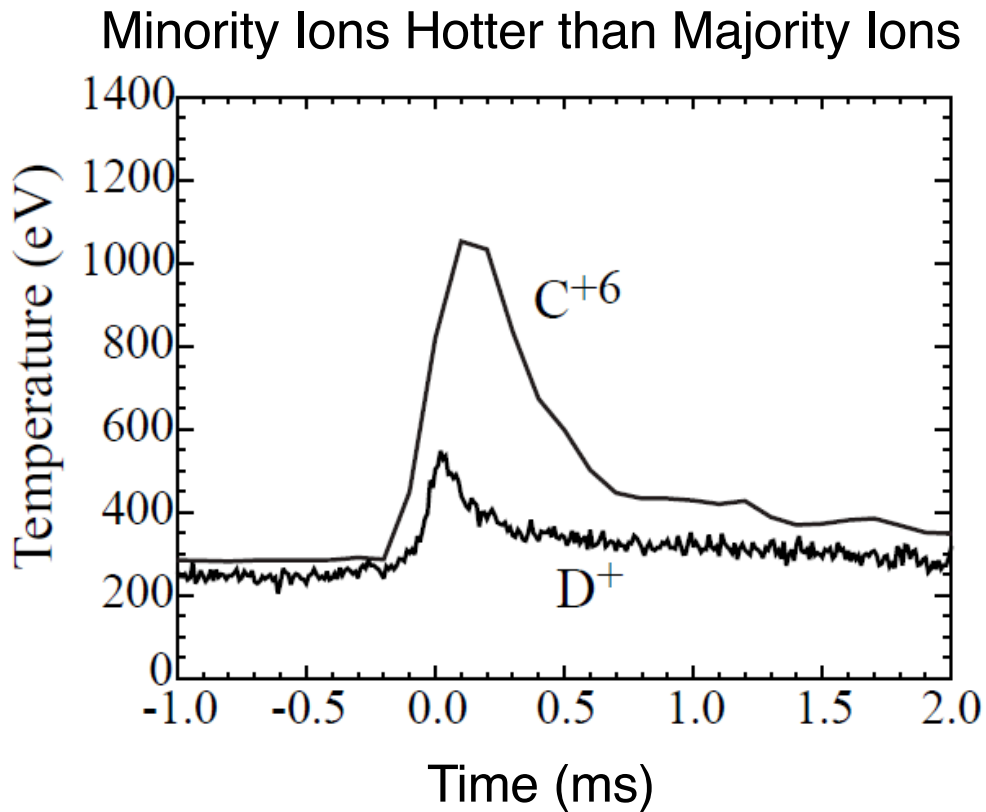
Relative to Reconnection Event



# Heating is anisotropic and species dependent



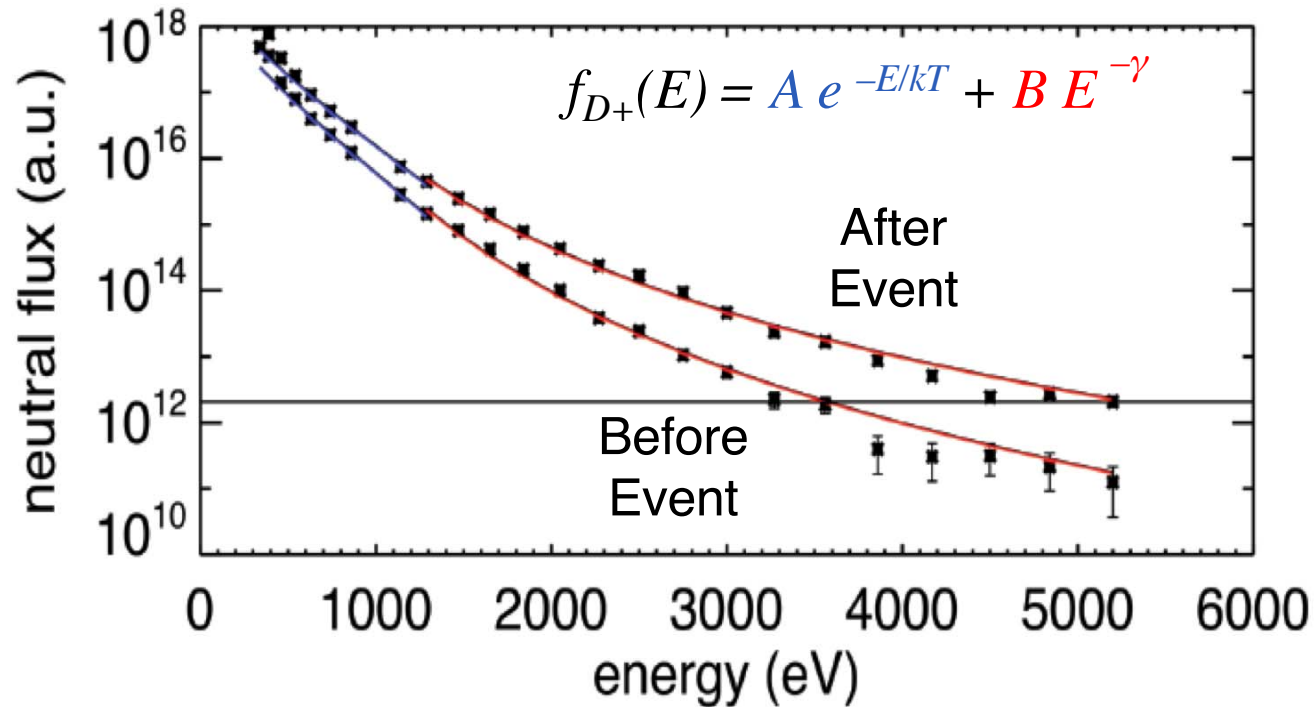
- Ion distribution diagnostics on MST:
  - Rutherford scattering for majority ion temperature
  - Charge-exchange recombination spectroscopy (CHERS) for minority ions
  - Neutral particle energy analyzers



# An energetic ion tail is generated and reinforced at each reconnection event



- Distribution is well-fit by a Maxwellian plus a power-law tail



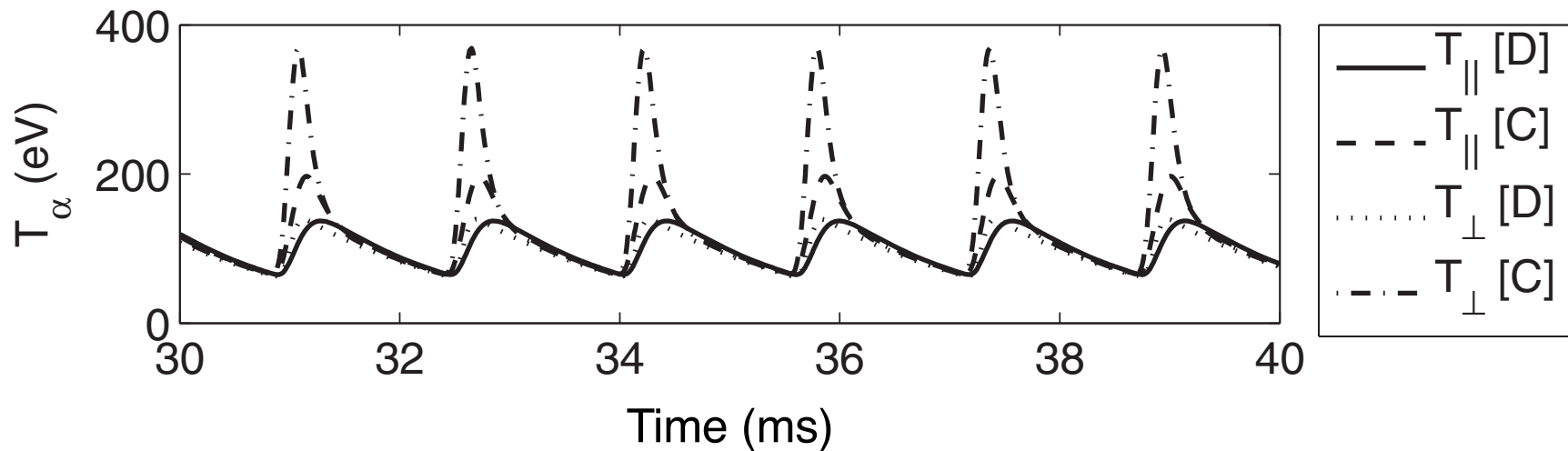
# Proposed Ion Heating Mechanisms

# The models proposed for ion heating in the RFP are similar to those for the solar corona and wind



- **Cyclotron-resonant heating:**

- Feeds off the turbulent cascade to gyro-scale
- Preferential perpendicular heating, but with collisional relaxation
- Preferential minority ion heating, since  $\tilde{B}^2(\omega_{ci})$  is larger where  $\omega_{ci}$  is smaller
- Mass scaling is predicted with dominant minority heating and collisional relaxation



Tangri et al., PoP **15** (2008)  
(similar to Cranmer et al)



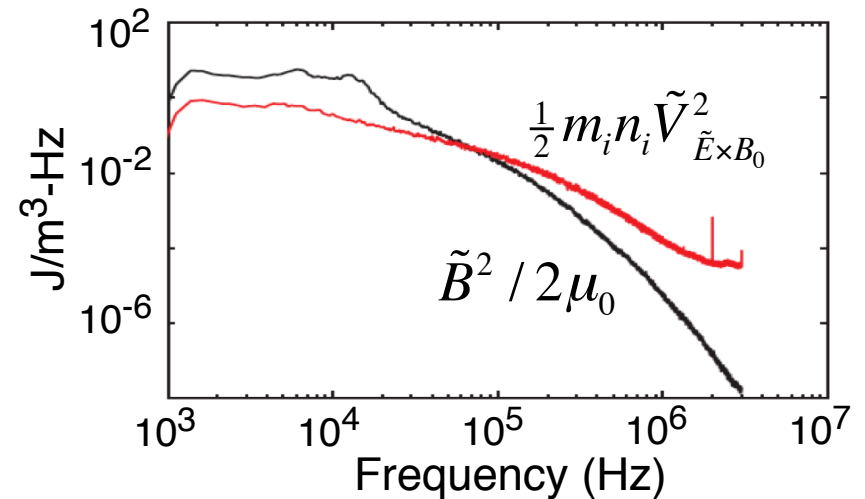
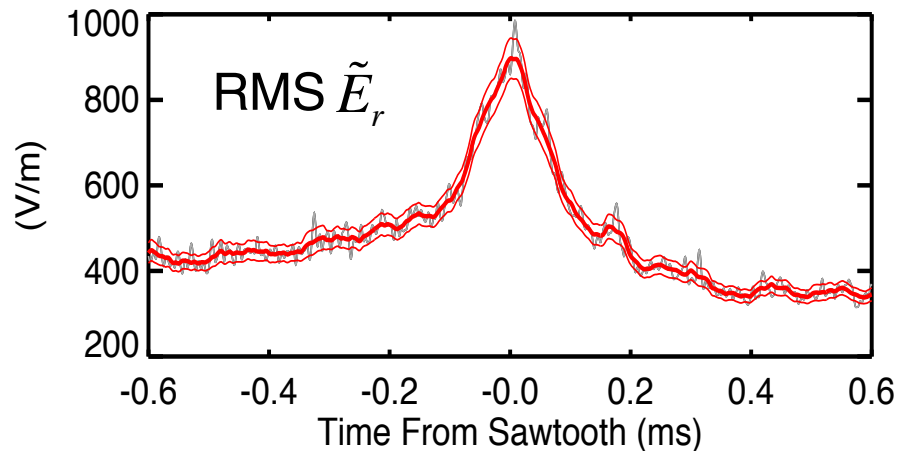


# The models proposed for ion heating in the RFP are similar to those for the solar corona and wind



- **Stochastic heating:**

- Feeds off large electric field fluctuations and a distinct chaotic diffusion process
- Monte Carlo modeling yields MST-like heating rates (Fiksel et al, PRL 2009)
- Predicts mass scaling close to that observed



Does non-Alfvénic cascade help make heating more powerful?



# Viscous heating is not sufficient



- No experimental evidence for the required large sheared flow
- Perpendicular flow is dominant for tearing modes for which the classical viscosity is small
- Difficult to achieve the large impulsive heating rates seen in MST plasmas
- See, e.g., Svidzinski et al, PoP **15** (2009)



# The need for multi-scale modeling



- Fluid treatment of tearing instability and the self-organizing feedback is mature but does not encompass small scales
  - Nonlinear MHD and two-fluid models (NIMROD, DEBS)
  - Typical spatial resolution  $k\rho_i \lesssim 0.3$
  - Braginskii dissipation, sometimes artificially enhanced for numerical purposes
- A large-scale electric field (dynamo flux) is created in the self-organizing feedback
  - Energization on the largest scale, e.g., energetic ions
  - Two-fluid nature allows ions and electrons to respond differently
- Boundary conditions are important
  - Unavoidable consequence of being a confined plasma
  - Coupling to other modes like drift waves, even if they are stable
  - Boundary interfaces are generic, e.g., corona or magnetosphere



# (Some) next step RFP experimental opportunities



- Diagnostics and methods that discriminate heating mechanisms
  - Limited measurements suggest ion heating is preferentially perpendicular
  - Test bed for wave-particle correlation study
  - Diagnostic challenge is significant, so understanding probably hinges on modeling predictions for turbulence characteristics
- Plasma control improvements, e.g., current magnitude and duration
  - *Lower* the Lundquist number to isolate MHD regime (versus two-fluid)
  - Increase access for both intrusive and nonintrusive diagnostics
- Inject plasmoids
  - Mix with tearing-driven turbulence
  - Increase beta
  - Form shocks



# Computational model for tearing-relaxation recently extended to include two-fluid effects



- Nonlinear multi-mode evolution solved using NIMROD
- Motivated by measurements that suggest coupled electron and ion relaxation

Ohm's law: 
$$\mathbf{E} = -\mathbf{V} \times \mathbf{B} + \frac{1}{ne} \mathbf{J} \times \mathbf{B} - \frac{1}{ne} \nabla p_e + \eta \mathbf{J} + \frac{m_e}{ne^2} \frac{\partial \mathbf{J}}{\partial t}$$

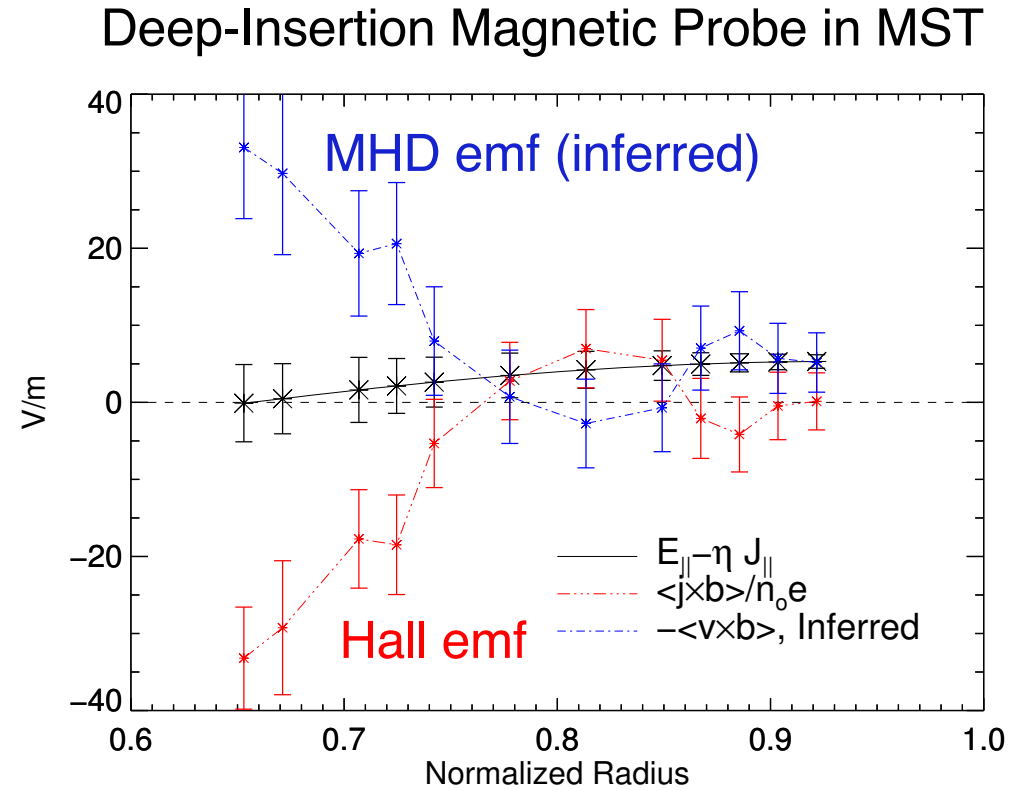
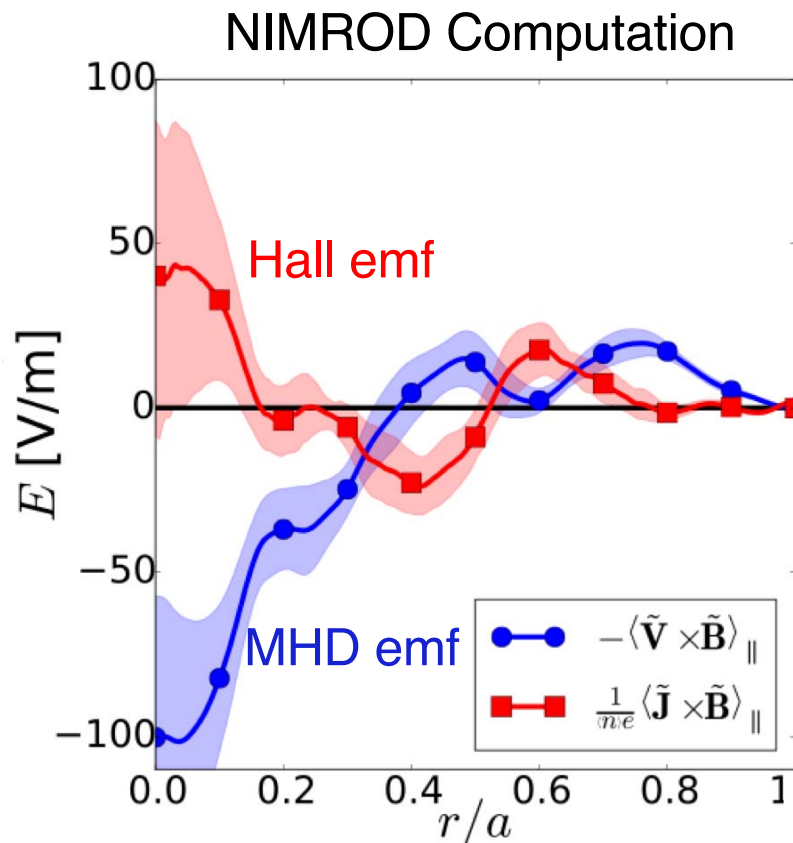
Momentum: 
$$nm_i \frac{d\mathbf{V}}{dt} = \mathbf{J} \times \mathbf{B} - \nabla p - \nabla \cdot \Pi_{gyro} - \nabla \cdot \mathbf{v} nm_i \mathbf{W}$$

A blue double-headed arrow points from the  $\mathbf{J} \times \mathbf{B}$  term in the momentum equation to the  $\frac{1}{ne} \mathbf{J} \times \mathbf{B}$  term in the Ohm's law equation, indicating a coupling between the two.

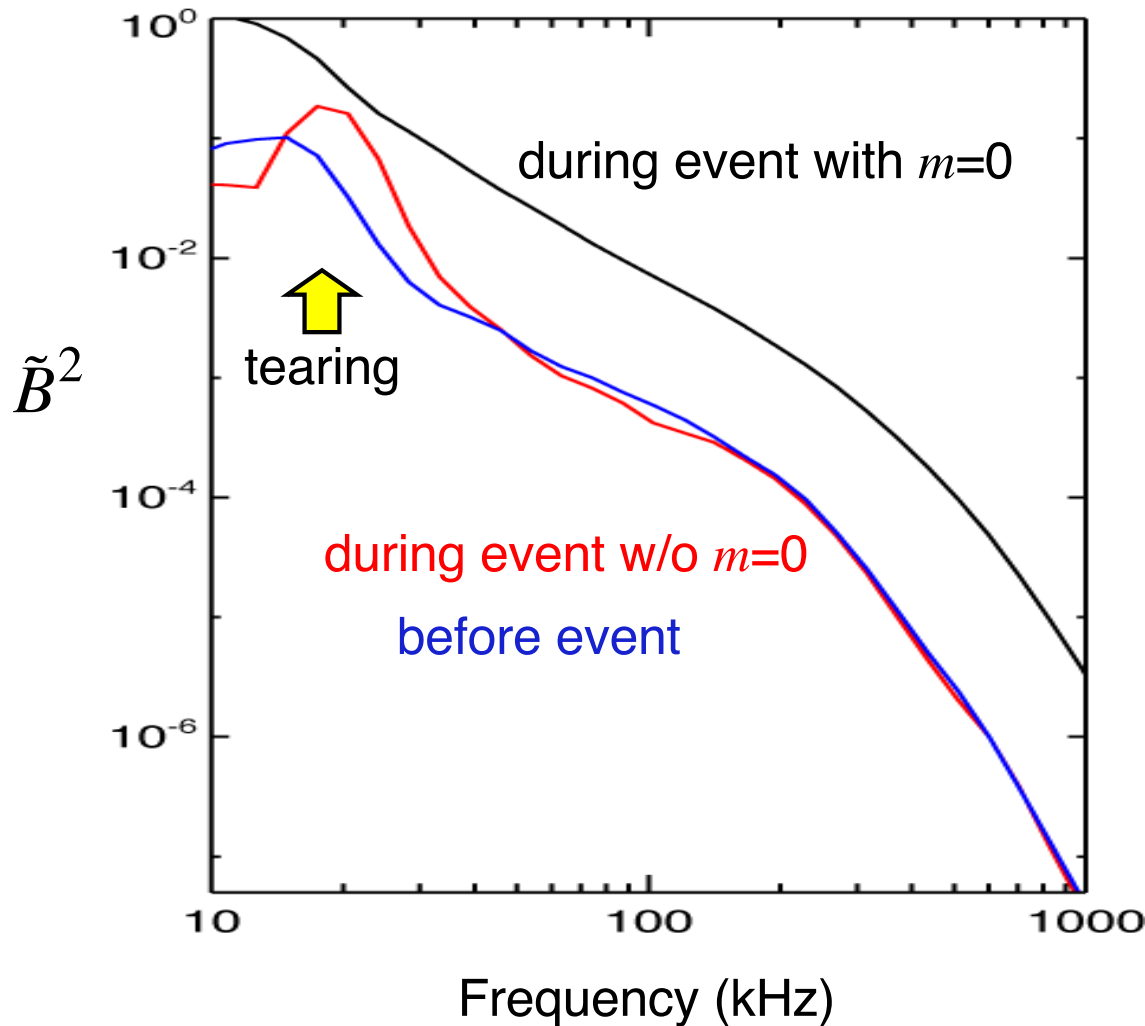
Relaxation process couples electron and ion momentum balance



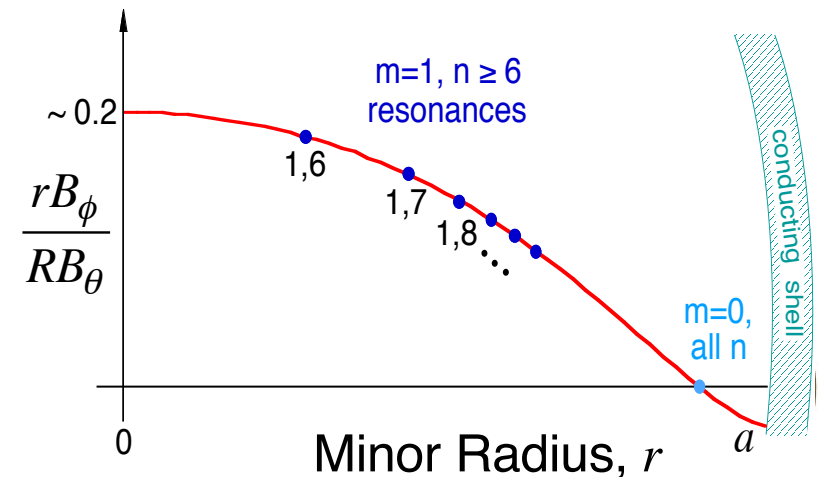
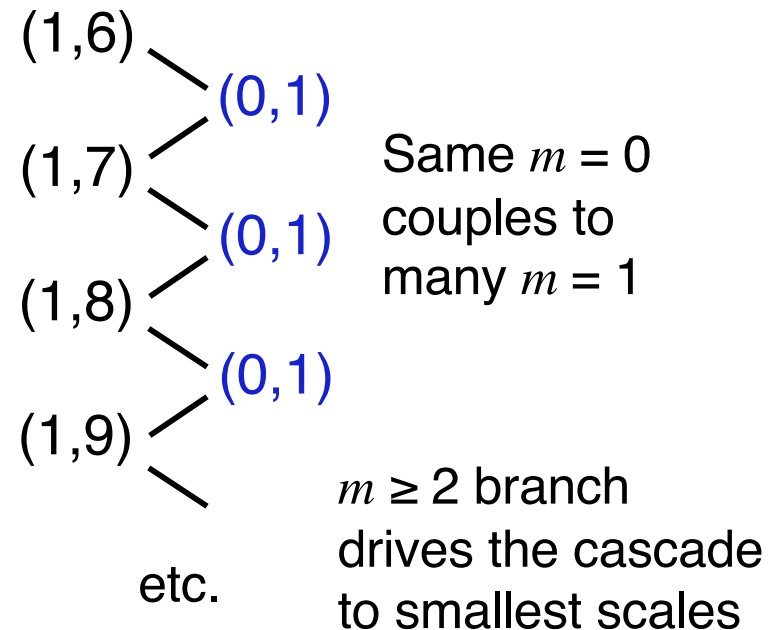
# Two-fluid relaxation signatures are measured, e.g., Hall emf



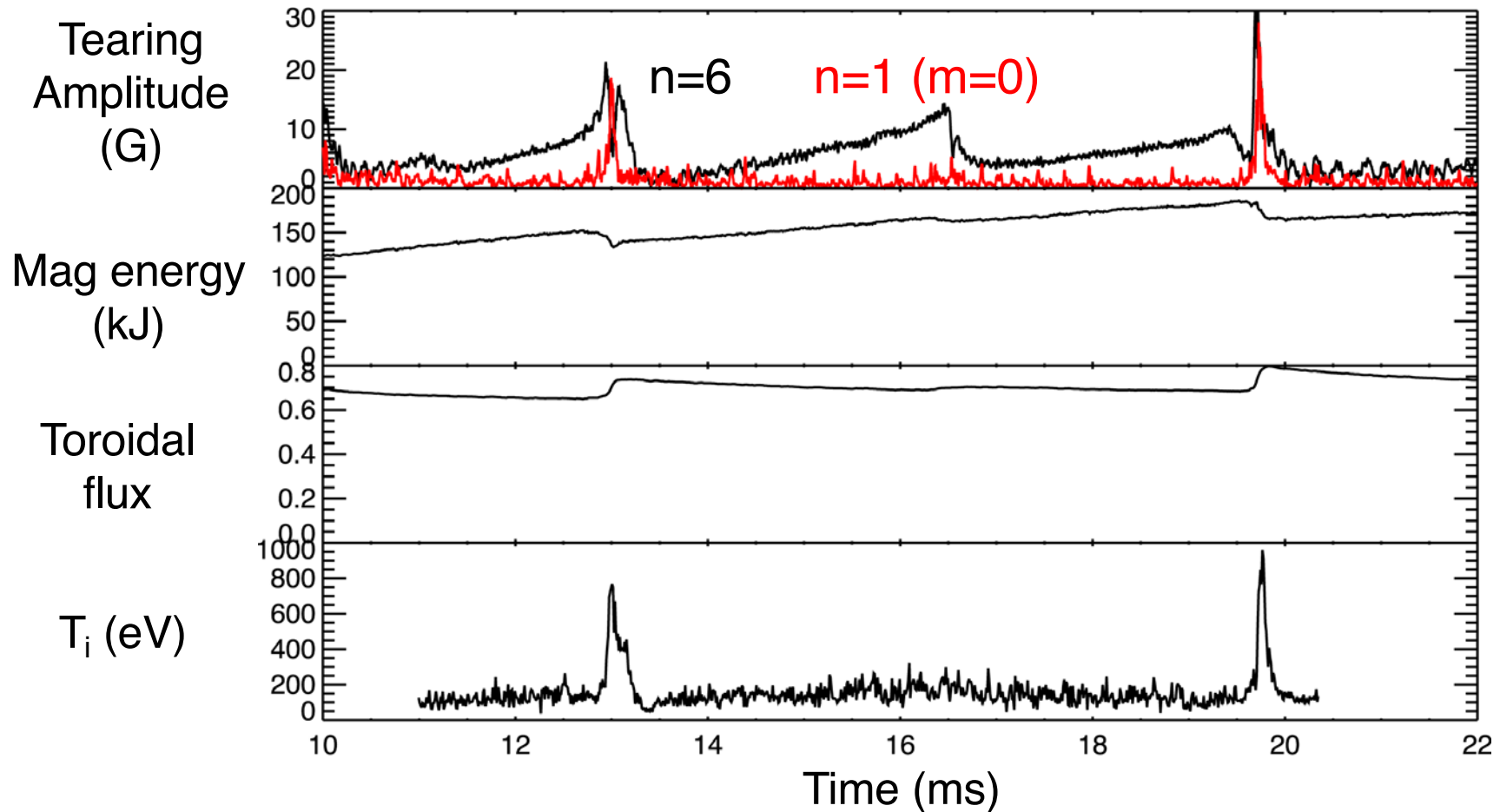
# The turbulent cascade is not as strong when nonlinear coupling at the driving scale is disabled



Schematic of 3-wave cascade:

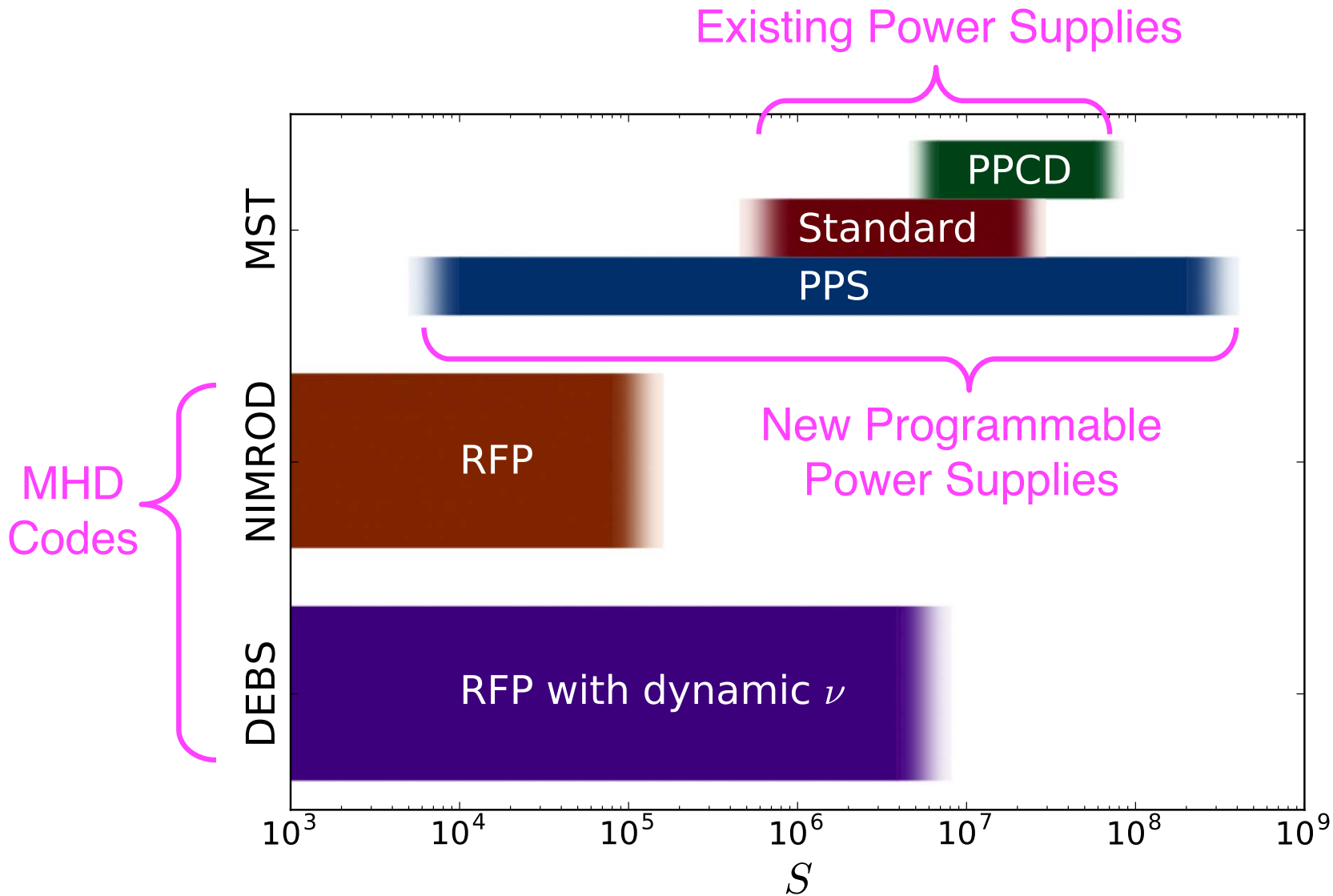


# Particle energization correlates with power flow into the cascade





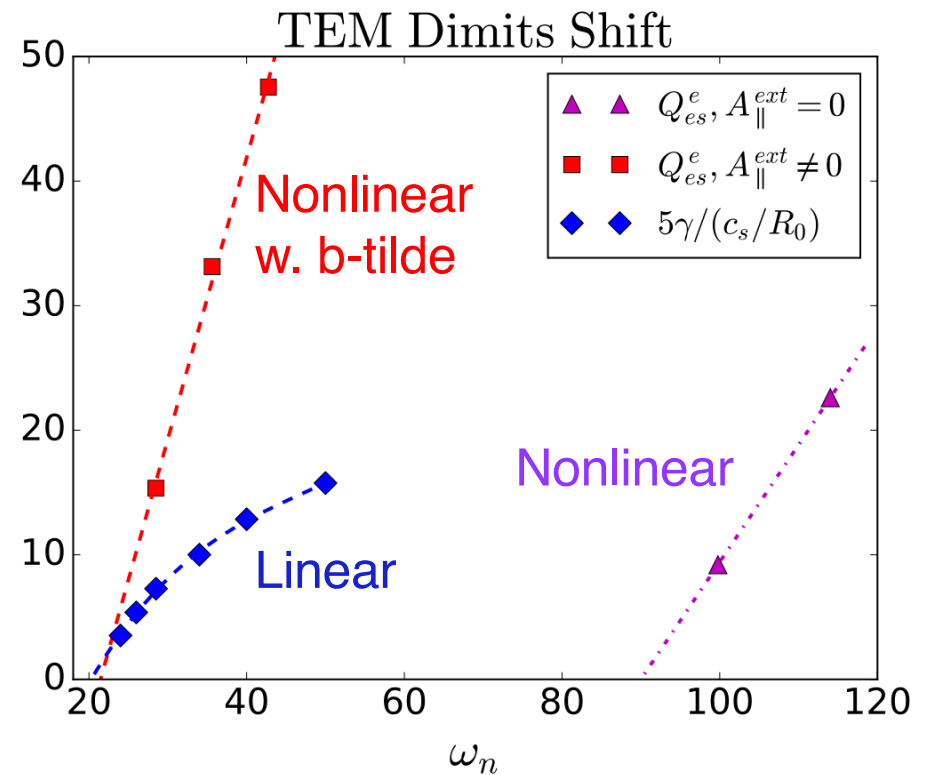
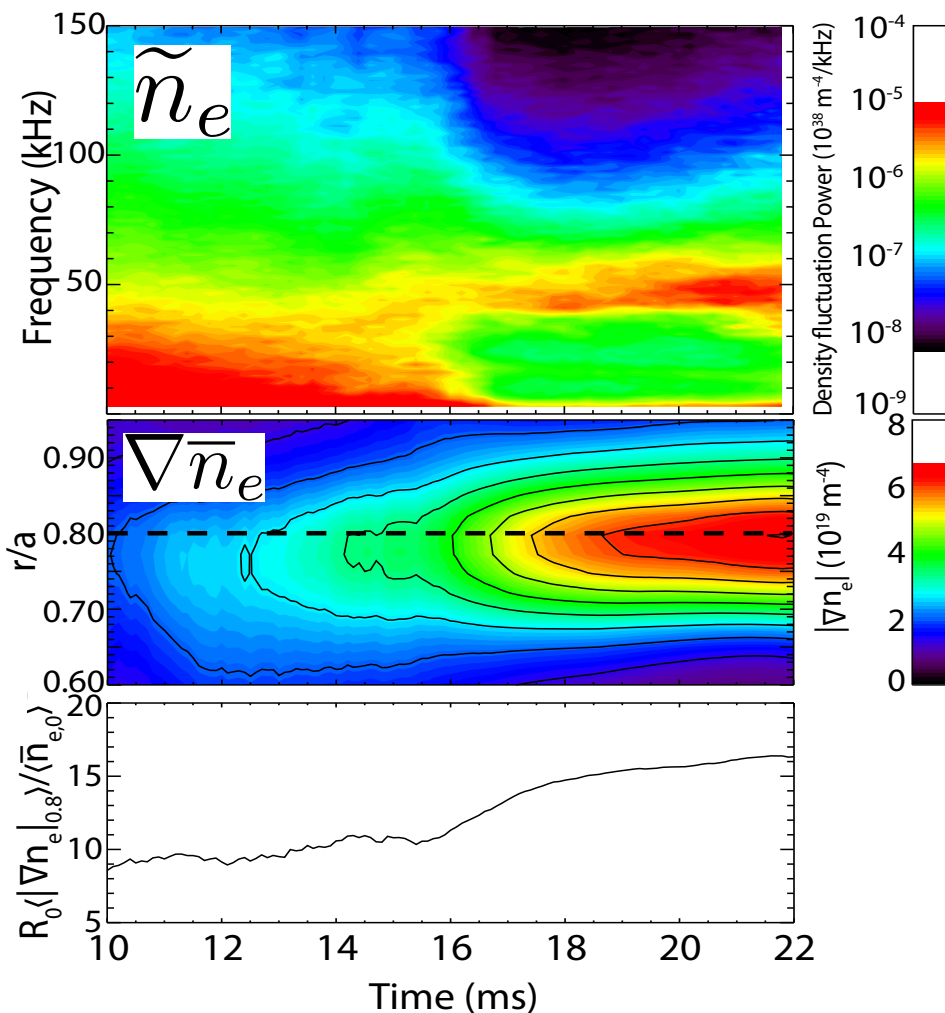
# New power supplies enable four orders of magnitude in Lundquist number, $S \approx 3 \times 10^{4-8}$



# Gyrokinetic modeling for the tearing cascade is a next challenge



- GENE modeling identifies standard drift mode branches in RFP equilibria
- Initially motivated by improved-confinement regime, but gradients in standard conditions are close to marginal stability (if not unstable)



- Ion heating and acceleration associated with magnetic reconnection from tearing instability is a powerful process in the RFP laboratory plasma
  - Gyro-resonant and stochastic processes are likely candidates to support the observed rapid heating and other features
  - Energetic tail formation for ions and electrons
- Global self-organization strongly coupled to turbulence and dissipation
  - Correlations in electric and magnetic field fluctuations are a hallmark of dynamo feedback
  - Inhomogeneity on the system scale, e.g., strong edge gradients
  - Global magnetic flux change drives produces ion runaway energization
  - Impact of transport processes, which can be quite different for ions and electrons